

Recent Progress on Hyper-Kamiokande Project



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Contents

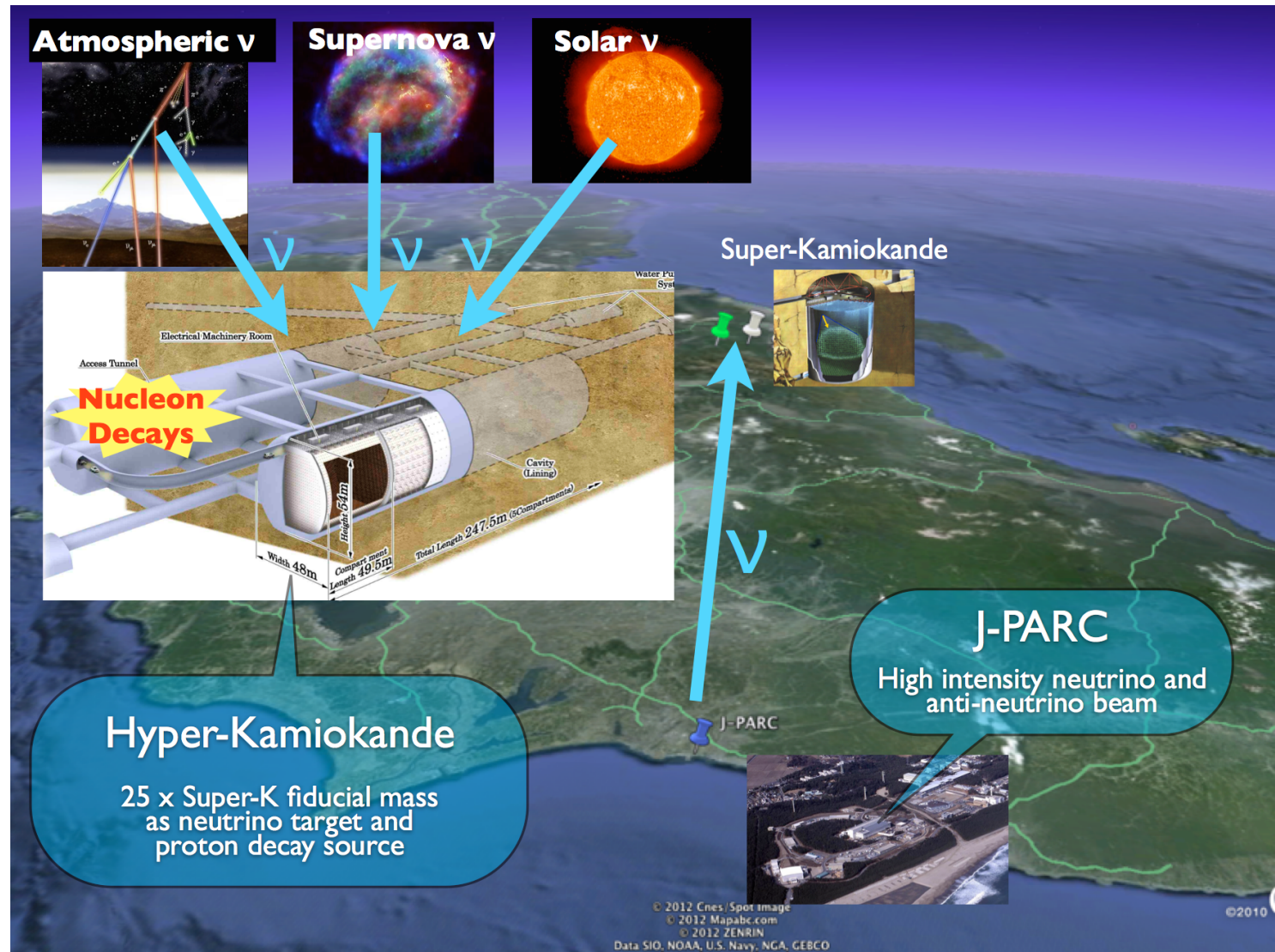
1. Overview
2. Physics
3. Recent progress
4. Summary

Contents

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2. Physics
3. Recent progress
4. Summary

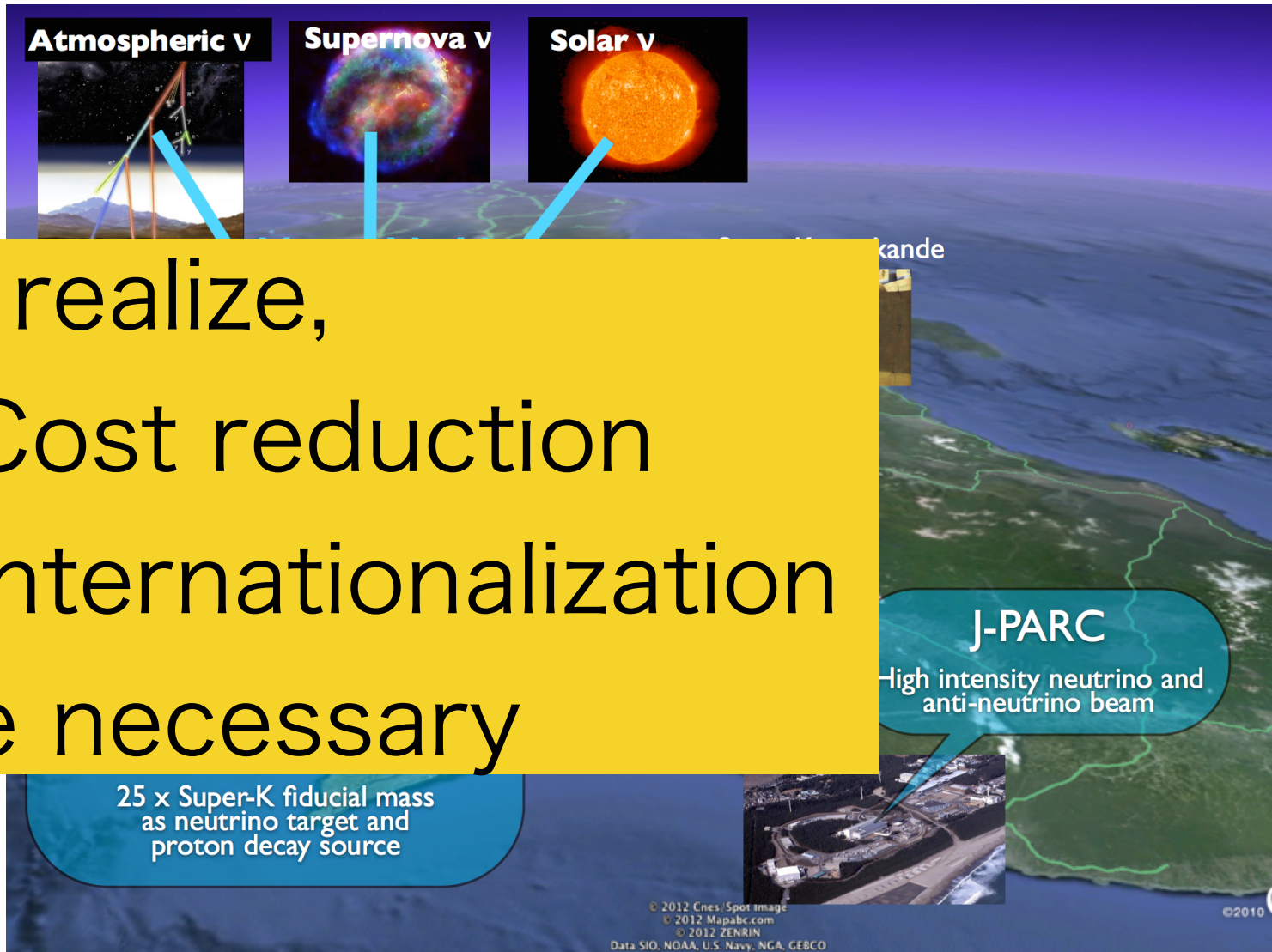
Hyper-Kamiokande

- Next generation water cherenkov detector
- Multi-purpose detector for various physics

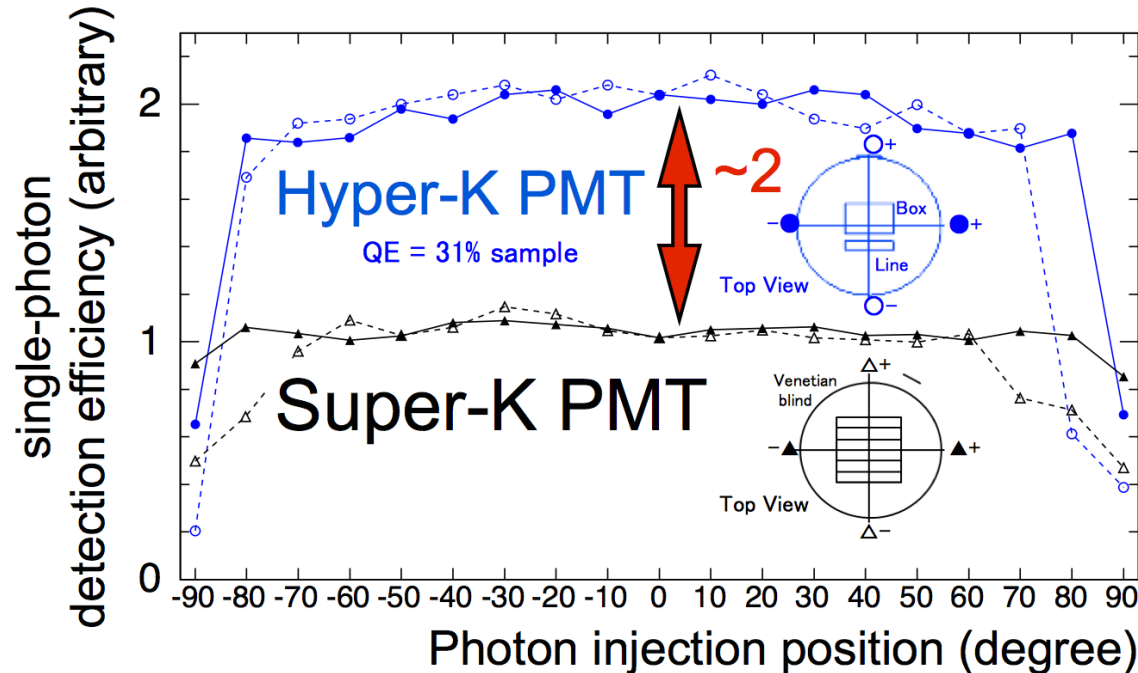


Hyper-Kamiokande

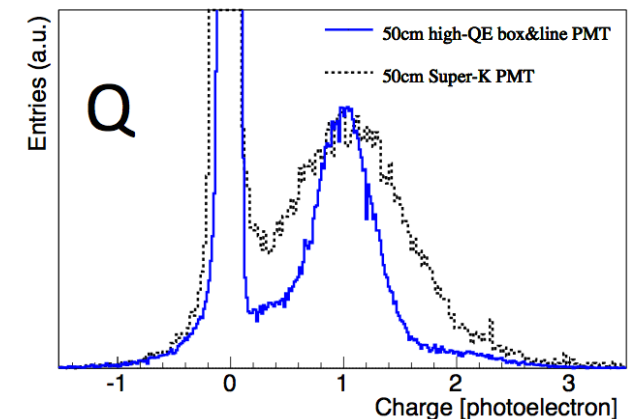
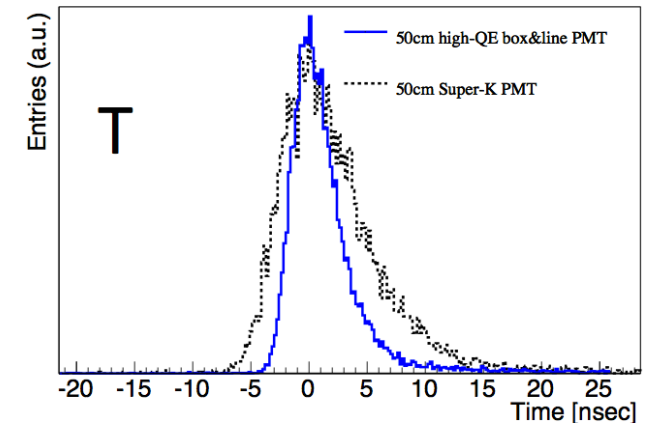
- Next generation water cherenkov detector
- Multi-purpose detector for various physics



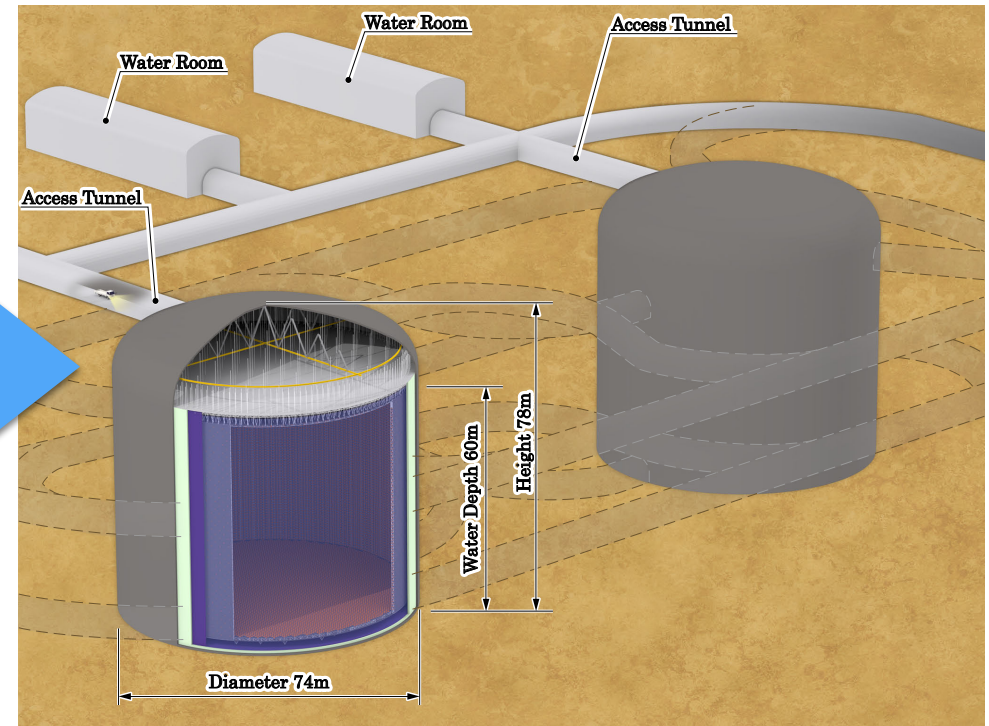
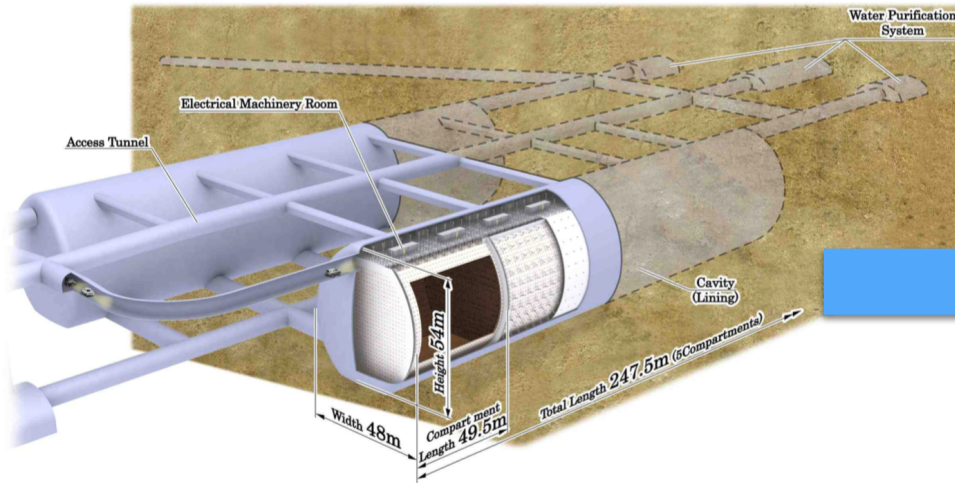
New Photo-Sensor



- New 20-inch photo-sensors: higher performance
 - Single-photon efficiency: $\times 2$
 - 1 p.e. timing resolution: $2\text{ ns} \rightarrow 1\text{ ns}$
 - 1 p.e. charge resolution: $53\% \rightarrow 35\%$
- Large impact on detector performance/physics sensitivity



Design Optimization



- SK-like cylindrical vertical tank: $\Phi 74\text{m} \times \text{H}60\text{m}$
 - Total volume: 260kton/tank, Fiducial volume: 190kton/tank
- Photo-coverage = 40% \rightarrow 40k ID PMTs/6.7k OD PMT
- 2 tanks with staging (1 tank at day1)
 - In this talk, assume 2nd tank operation starts from 7th year after 1st tank operation.

Contents

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2. Physics
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Physics at Hyper-K

- Rich physics topics!

- Neutrino oscillation (acc. ν , atom. ν , solar ν)

- CP violation

- Mass hierarchy

- $\theta_{23} = 45^\circ?$, $< 45^\circ?$, $> 45^\circ?$

- Day/night asym. in solar ν

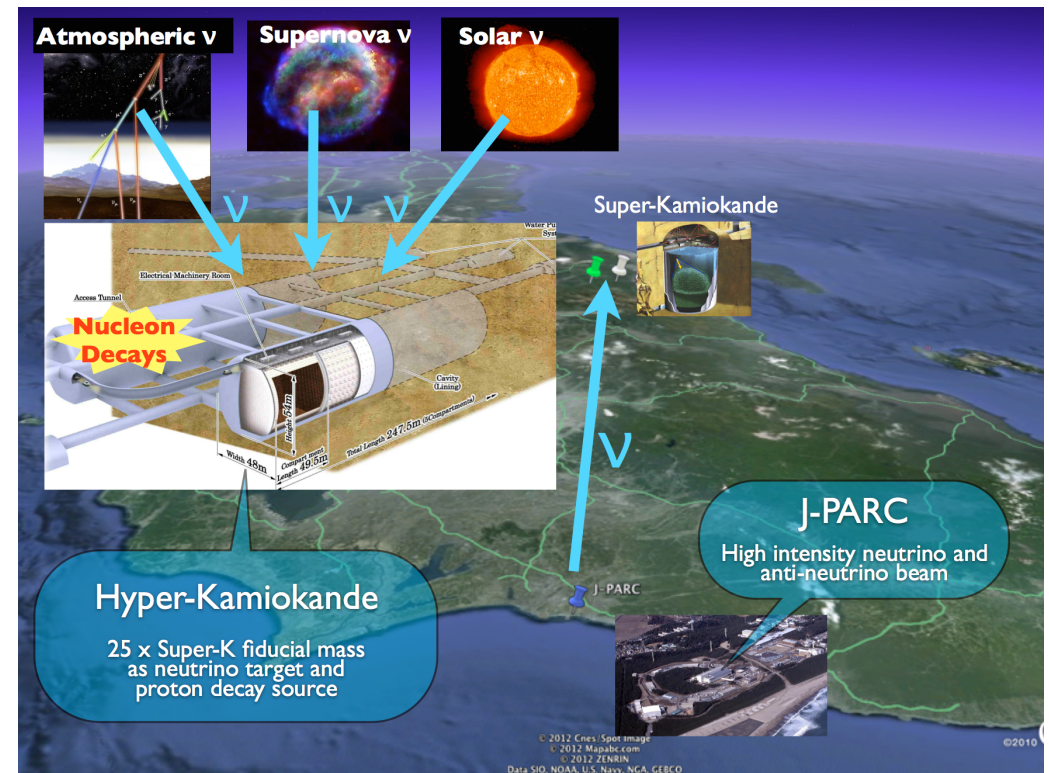
- Proton decay \rightarrow test of GUT

- $p \rightarrow e^+ \pi^0$ (SK 90% limit = 1.7×10^{34} y)

- Supernova ν

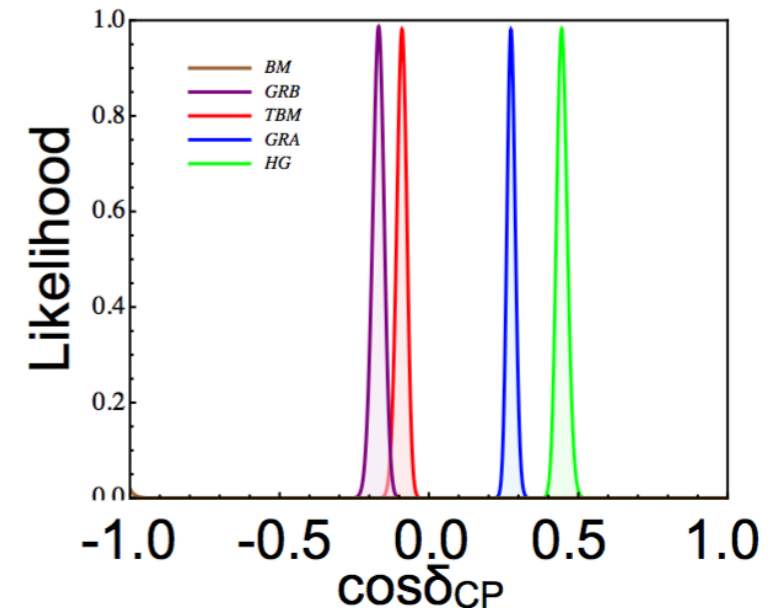
- Supernova burst ν

- Supernova relic ν



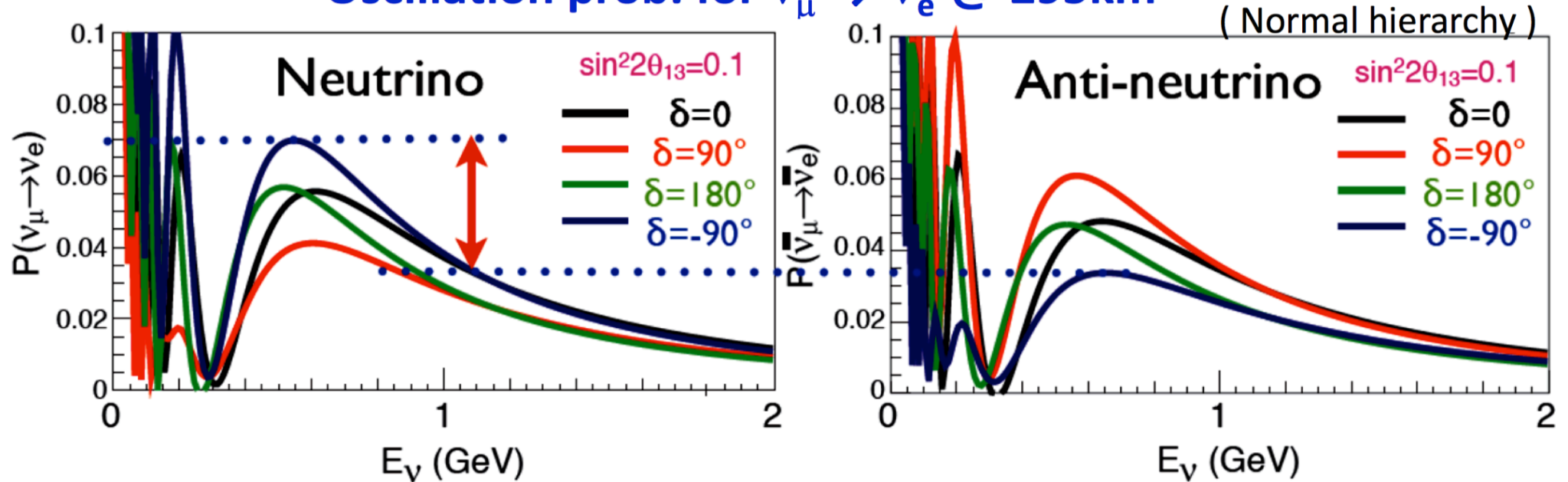
Why ν CPV is important?

- Leptonic (ν) CPV search is very important
 - The only known CPV source = CKM phase
 - Need other CPV source to explain the matter-antimatter asymmetry in the universe.
- Leptogenesis scenario only with Dirac CP phase
 - S. Pascoli et al., PRD 75, 083511 (2007) PDG review 2014
 - $|\sin \delta_{CP}| > \sim 0.6$
- Flavor symmetry prediction on δ_{CP}
 - e.g. Petrov 1504.02402v1
- Precise measurement is also important!



Measurement of CP Asymmetry with ν Beam

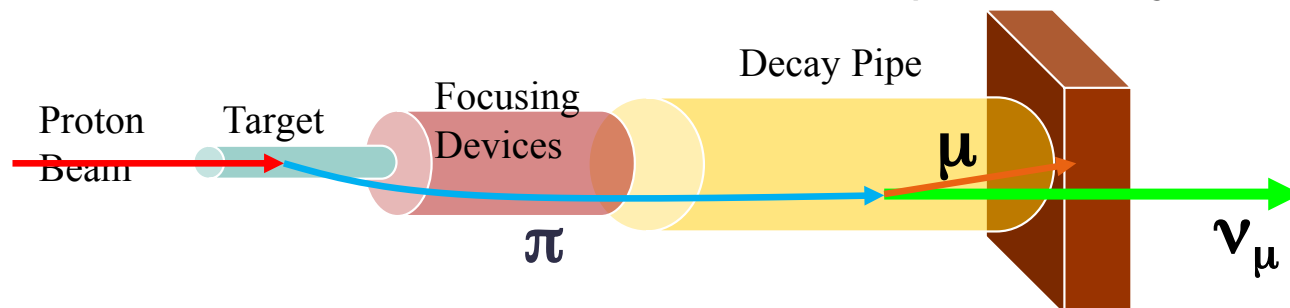
Oscillation prob. for $\nu_\mu \rightarrow \nu_e$ @ 295km



- Comparison of $P(\nu_\mu \rightarrow \nu_e)$ and $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$
 - Max. $\sim \pm 25\%$ difference from $\delta = 0$ case
 - Sensitive to exotic (non-MNS) CPV source

J-PARC Neutrino Beam

Conventional neutrino beam from pion decay



- High intensity beam

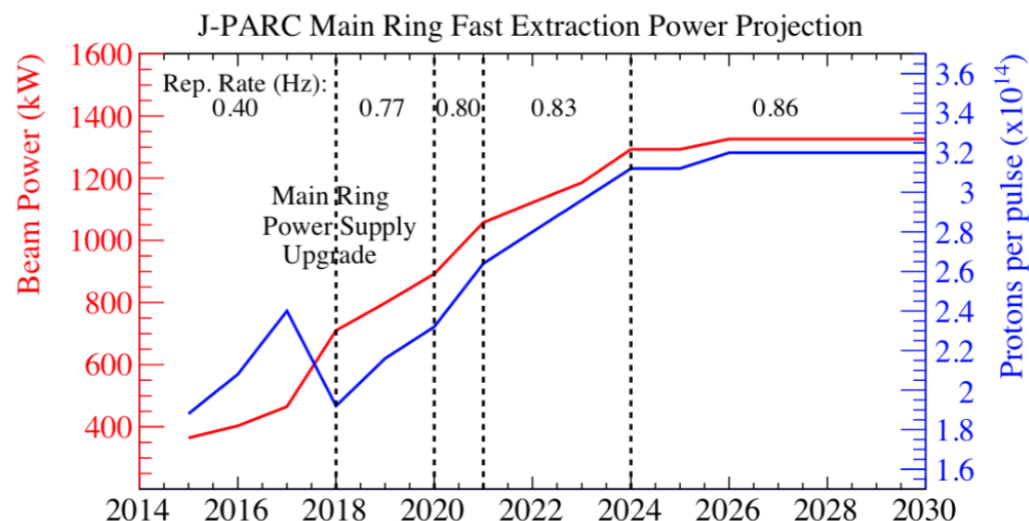
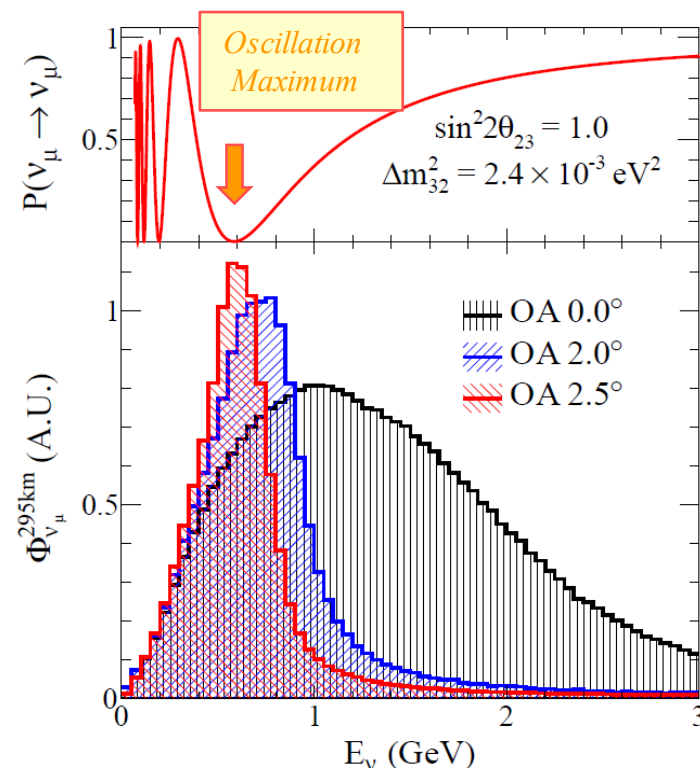
- 30 GeV, 750 kW proton beam (2×10^{14} ppp, 1.3s cycle)

- Off-axis beam (2~2.5°)

- 99% ν_μ purity
- Low energy narrow-band beam ~ 0.6 GeV
 - peak at 1st osc. max. with $L=300\text{km}$.

- Future beam power upgrade

- Aim to achieve **1.3 MW** by 2026
 - 3.2×10^{14} ppp, 1.16s cycle
- Great impact on HK LBL measurements

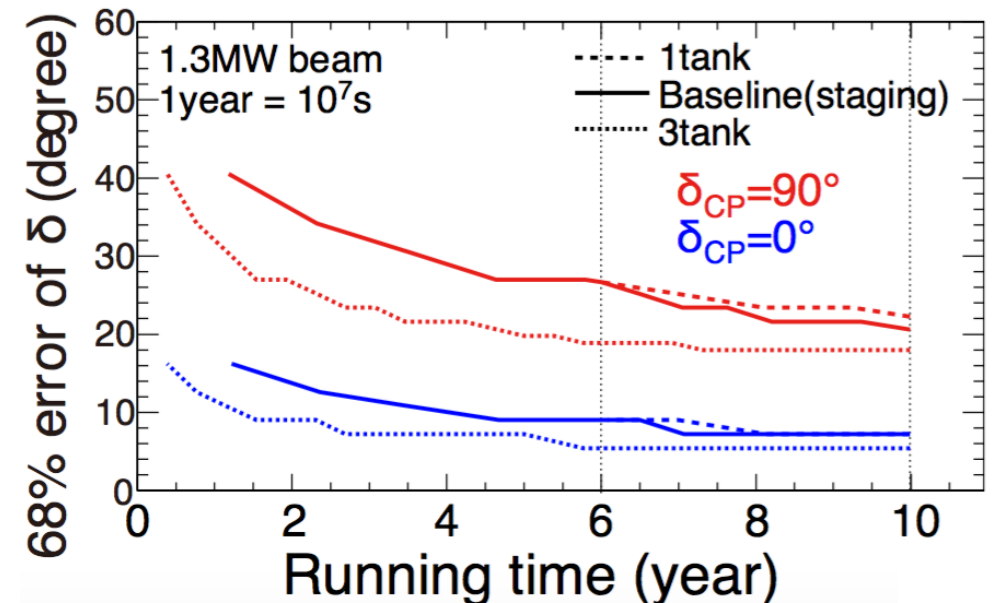
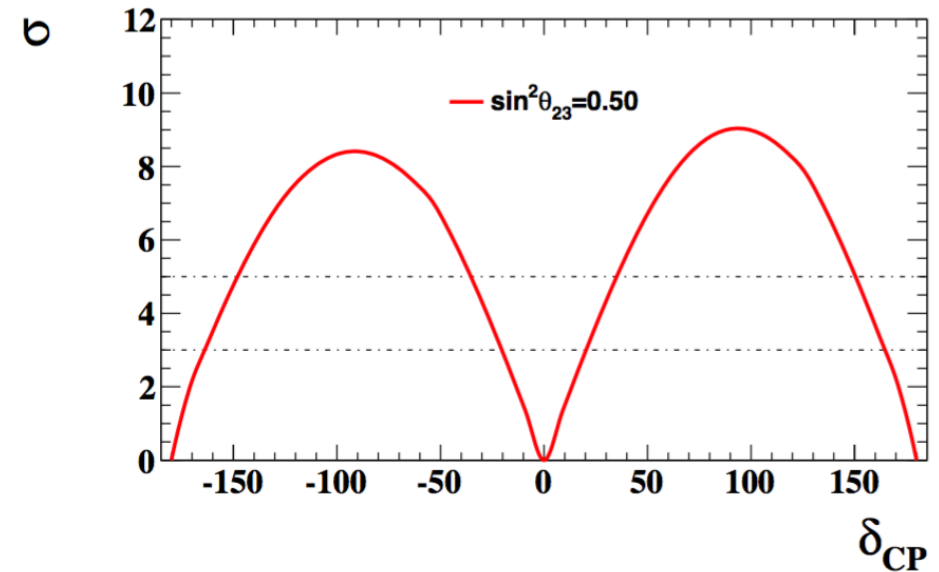


CPV Sensitivity

- Exclusion of $\sin \delta_{CP}=0$
 - $>8\sigma$ (6σ) for $\delta = -90^\circ$ (-45°)
 - $\sim 80\%$ coverage of δ parameter space with $>3\sigma$
- δ_{CP} measurement precision
 - $7\sim 21^\circ$ precision

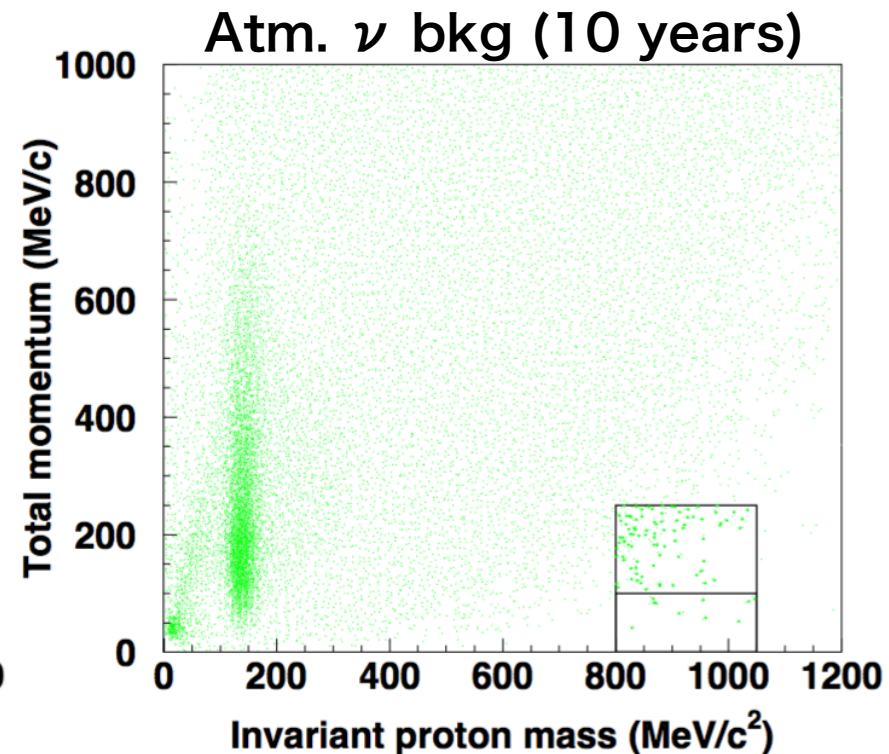
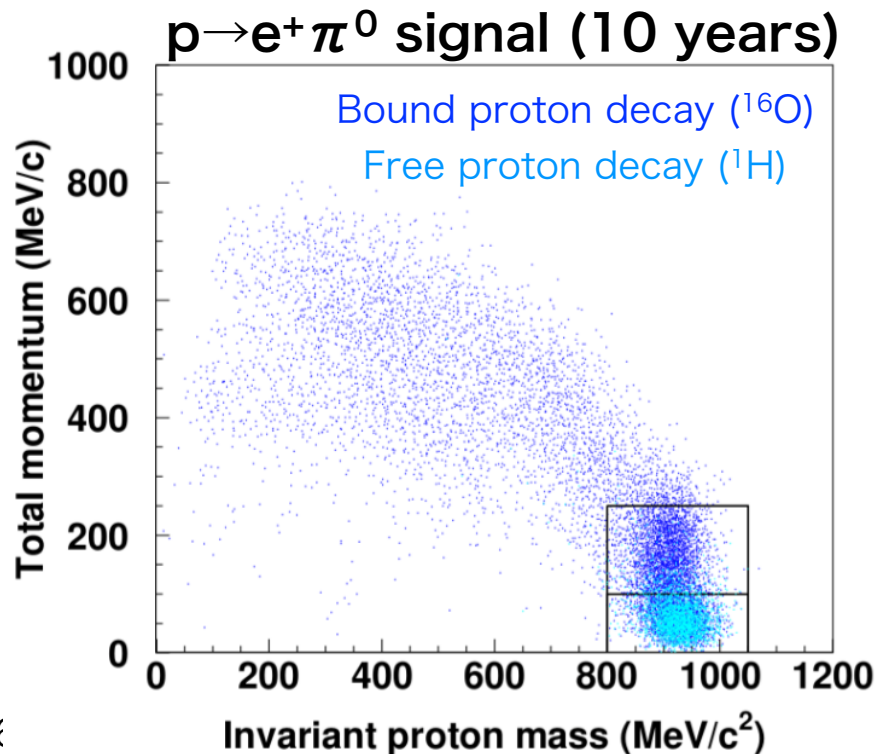
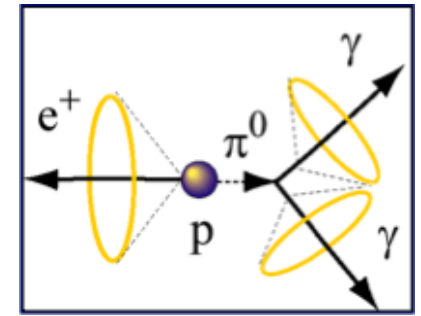
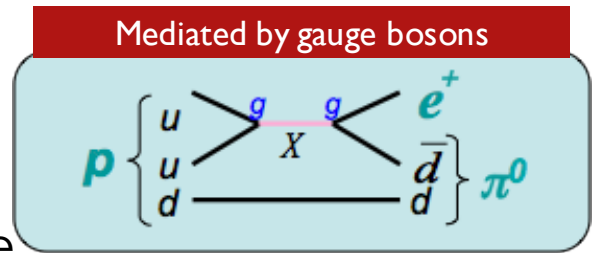
		sin $\delta=0$ exclusion		68% error	
		$>3\sigma$	$>5\sigma$	$\delta=0^\circ$	$\delta=90^\circ$
Old	7.5MWy	76%	58%	7.5°	19°
2tank (staging)	13MWy	78%	62%	7.2°	21°

Normal mass hierarchy

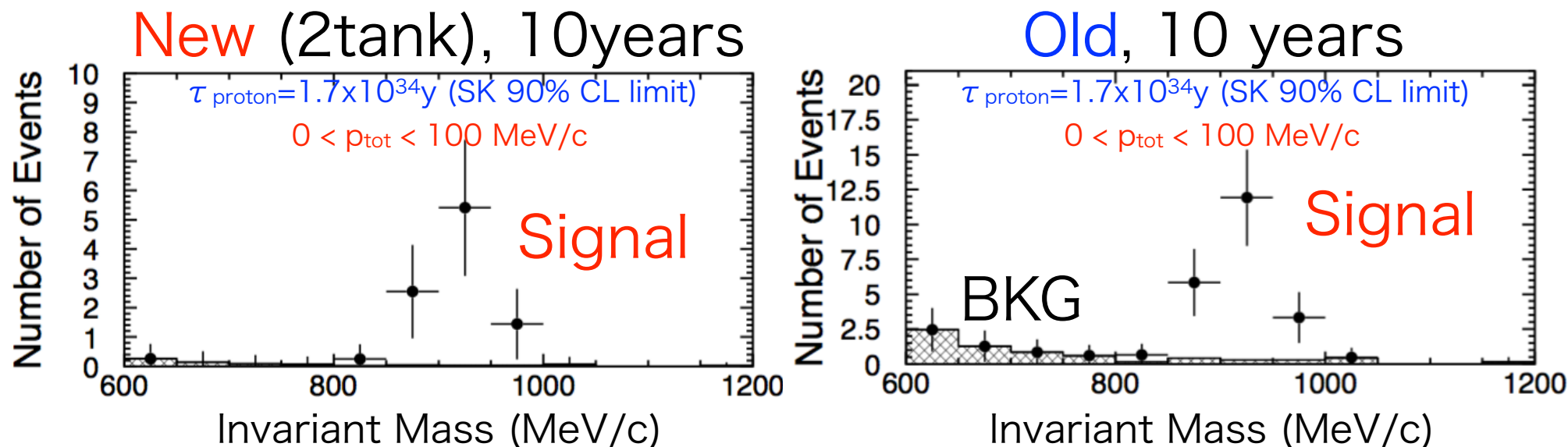


Search for Proton Decay

- Proton decay = **direct observation of GUT**
- $p \rightarrow e^+ \pi^0$: leading decay mode in many models
 - Water cherenkov detector has advantages for this mode
 - All decay products are visible.
 - Good PID and efficiency @ 1 GeV
 - Free proton in hydrogen atom \leftrightarrow bound proton in ^{16}O



S/N improvement for Proton Decay



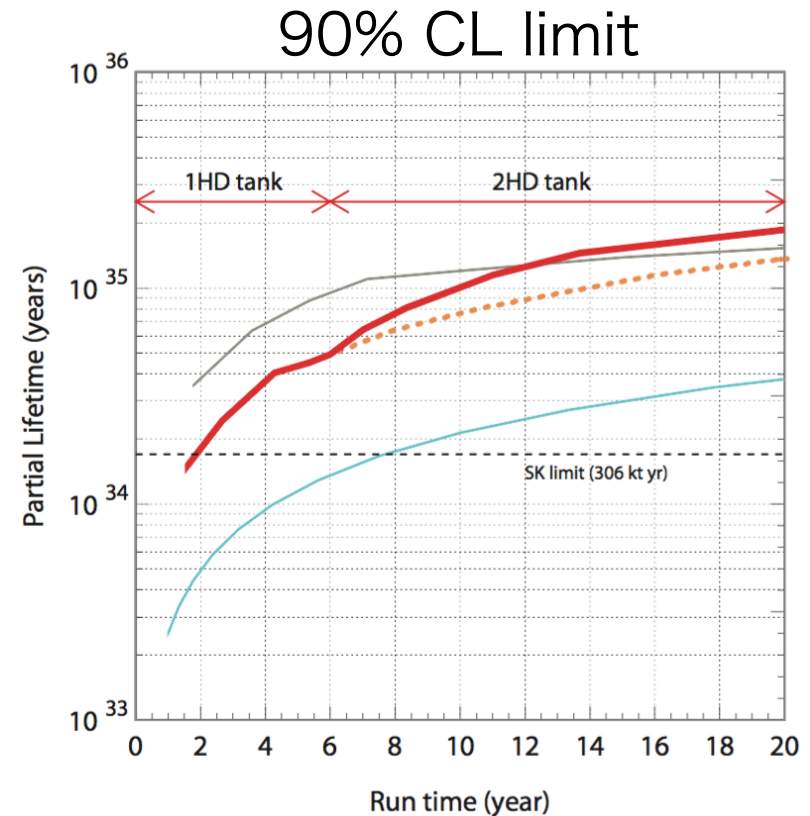
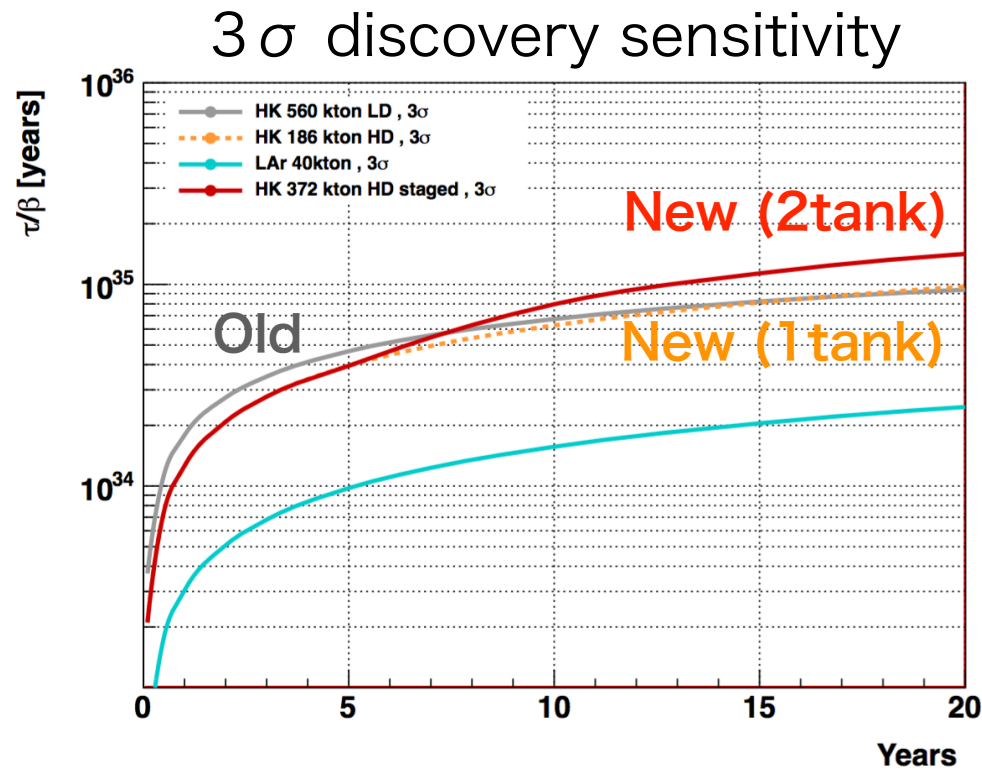
- BKG reduced to $\sim 1/5$ thanks to higher photo-coverage (40%)
- No reduction in signal eff.

	$p_{\text{tot}} < 100 \text{ MeV/c}$		$100 < p_{\text{tot}} < 250 \text{ MeV/c}$	
	Sig. $\epsilon(\%)$	Bkg (/Mtyr)	Sig. $\epsilon(\%)$	Bkg (/Mtyr)
Old	18.8	0.27	20.4	2.17
New	18.7	0.06	19.4	0.62

$\sim 9 \sigma$ discovery potential !

(in case of $\tau_{\text{proton}} = 1.7 \times 10^{34}$ years: SK 90% CL limit)

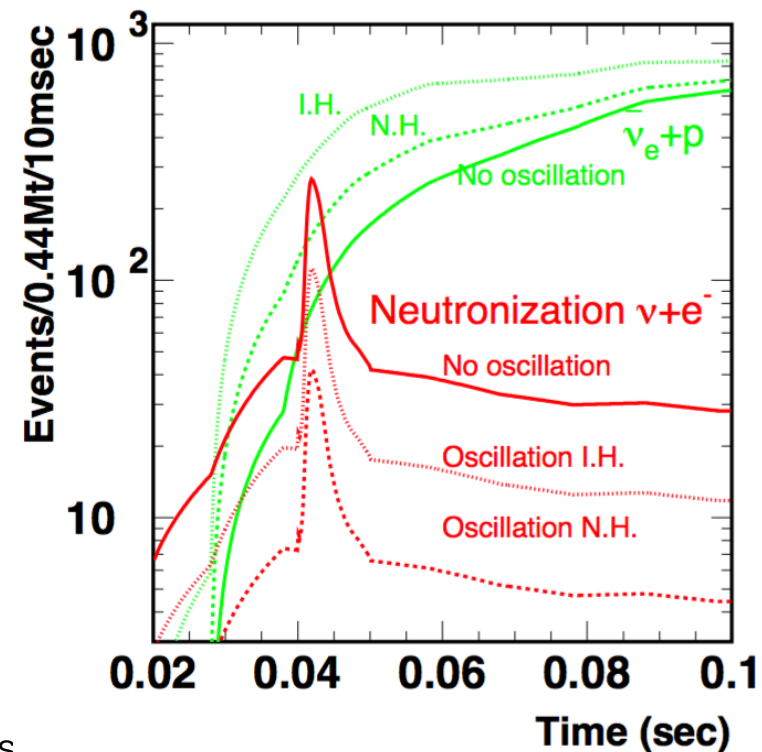
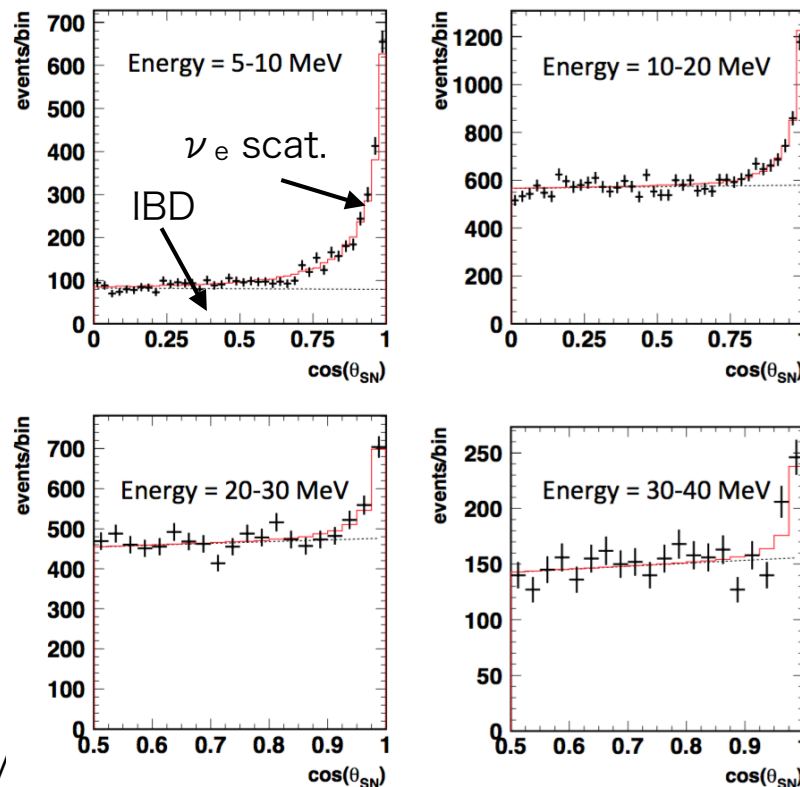
Proton Decay Sensitivity



- 3 σ discovery can be achieved within 15 years if lifetime = 1×10^{35} years
- Higher photo-coverage (40%) gives better sensitivity even with smaller fiducial volume.
 - 3 σ discovery sensitivity: New(1tank) ~ old
 - 90% CL limit: New(2tank) ~ old

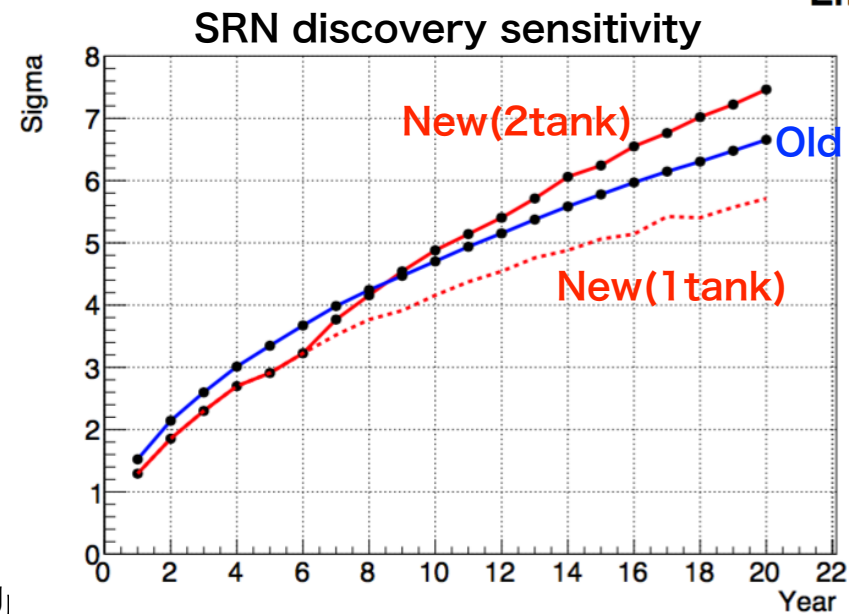
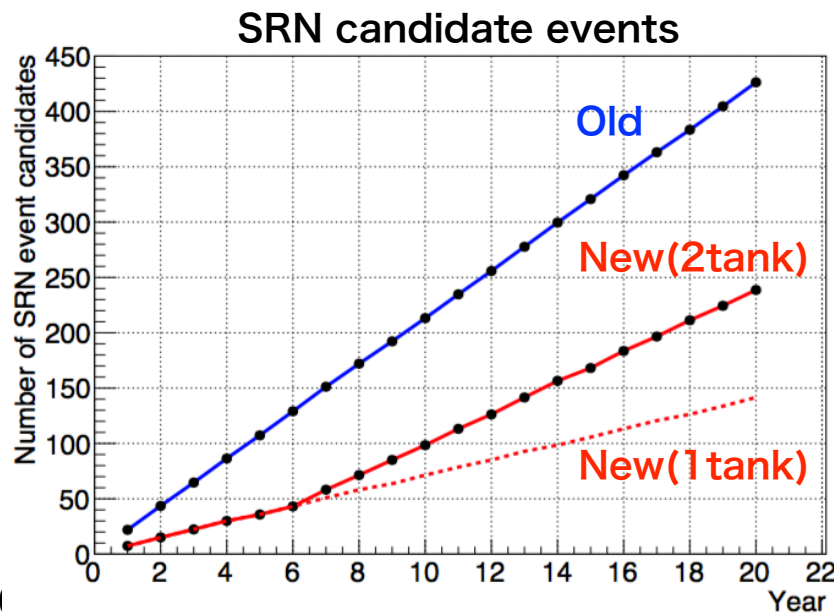
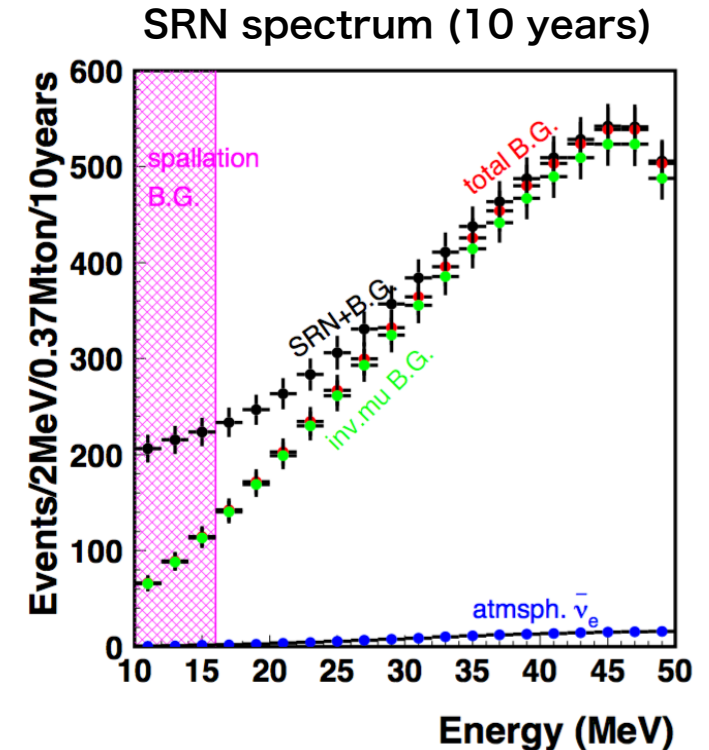
Supernova Burst Neutrinos

- 100k~160k ν events from SN at 10kpc \rightarrow very rich info.
- Inverse beta decay ($\bar{\nu}_e + p \rightarrow n + e^+$): 98k~136k evt. \rightarrow isotropic
- $\nu_e + e^-$ scattering: 4k~5k evt. \rightarrow directional information
- ν_e from neutronization: 12~80 evt. \rightarrow SN explosion mechanism
- Property of neutrino: absolute mass, mass hierarchy



Search for Supernova Relic Neutrinos

- $O(10^{17})$ SNs occurred in the past universe
- Expected flux: $0.3 \sim 1.5$ evt/cm²/s (>17.5 MeV)
- Background: spallation, atm. ν
- Higher photo-coverage helps to increase neutron-tagging efficiency
- 4.8σ discovery sensitivity expected after 10 years



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4. Summary

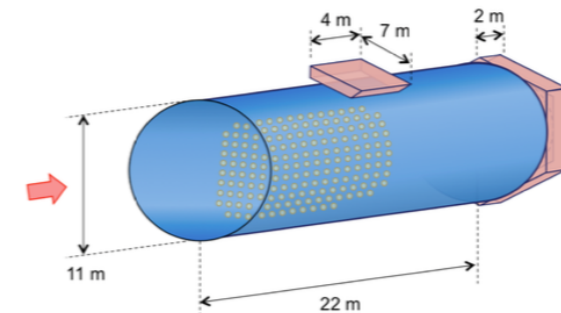
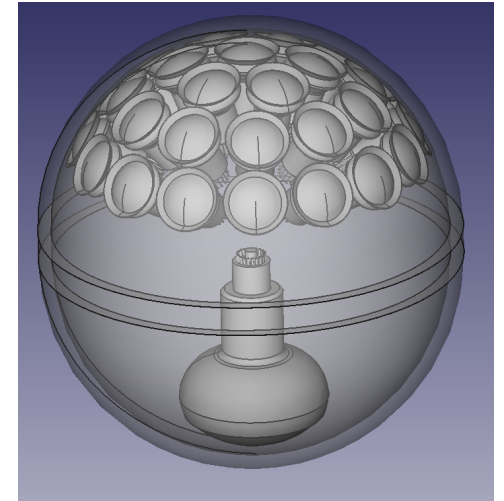
HK Proto-Collaboration

- HK proto-collaboration formed (2015. Jan.)
 - ~250 members from 13 countries
- MoU for cooperation in HK project between ICRR/U-Tokyo and IPNS/KEK (2015. Jan.)



International Contribution

- Japanese contribution
 - Cavern, tank, half of photo-sensors
- Foreign contribution
 - Half of photo-sensors
 - Multi-PMT module (for ID/OD)
 - Texas 11" PMT by ETL (for OD)
 - Electronics, DAQ, and so on.
 - Near detectors
 - Upgrade in near detector @ 280m
 - Intermediate detector @ 1~2km



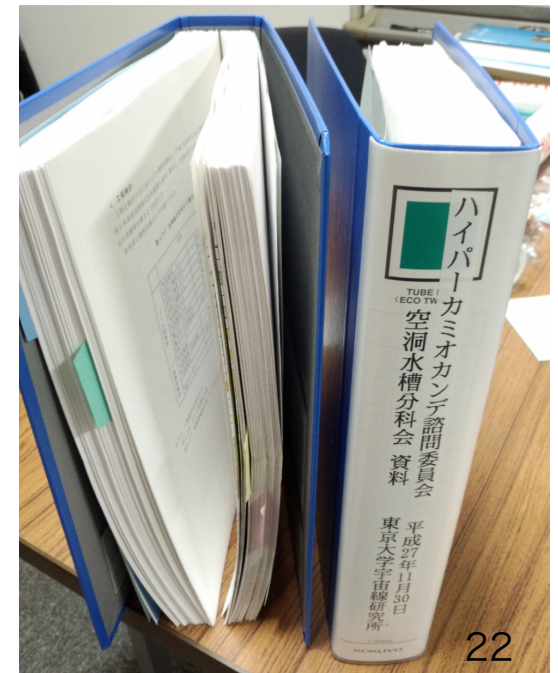
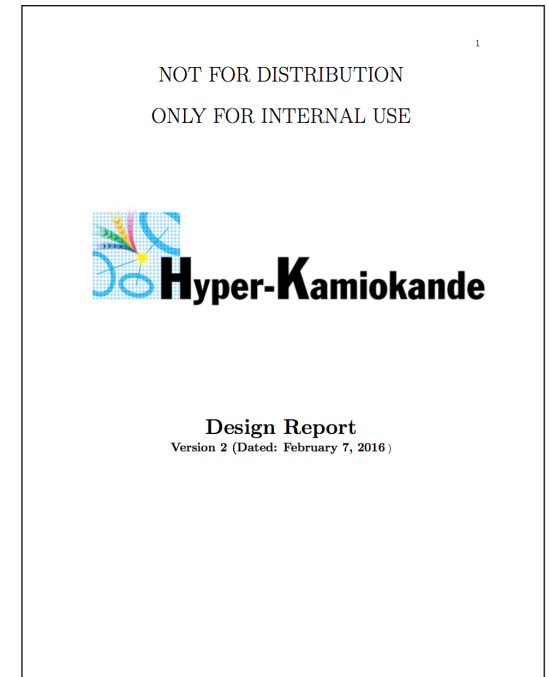
HK Design Review

- **HK Design Report**

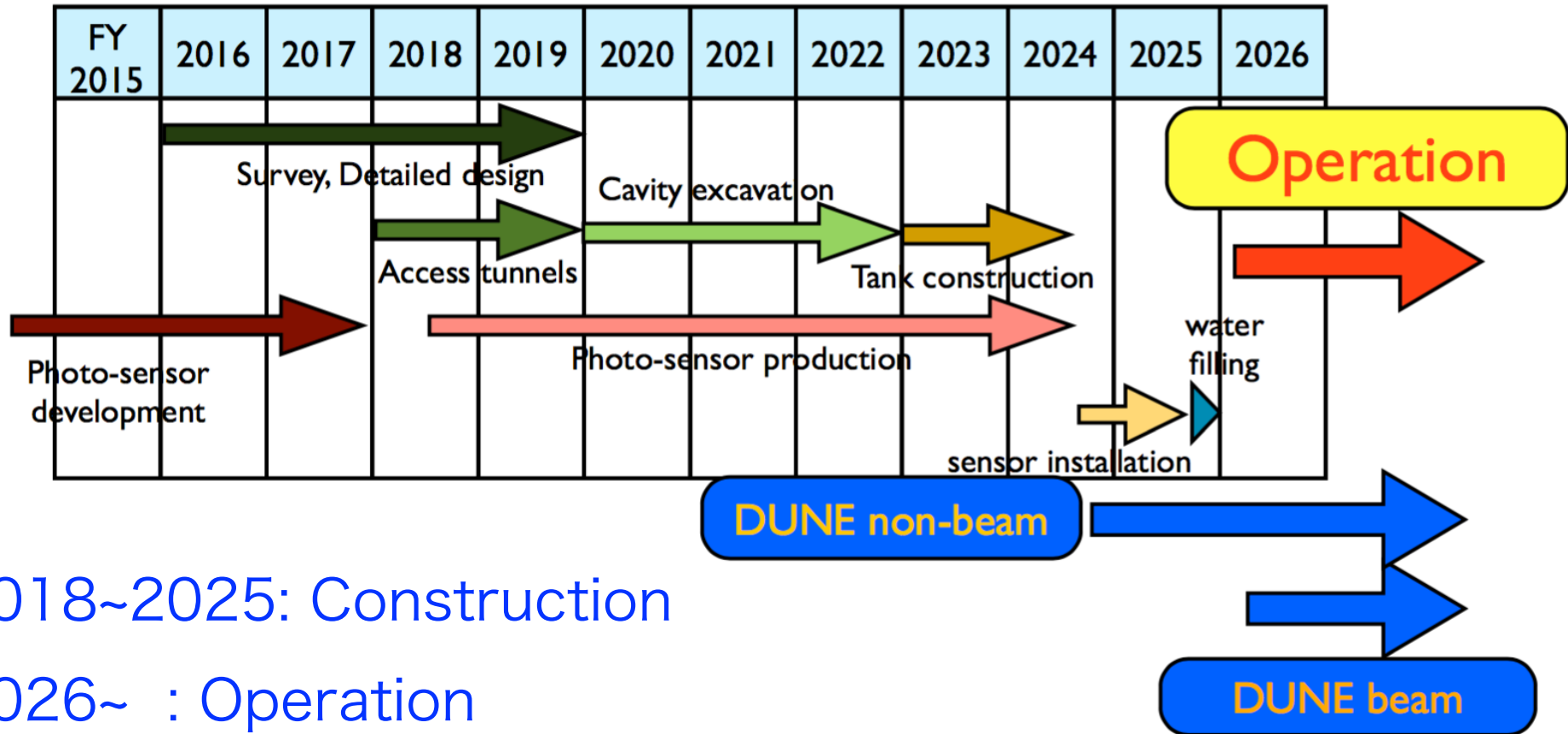
- Physics, detector, cost, organization, etc. (in English)
- Cavern and tank design (in Japanese)

- **Hyper-Kamiokande Advisory Committee**

- Formed under directors of IPNS/KEK and ICRR/U-Tokyo
- **Main committee** → review the HK Design Report
 - HKAC report will be submitted to the directors soon
- **Sub-committee** → review cavern and tank construction



Timeline (1st Tank)



- 2018~2025: Construction
- 2026~ : Operation
 - CPV study
 - Atm•Solar•Supernova ν study, Proton decay searches
- Timely budget allocation is very important for international competition!!

Summary

- Hyper-Kamiokande
 - Next generation water cherenkov neutrino & nucleon decay detector.
- Design optimization
 - Vertical cylindrical shape ($\Phi 74\text{m} \times \text{H}60\text{m}$) \rightarrow 26kton/tank
 - 2 tanks with staging (to start as early as possible)
- Many physics topics can be studied.
 - Discovery of CPV \rightarrow Precise measurement
 - Search for Proton decay
 - Detection of astrophysical neutrinos (Solar ν , SN, SRN)
- Recent progress
 - Porto-Collaboration, Design report, Review committee, submission to SCJ.
- We are aiming to realize HK project and to start operation from 2026.