



Kobayashi-Maskawa Institute  
for the Origin of Particles and the Universe



NAGOYA UNIVERSITY



# NEWSdm experiment

Directional Dark Matter Search with Super-high  
resolution Nuclear Emulsion

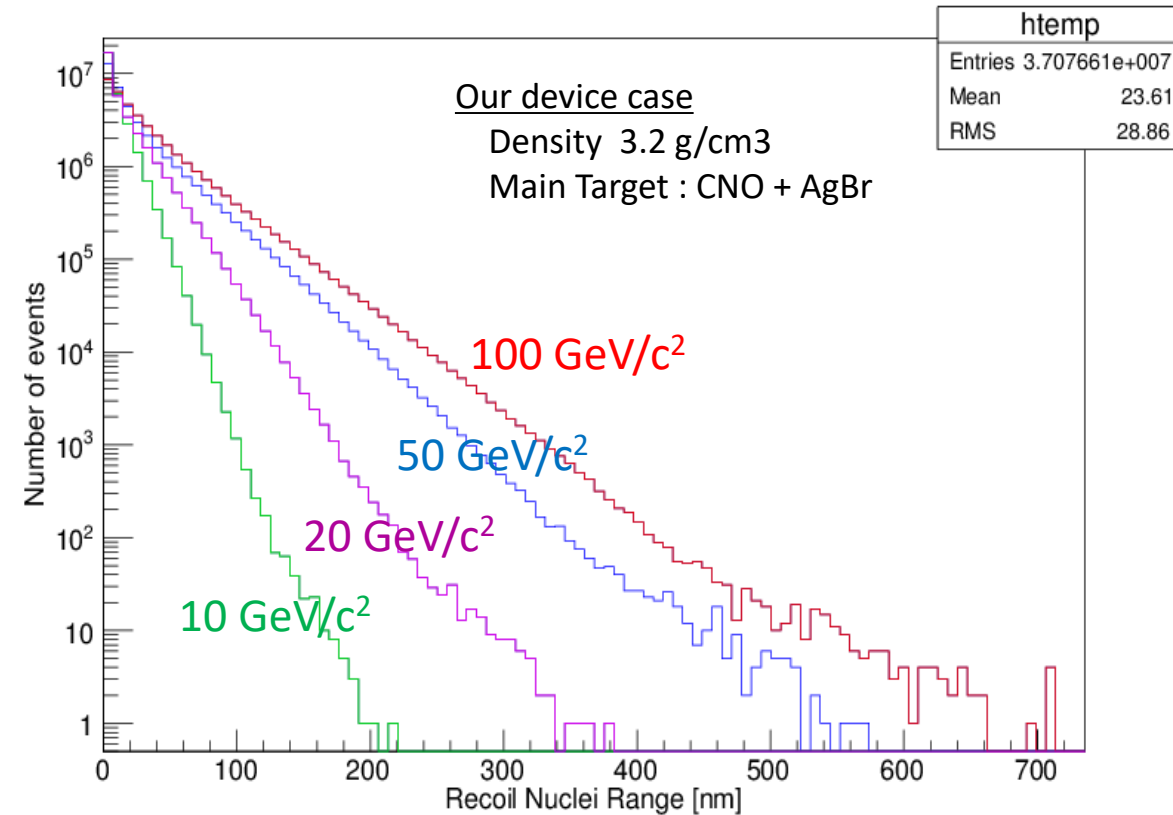
**Tatsuhiro NAKA**

KMI, Nagoya University

on behalf of NEWSdm collaboration

# Direction sensitive dark matter with solid detector

- ❑ Super-high resolution device using capability of detecting nano-scale tracks
- ❑ Readout technologies for such very short length tracks
- ❑ Understanding and rejection of backgrounds



# Direction sensitive dark matter with solid detector

## This talk

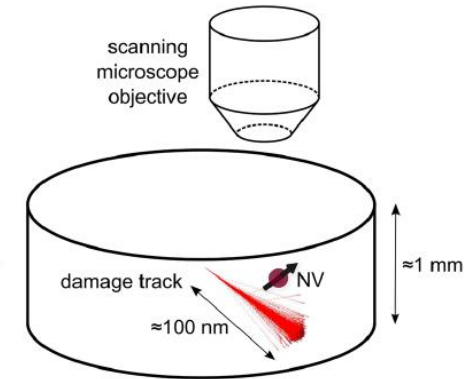
### Super-fine grained Nuclear emulsion (Nano Imaging Tracker : NIT)

First detector demonstrated capability of tracking  
to low-velocity nuclear recoil

## New Idea and on studying

### ➤ Diamond

Microscope imaging of luminescence  
due to N-V center in diamond  
*Phys. Rev. D. 96 035009 (2017)*



### ➤ Anisotropic crystal (e.g., ZnWO<sub>4</sub>)

### ➤ Carbon nano tube

Carbon nanotube target + gaseous TPC  
*arXiv:1412.8213 [physics.ins-det]*

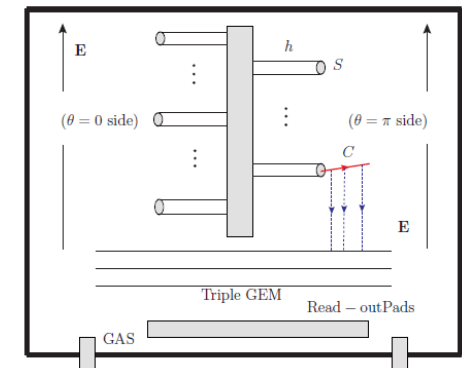


### ➤ Rock (but not directional search)

Crystal defect tracking in Ancient mineral

→ already M or G year exposure

*arXiv:1811.06844v1 [astro-ph.CO] 16 Nov 2018*



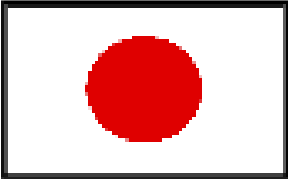



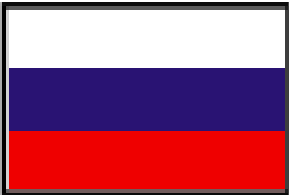

# NEWSdm experiment [Nuclear Emulsion for WIMPs Search – directional measurement]



<http://news-dm.lngs.infn.it>

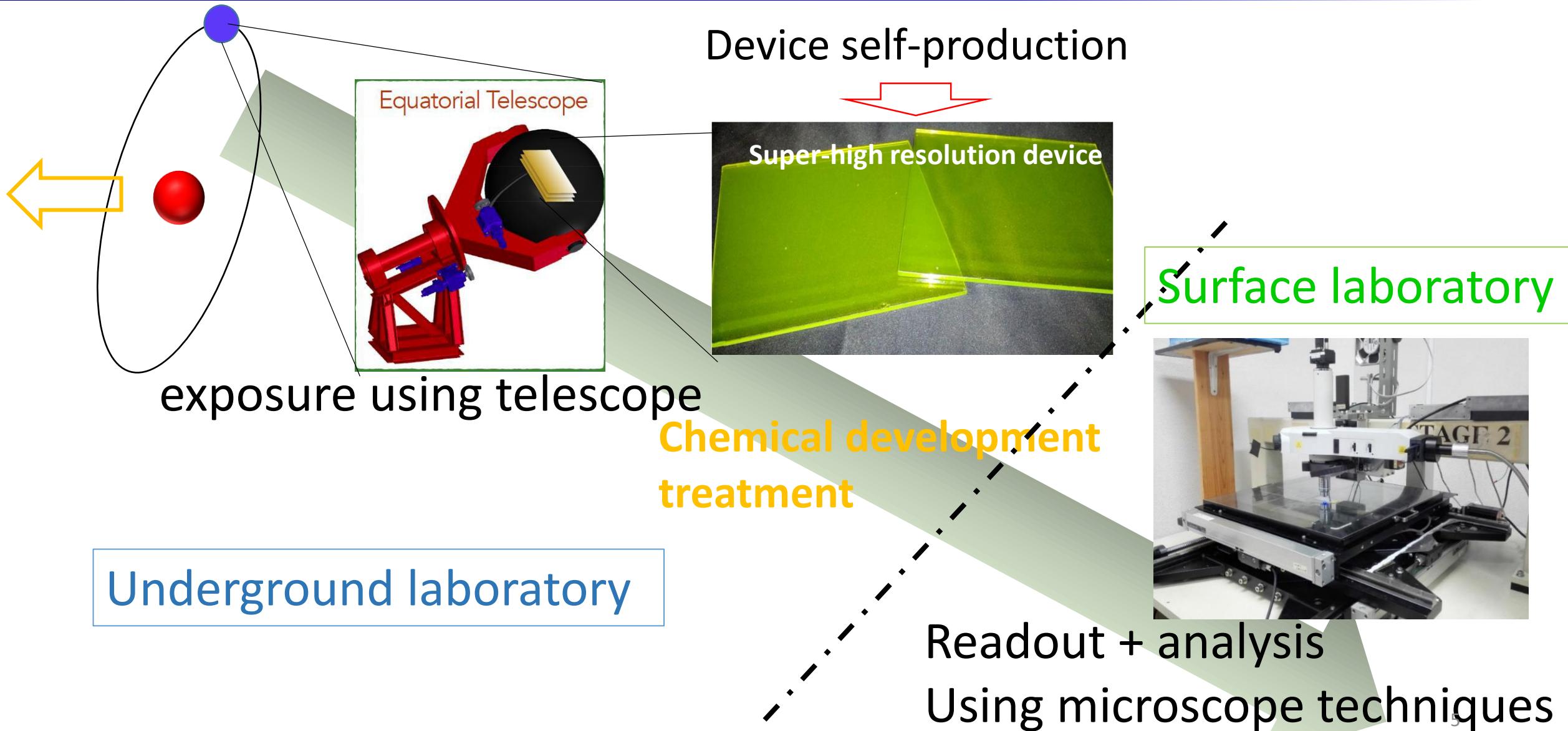
LOI under review by the LNGS science committee

NEWS: Nuclear Emulsions for WIMP Search  
Letter of Intent  
(NEWS Collaboration)

		
Chiba Nagoya		METU Ankara
		
Bari GSSI LNGS Napoli Roma	LPI RAS Moscow JINR Dubna SINP MSU Moscow INR Moscow Yandex School of Data Analysis	Gyeongsang

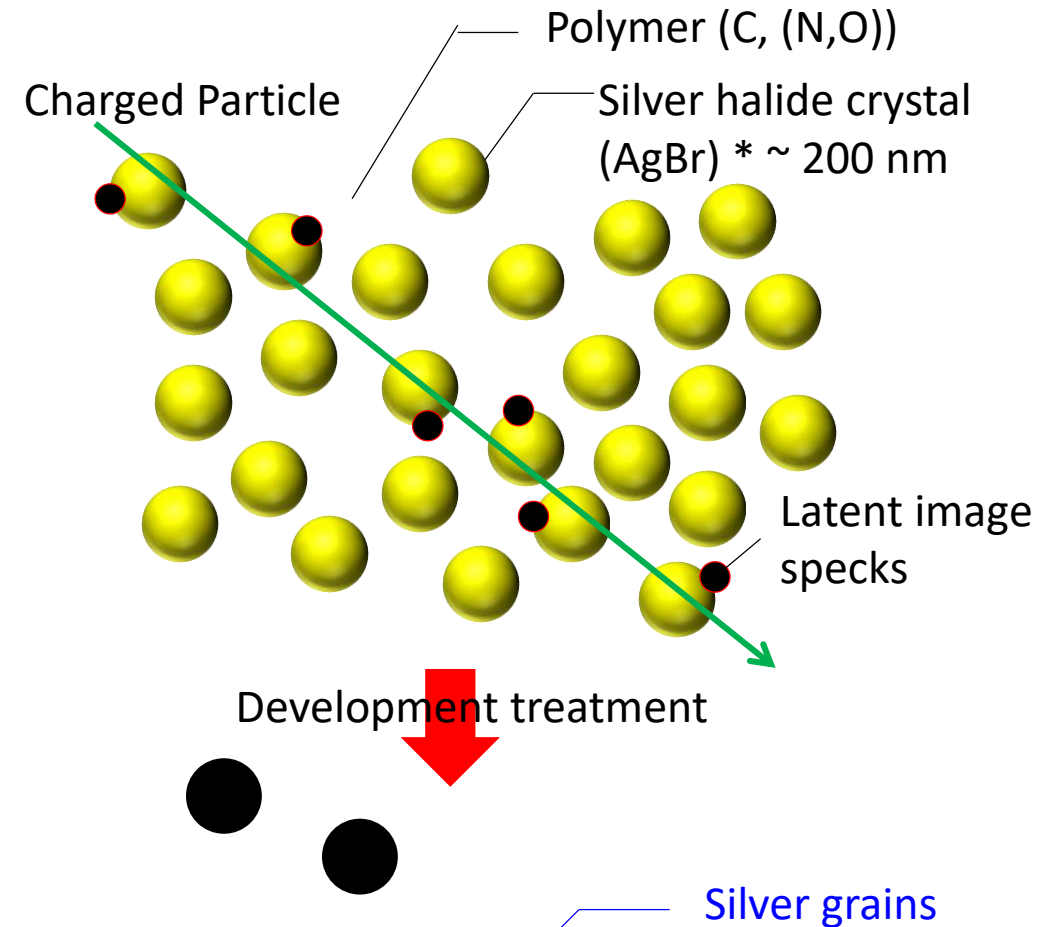
<https://arxiv.org/abs/1604.04199>

# Concept of NEWSdm experiment



# Nuclear Emulsion Device

- Kind of photographic film
- High spatial resolution
- $4\pi$  tracking



Standard nuclear emulsion

Crystal size : 200 nm

Detectable track length :  $> O(1) \mu\text{m}$

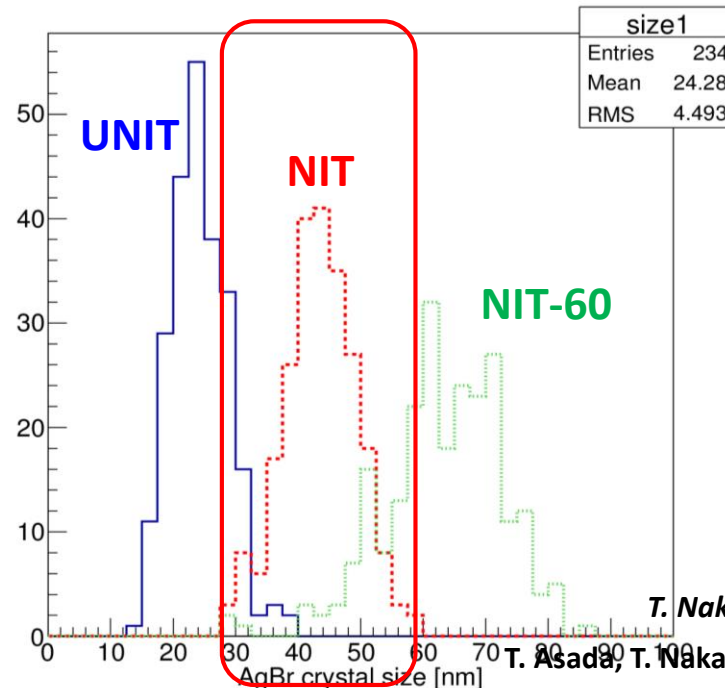
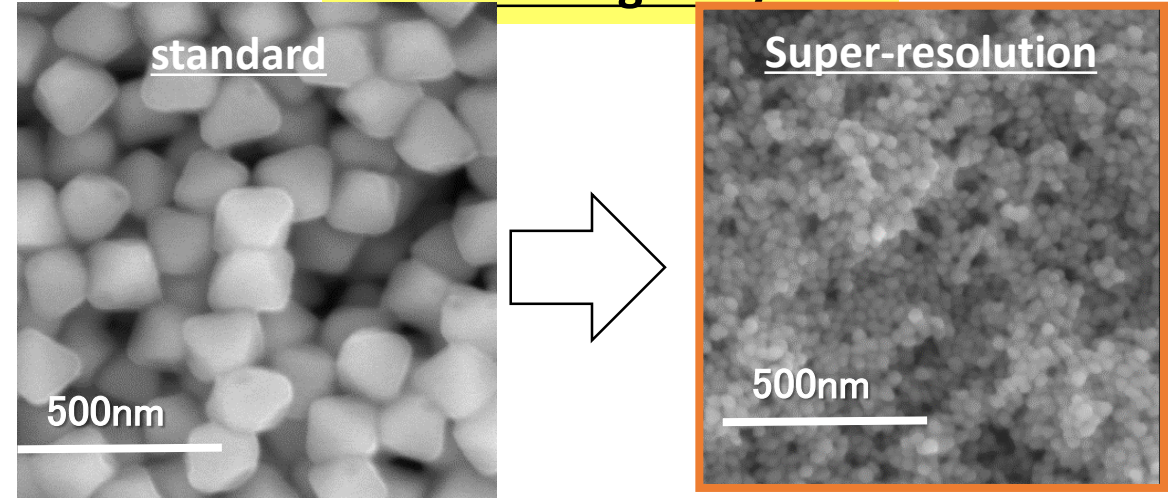
**Very fine crystal controlled about 10 nm to detect 100 nm scale tracks**

# Self-production of Nano Imaging Tracker(NIT)



- Production time : 4-5 hours /batch
  - One butch : ~ 100 g (+ 300 g)  
(there are 2 type machines)
- ⇒ kg scale production is possible using this machine.

## Controlled AgBr crystal



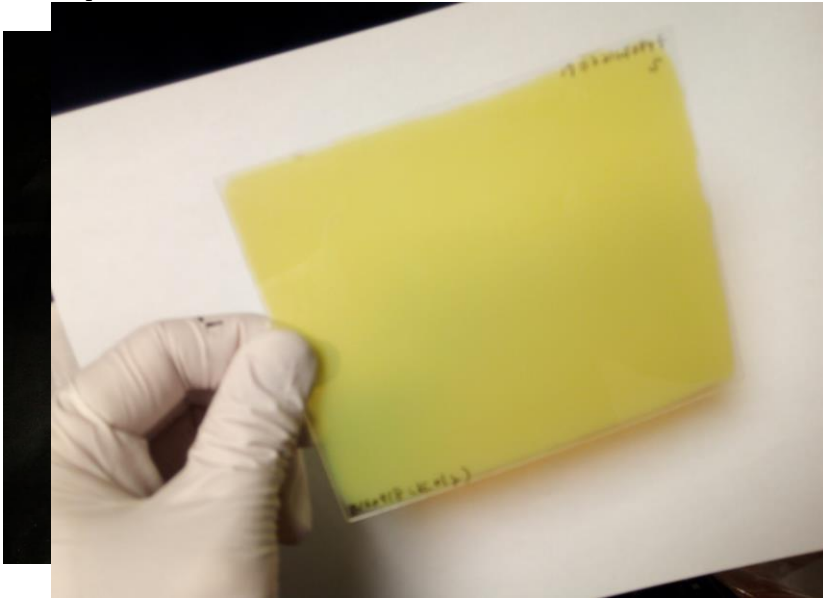
Current standard Device :  
Nano Imaging Tracker [NIT]  
crystal size : 44 nm

Finest grain emulsion :  
Ultra-NIT [UNIT]  
crystal size : 25 nm

T. Naka et al., Nucl. Inst. Meth. A 718 (2013) 519-521

T. Asada, T. Naka + , Prog Theor Exp Phys (2017) 2017 (6): 063H01

# Properties of NIT device



## Elemental composition of NIT

	Mass fraction	Atomic Fraction
Ag	0.44	0.10
Br	0.32	0.10
I	0.019	0.004
C	0.101	0.214
O	0.074	0.118
N	0.027	0.049
H	0.016	0.410
S, Na + others	~ 0.001	~ 0.001

For high-mass DM s

For low-mass DM

## ◆ Intrinsic radioactivity :

U-238	Th-232	K-40	Ag-110m	C-14	
27	6	35	(~400)	24000	[mBq/kg]

- K-40 reduction : 69020 (first type) → 35 mBq/kg  
by KBr → NaBr for AgBr creation and use high deionized gelatin
- Ag-110m : not confirmed yet
  - first measured batch : ~ 400 mBq/kg
  - recent batch : < 150 mBq/kg
- C-14 : AMS measurement result. Consistent with natural abundance.  
→ if replace to synthetic polymer, it will be reduced more than  $10^{-3}$

## ◆ Intrinsic neutron background (SOURCES + Geant4):

	Emission [/kg/y]	Rate for > 100 nm tracks [/kg/y]
Intrinsic neutron	~ 1.2	~ 0.1

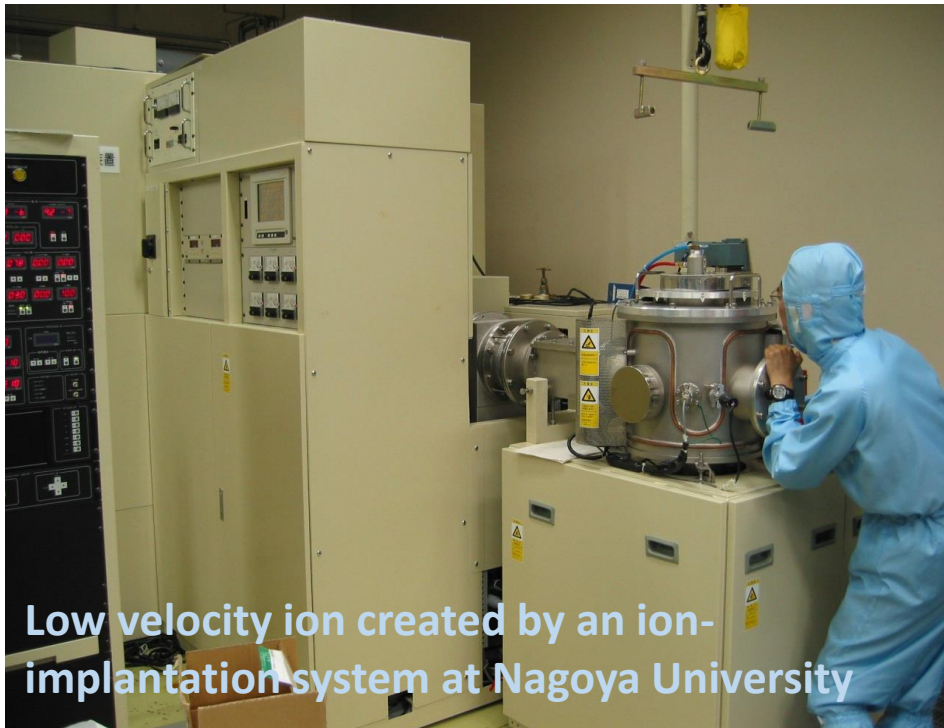
Detail shown in Astropart. Phys. 80 (2016)16-21



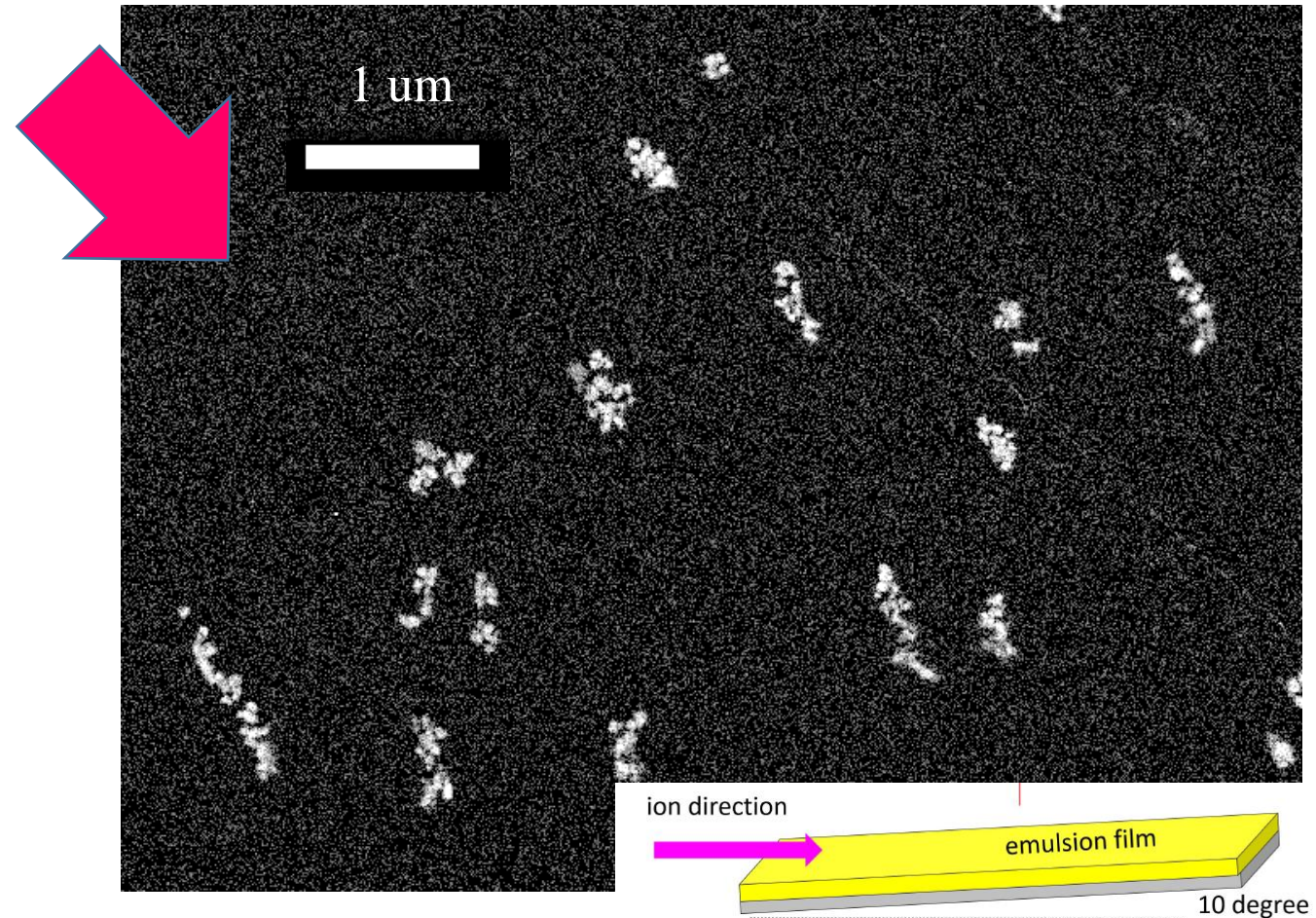
# Low-velocity ion tracking

Can use ion implantation as calibration source

- Mono energy ( $\pm 0.1$  keV)
  - Good direction uniformity ( $<10$  mrad)
  - Now, C from  $\text{CO}_2$  • Ar, Kr
- ( various kind ions are also possible)

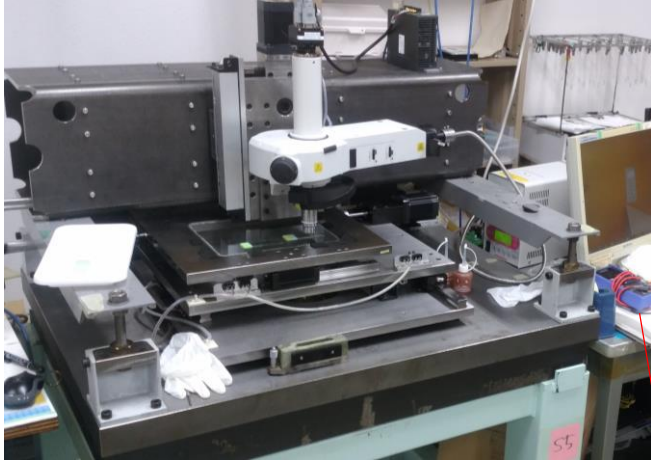


SEM image of low-velocity Carbon ion (100keV)

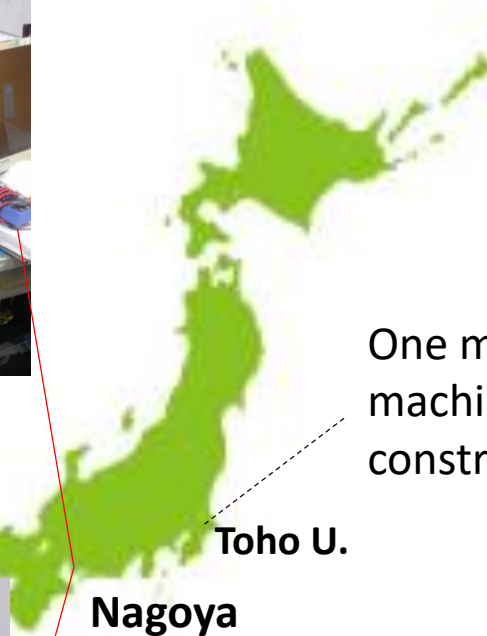
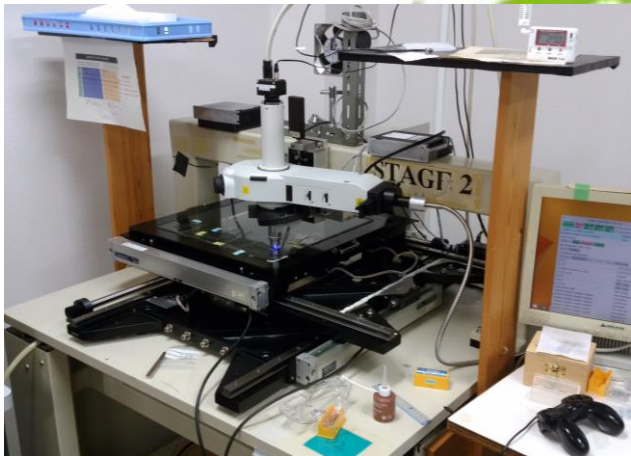


AgBr crystal has good sensitivity about Carbon (100 % consistent sensitivity)

# Readout technologies



- Event selection
- Phase contrast imaging



Toho U.

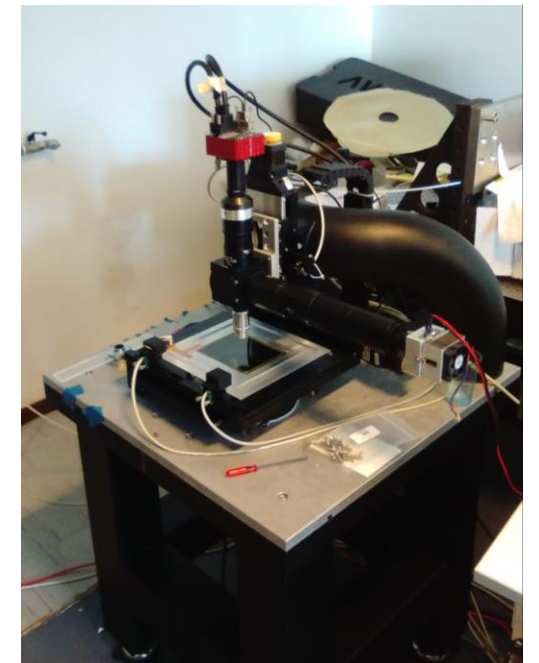
Nagoya

One more machine will be constructed

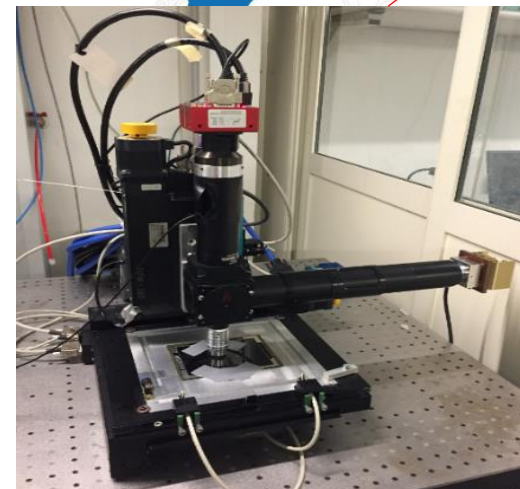


LNGS

Napoli



Machine for device quality check



- Event selection
- Plasmon analysis

x 2

# Optical microscope system and analysis flow

Standard optical microscope scanning [on going]

**Current Speed : ~30 g/y**

- Elliptical event selection
- Roughly event selection with high speed
- On-line event analysis

**~ 100 g/month scale (~ kg/y)**

**~ kg /month scale (~ 10 kg/y)**

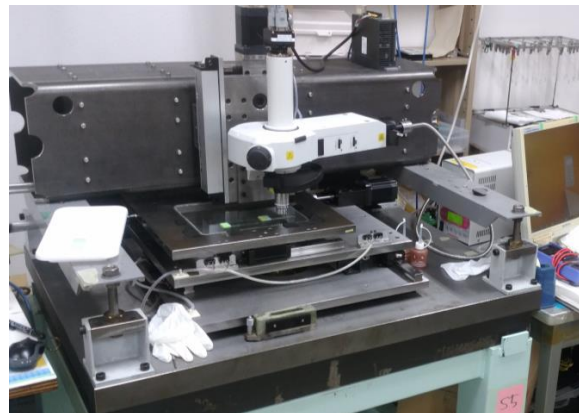


Phase contrast imaging [will be newly installed]

**10<sup>7</sup> events/month**

- Phase contrast imaging
- Contaminated dust discrimination

**To be constructed soon**

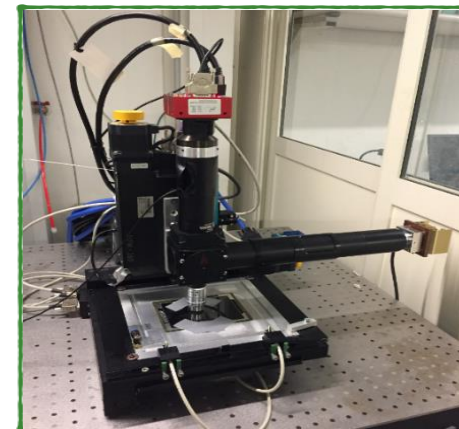


LSPR analysis [under studying]

**10<sup>5</sup> events/month**

- Super-resolution : ~10 nm
- Spectrum analysis
- Machine learning

Yandex@Russia, Napoli



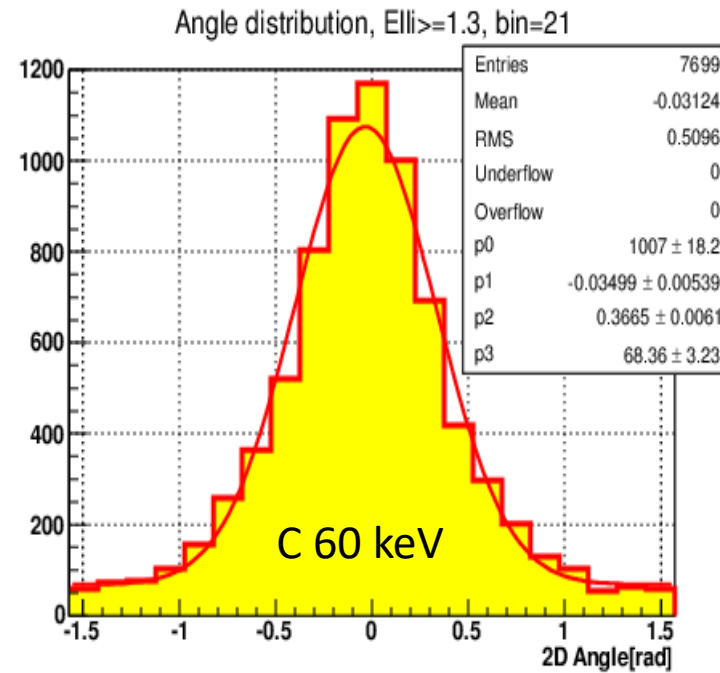
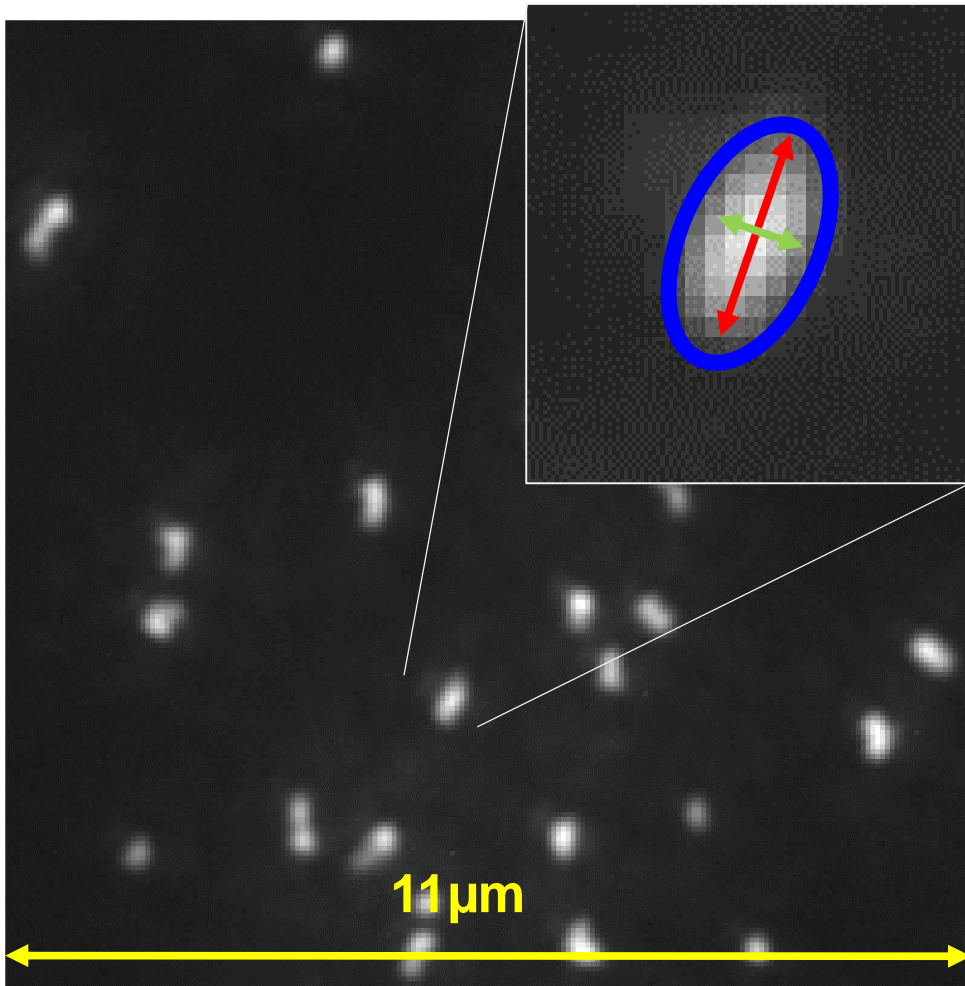
Further new analysis [ under studying ]

**~10<sup>3</sup> events/month**

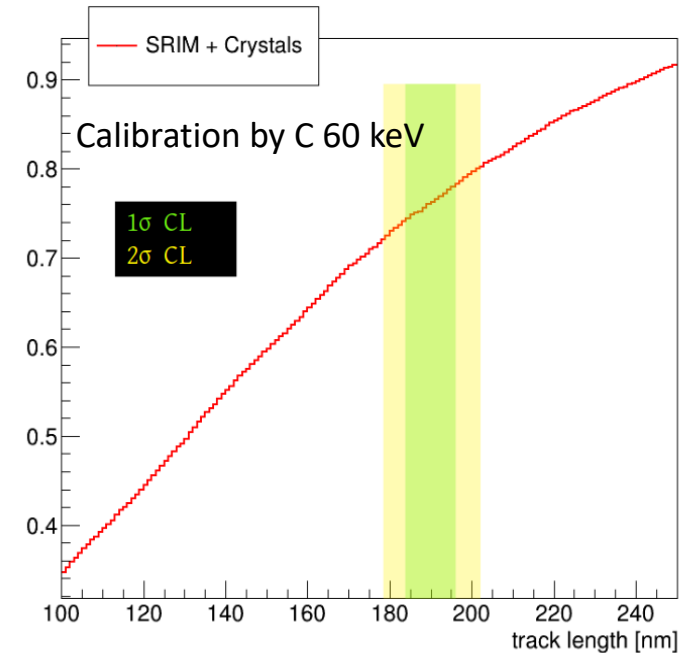
- 3D super-resolution analysis with plasmonics
- Destructive analysis using oxidation method
- Expansion method

**Cutting-edge technologies will be installed**

# Sub-micron length track readout capability



Clearly observed angular distribution  
⇒ angular resolution  $\sim 30$  deg.



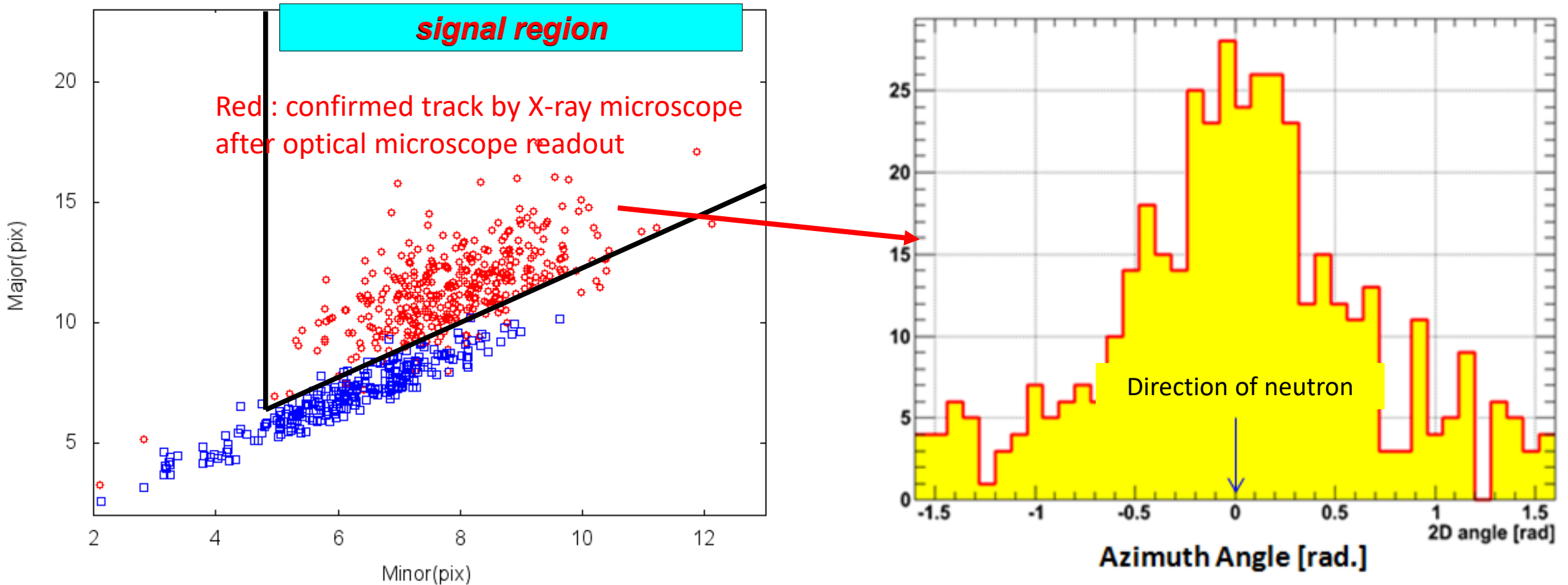
**Direction sensitive track length threshold in this algorithm**  
⇒  $> \sim 190$  nm

**Energy threshold**  
 $> \sim 60$  keV (eff.  $\sim 10\%$  ⇒ to be improve by upgrade optical condition)

*K. Kimura and T. Naka, Nucl. Inst. Meth. A 680 (2012) 12-17*

*T. Katsuragawa et al, JINST 12 T04002 (2017)*

# Demonstration of direction sensitive nuclear recoil detection due to 14.8 MeV neutrons



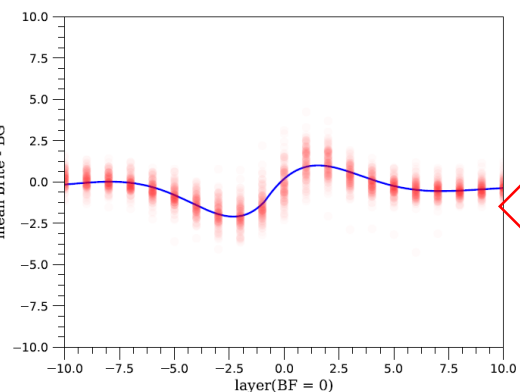
**Mostly detected target was Br recoil [ < 200 keV ] → difference condition from current one  
Now on studying CNO recoil demonstration due to 565-700 keV (Li-p nuclear fission reaction)**

	Main source	Technologies	Expected rejection power or event rate
<b>Physical BG</b>			
Electrons	C-14 $\beta$ Environment gamma	Crystal temperature dependence ( <i>M. Kimura et al., NIM A 845 (2017) 373</i> ) Crystal sensitivity control Image and plasmonic analysis	(> 10 <sup>6</sup> or more rejection power (< O(1) /kg/day)) *now on studying
		Synthetic Polymer	> 10 <sup>3</sup> or more
Neutron	Intrinsic ( $\alpha, n$ )	-	$\sim 3 \times 10^{-4}$ /kg/day or less Astropart. Phys. 80 (2016)16-21
	Environment	Water shield	< 1E-4/kg/day
Cosmic-ray	Recoiled nuclei	Coincidence with MIP sensitive emulsion	*on studying using simulation
	Spallation neutron	(under studying with simulation)	( $\sim O(10^{-4})$ )/kg/day * now on study)
<b>Nonphysical BG</b>			
Contaminated dust	(under studying)	Clean room Phase contrast imaging Plasmonic analysis and image processing Machine learning Chemical treatment	Under studying (at least > 10 <sup>6</sup> or more, in principle it should not be background )

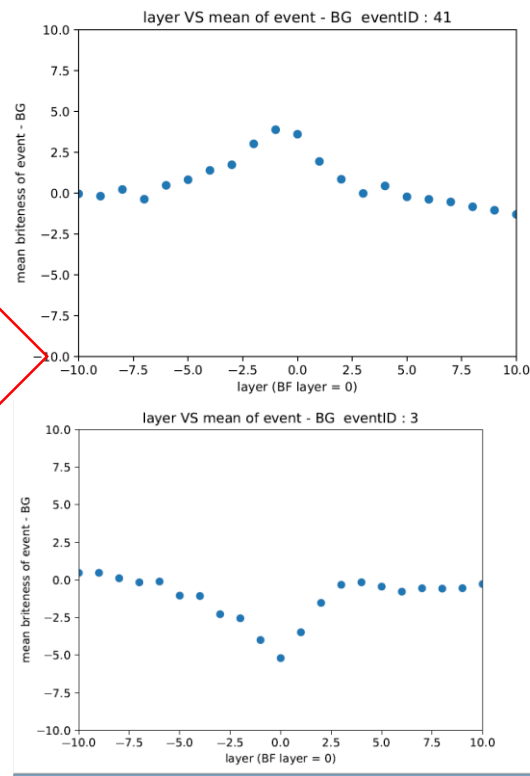
# Further signal discrimination from backgrounds

## Phase contrast imaging

### Low-velocity ion (signal)



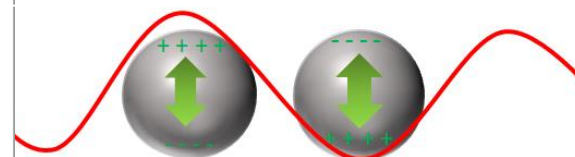
### Contaminated background



New information to distinguish signal from background by phase-contrast imaging

## Plasmonic optical response + machine learning

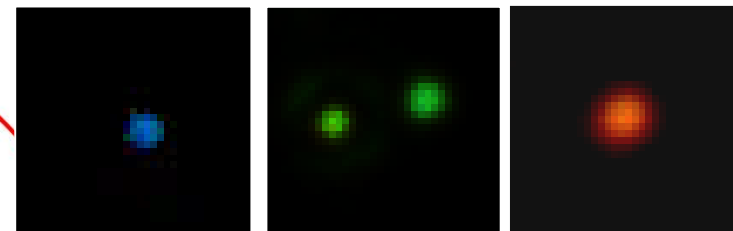
### Localized Surface Plasmon Resonance



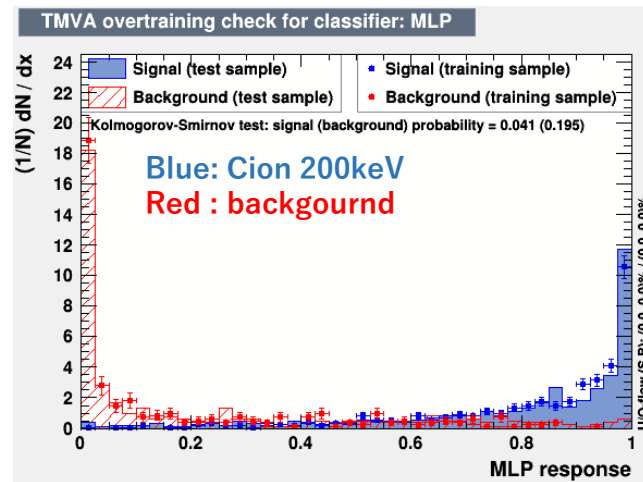
silver nano particle

$$p = 4\pi\epsilon_m a^3 \frac{\epsilon_1(\lambda) - \epsilon_m(\lambda)}{\epsilon_1(\lambda) + 2\epsilon_m(\lambda)} E_0$$

$$\epsilon_1(\lambda_j) + 2\epsilon_m(\lambda_j) \approx 0$$

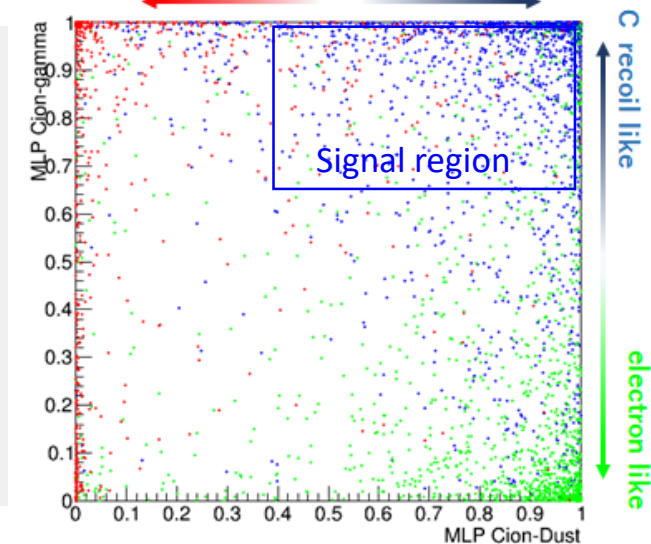


Scatter light spectrum information due to plasmonic effect



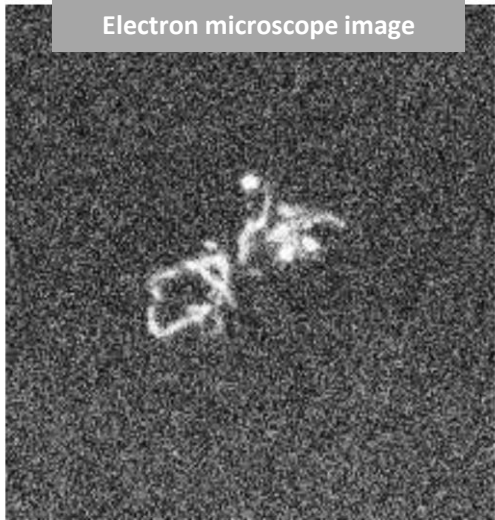
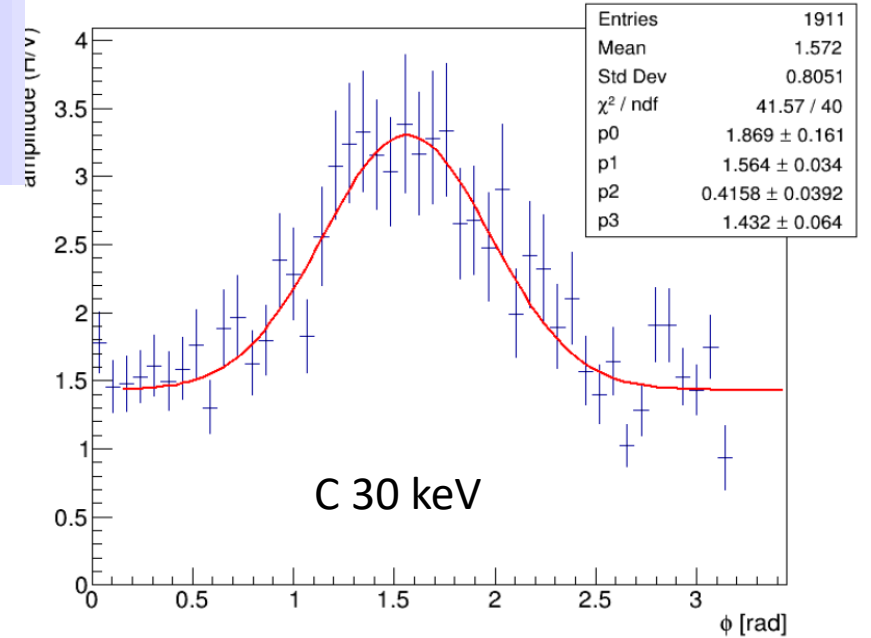
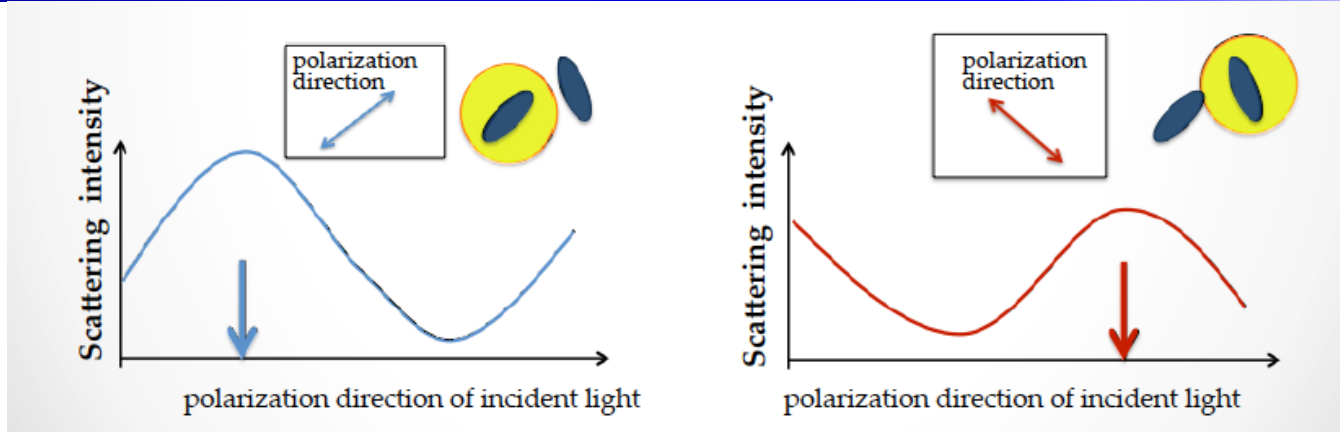
Background (dust) like ←

→ C recoil like

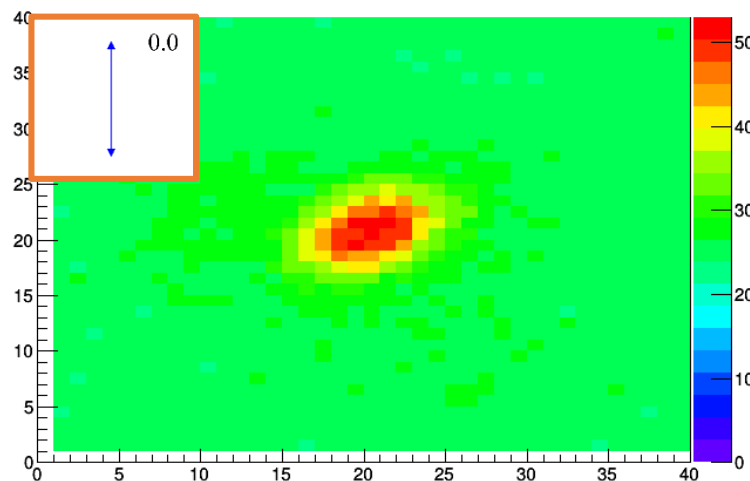


Such new analysis studies are now on going

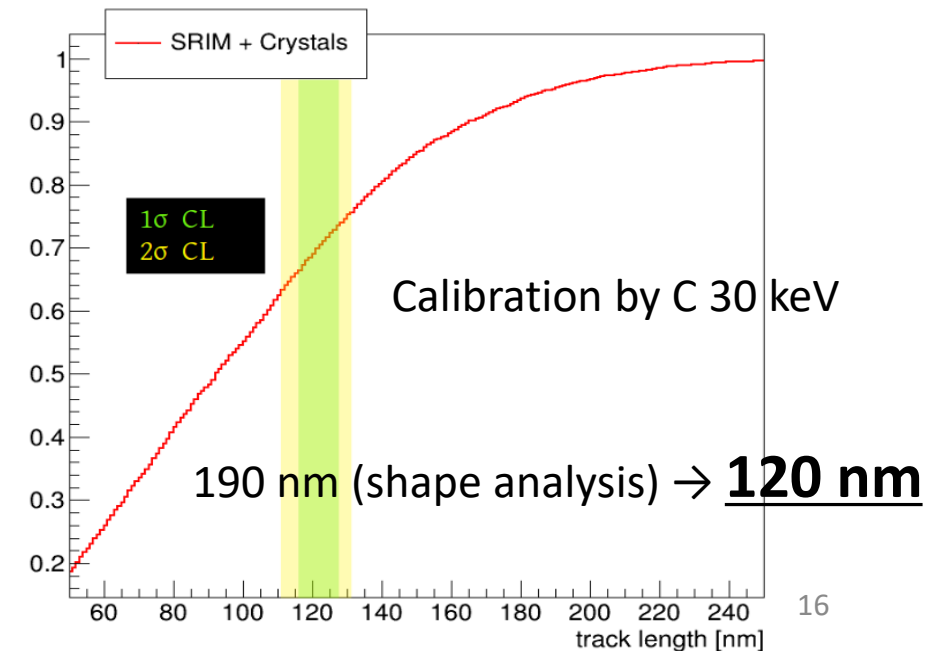
# Super-resolution microscopy using LSPR information toward lower-threshold tracking



## Polarization light dependence



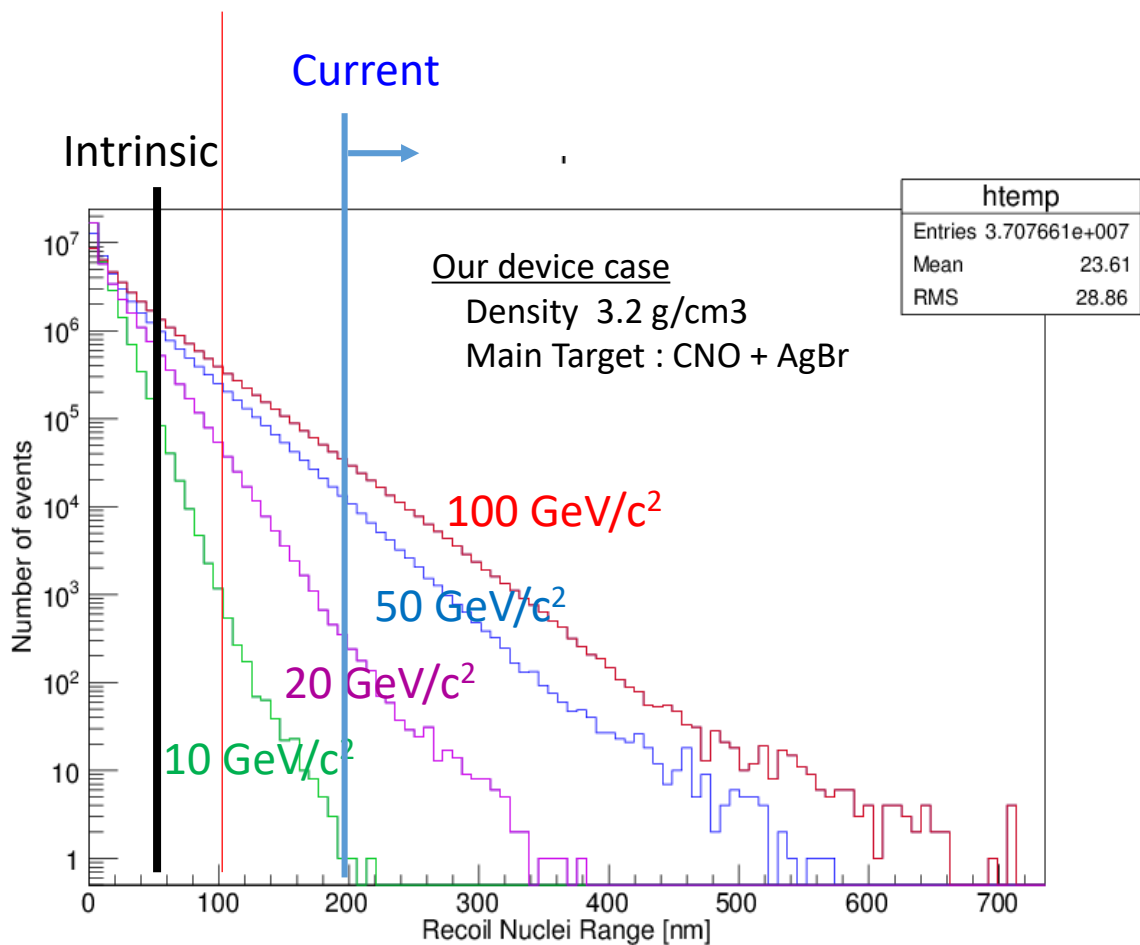
Shift of barycenter is important information for nano-scale structure



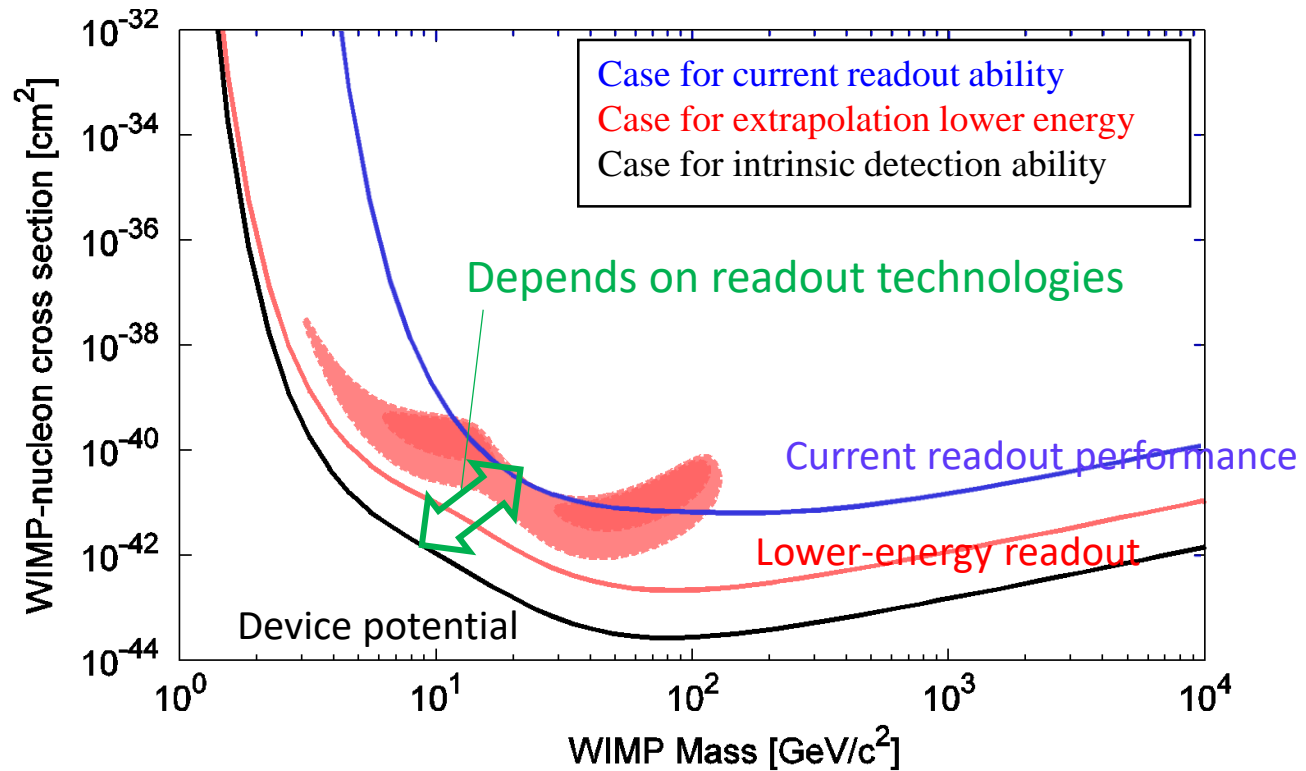


# Dark matter sensitivity

Demonstrated new tech.



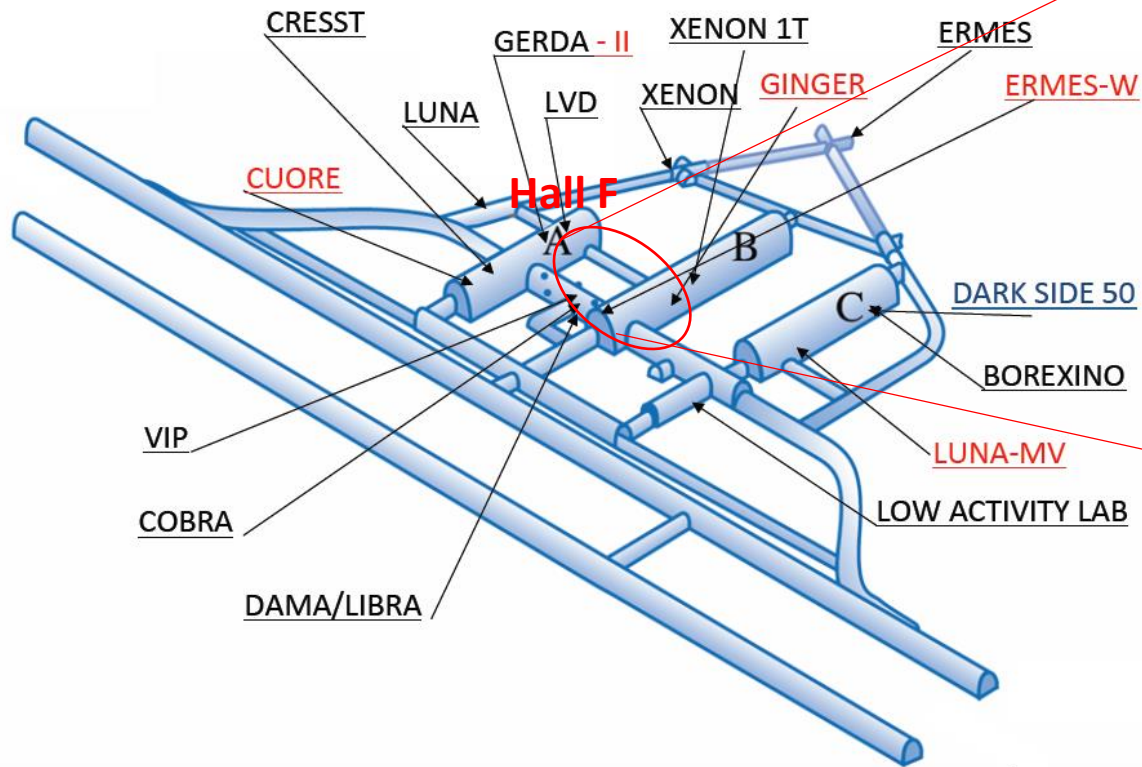
10 kg·year simulated sensitivity [90 % C.L.] + zero BG



NIT detector / CNO sensitive / no Bkg no directionality  
Simulation limit is “energy > 5 keV for all atoms (SRIM limit)”  
& “Sensitivity > 0.1 % (Simulation statistics limit;10 event)”

**Device potential : 10 keV of C recoil (> ~ 10% eff. and 45° angl. Res.**

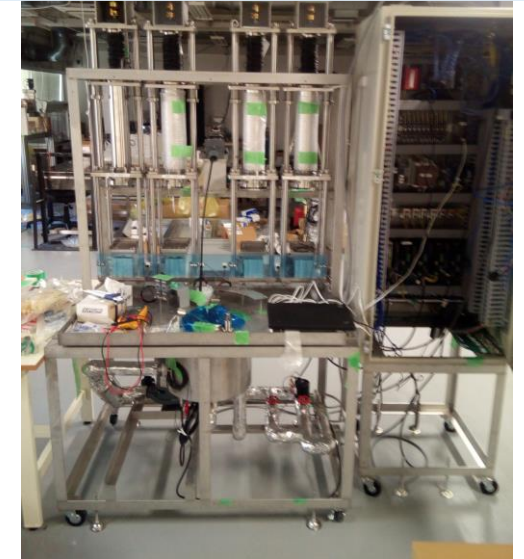
# Underground laboratory at LNGS



Device Production facility



New production machine



## Motivation of New Underground facility

- Device self-production in underground
- Device handling in clean room
- Chemical development

Discussion started from 2017, and construction from beginning of 2018

# New Underground emulsion facility

Feb. 2018 ~ : started construction and commissioning of the production machine at Nagoya (⇒ transported to LNGS from Sep. 2018)

**Feb. 2019 ~ : Started test production first time at underground  
+ clean room and other infrastructure are on constructing**

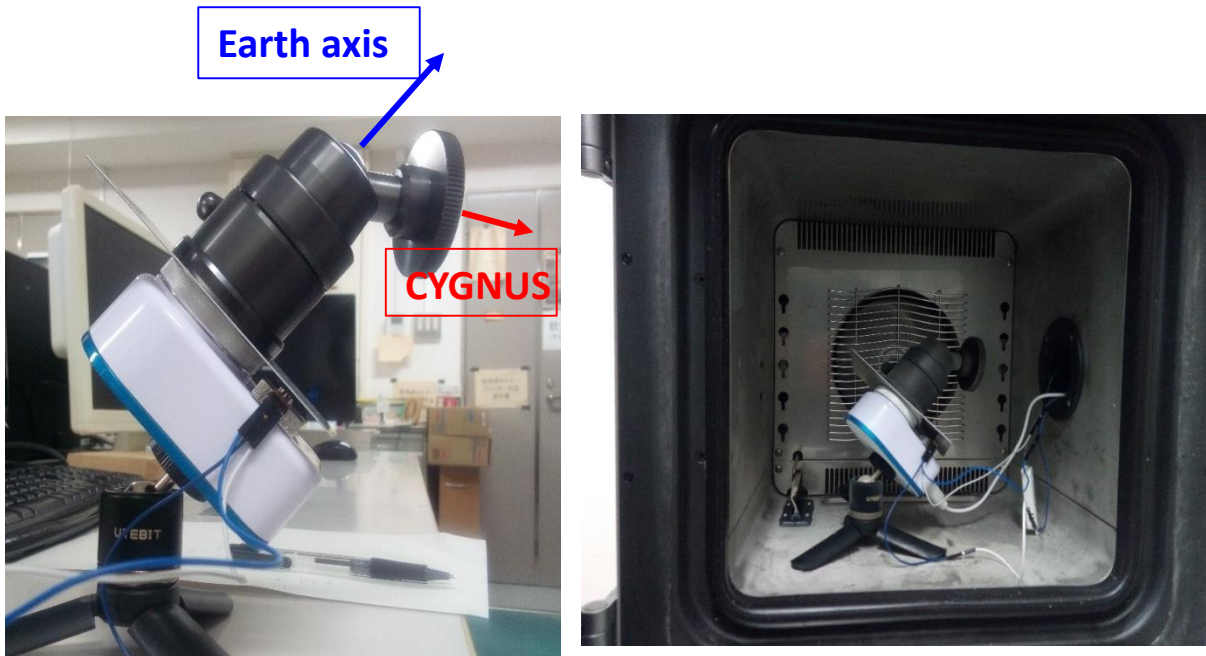
**Up to April : overall confirmation of underground emulsion  
facility with clean room**



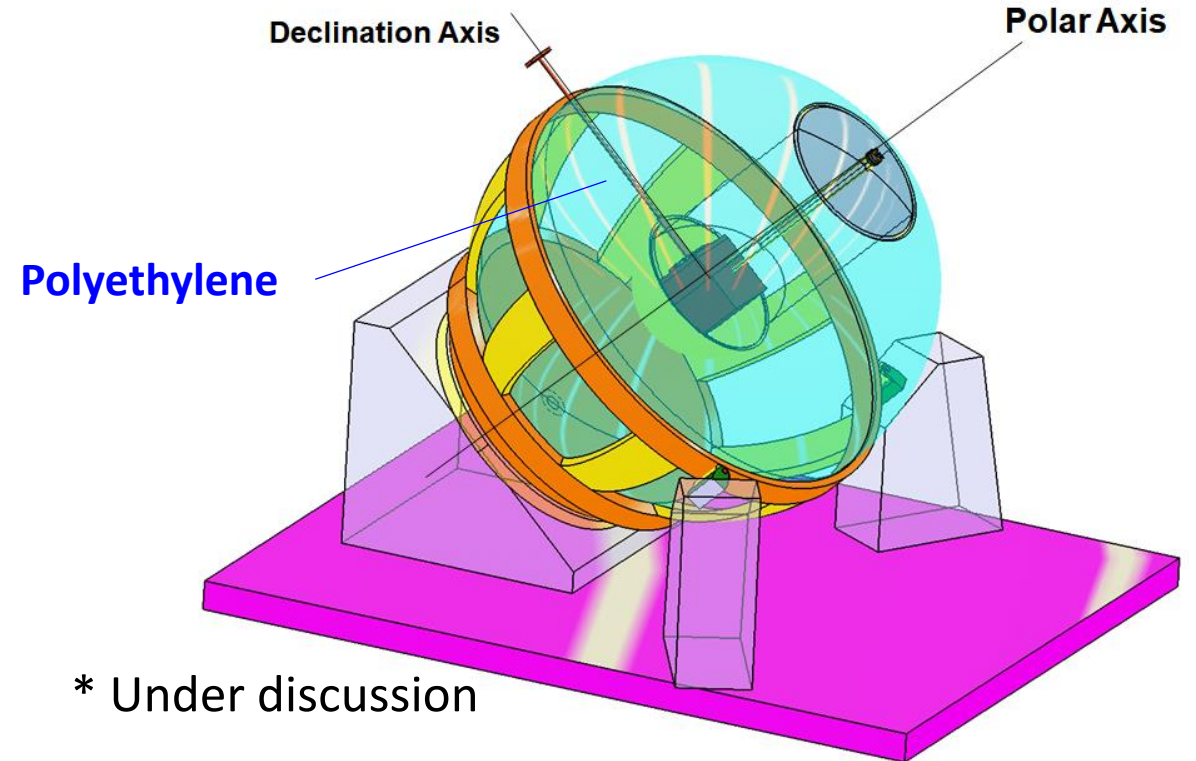
**First production in LNGS succeeded !!**

# Equatorial telescope for directional search

Future prospect for ~ 10-100 g scale detector



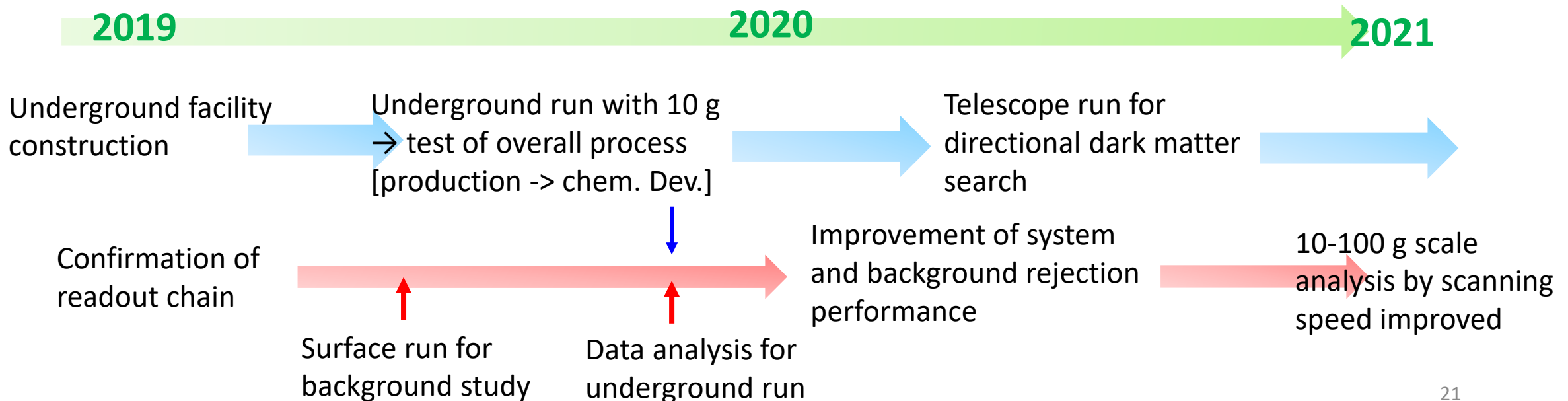
Future prospect for > 1kg scale detector



Source	Rate [/10kg/y]
Environmental $\gamma$ -rays	$(2.0 \pm 0.2) \times 10^4$
Environmental neutrons	$O(10^{-2})$
Cosmogenic neutrons	$1.4 \pm 0.1$ <sup>20</sup>

# Conclusion

- ❑ NEWSdm project is for direction-sensitive dark matter search with super-resolution nuclear emulsions as solid detector
- ❑ Device production and readout system demonstration have been done, and optimization and overall system are now on constructing and commissioning
- ❑ New underground facility with device production machine and clean room is now on constructing, and it will be ready around June, 2019.
- ❑ We will do underground experiment test there, and go forward for larger scale directional dark matter search

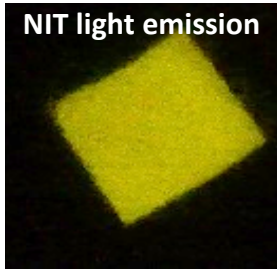


Back up

# Detector Application

## [Scintillation light emission]

- ✓ High emission efficiency  
→ possibility as scintillator
- ✓ Study for fundamental mechanism of AgBr nano crystal



T. Shiraishi, H. Ichiki, TN et al., accepted (2019)

## low-velocity heavy particle detector

- ✓ Exotic heavy low-velocity particle (e.g., monopole)
- ✓ Medical therapy
- ✓ Energy loss mechanism

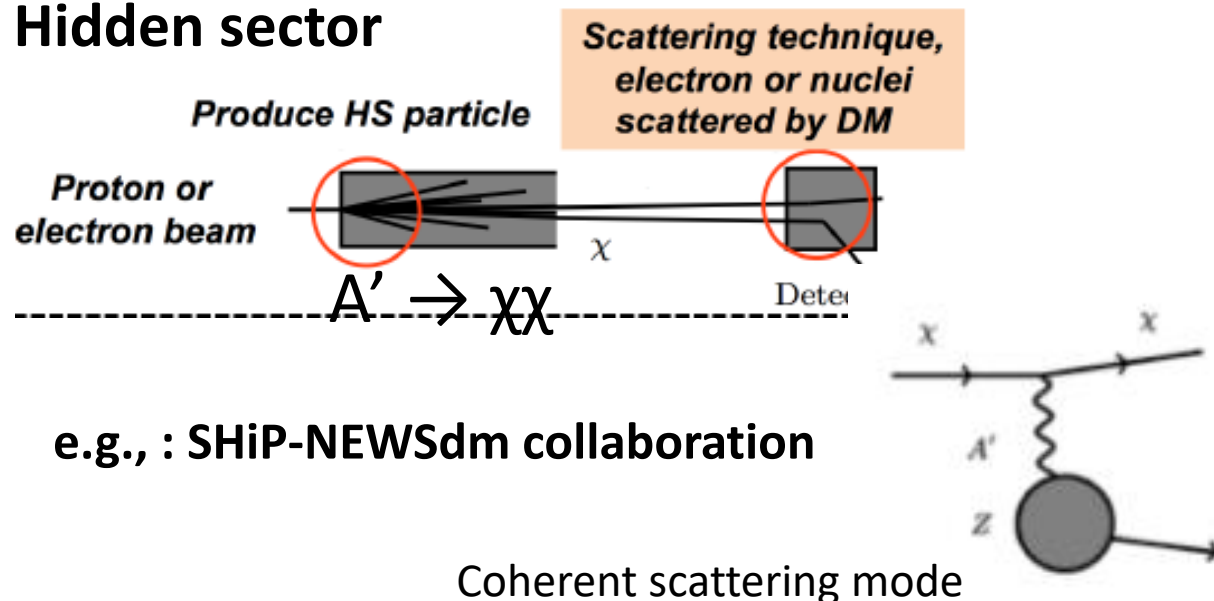


## [Neutron detector]

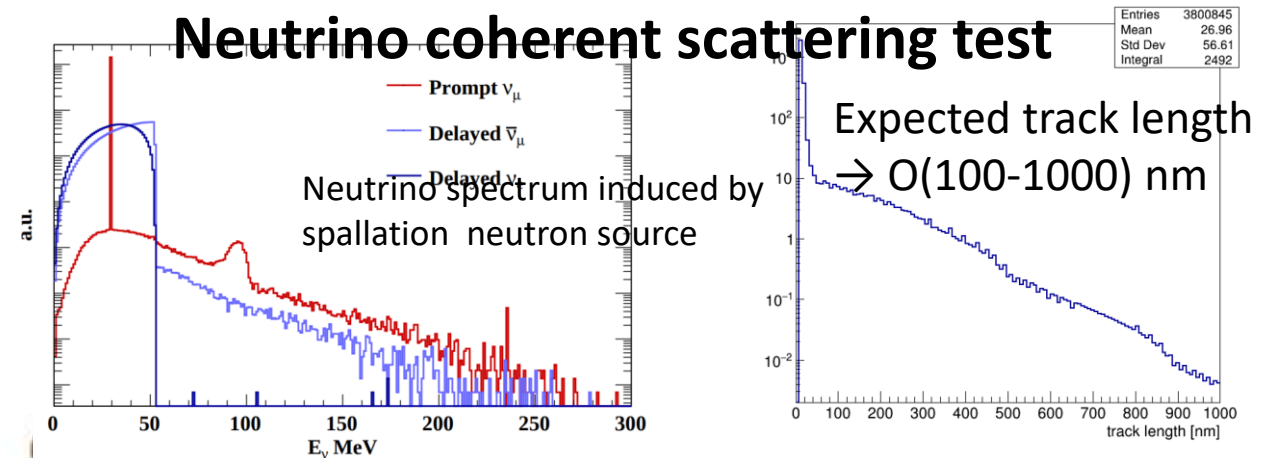
- ✓ Environment neutron measurement with direction information
- ✓ Low-energy (sub-MeV, UCN) neutron detector

# NEWSdm Application

## Hidden sector

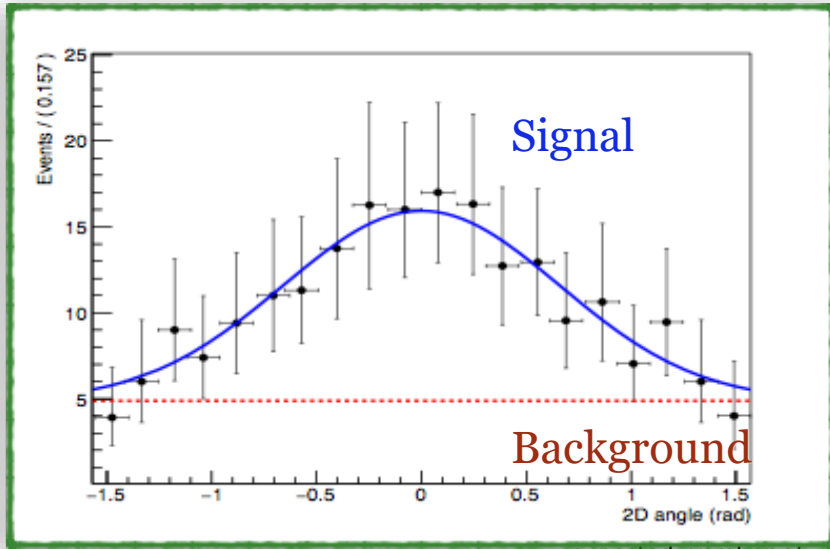


## Neutrino coherent scattering test

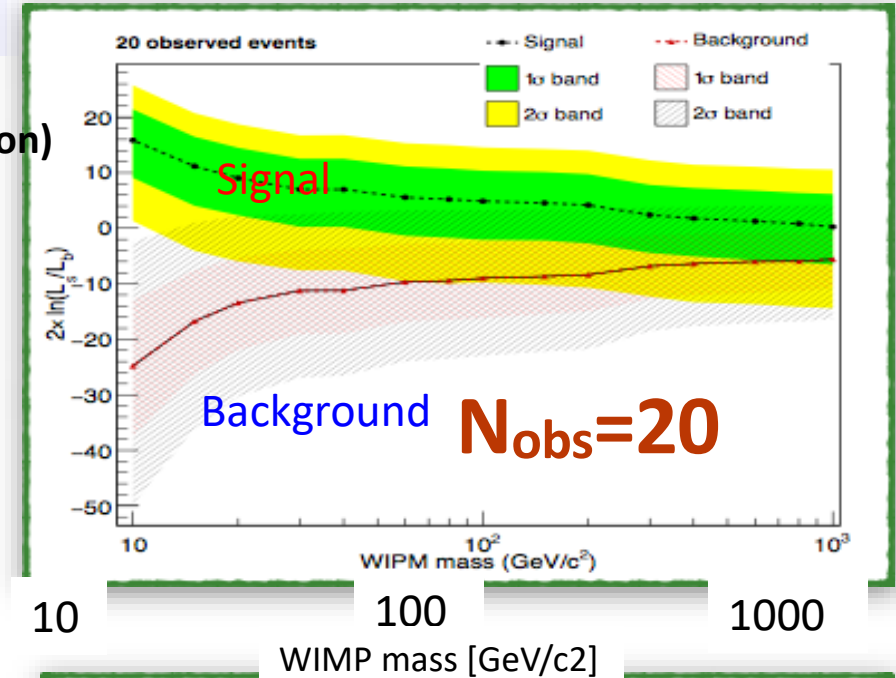
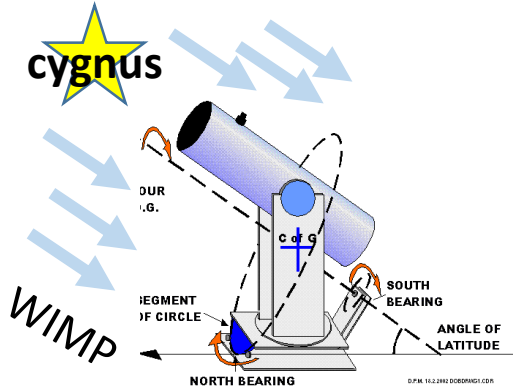


- ✓ Demonstration of neutrino floor for dark matter search
- ✓ High precision Demonstration using weak interaction of Dark matter search

# Potential of Directional Sensitive Search



N. Agfanova *et al.* (NEWSdm collaboration)  
Eur. Phys. J. C (2018) 78: 578

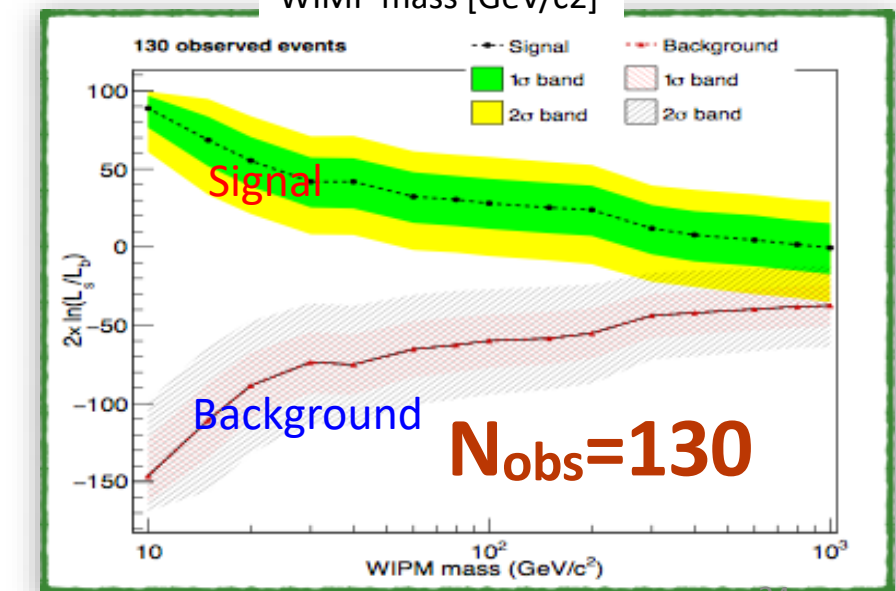


$$\mathcal{L}(\sigma_{\chi-n}, R_b) = \frac{e^{-(\mu_{\chi} + \mu_b)}}{N!} \times \prod_{i=1}^N [\mu_{\chi} f_{\chi}(\vec{q}_i; t_i) + \mu_b f_b(\vec{q}_i)]$$

expected number of WIMP events  $\rightarrow \mu_{\chi}$   
 expected number of background events  $\rightarrow \mu_b$   
 signal pdf  $\rightarrow f_{\chi}$   
 background pdf  $\rightarrow f_b$   
 total number of observed events  $\rightarrow N$   
 set of observables  $\rightarrow \vec{q}_i, t_i$

**Direction information : Several 10 events**

**Gain of 100 times**



**Annual modulation : Several 1000 events**

WIMP mass [GeV/c<sup>2</sup>]