

Summary of recent updates for SN burst detection at SK

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For Super-Kamiokande Collaboration
@ 10th SN workshop (Okayama Univ.)

Contents:

- Introduction
 - SK Gd status
 - Supernova monitor at SK
- Improvements of SN burst detection
 - GCN Notice
 - Study of SK performance for SN models
 - SN direction fitter improvement
- Summary

Super-K experiment

1000m underground = 2600 m.w.e

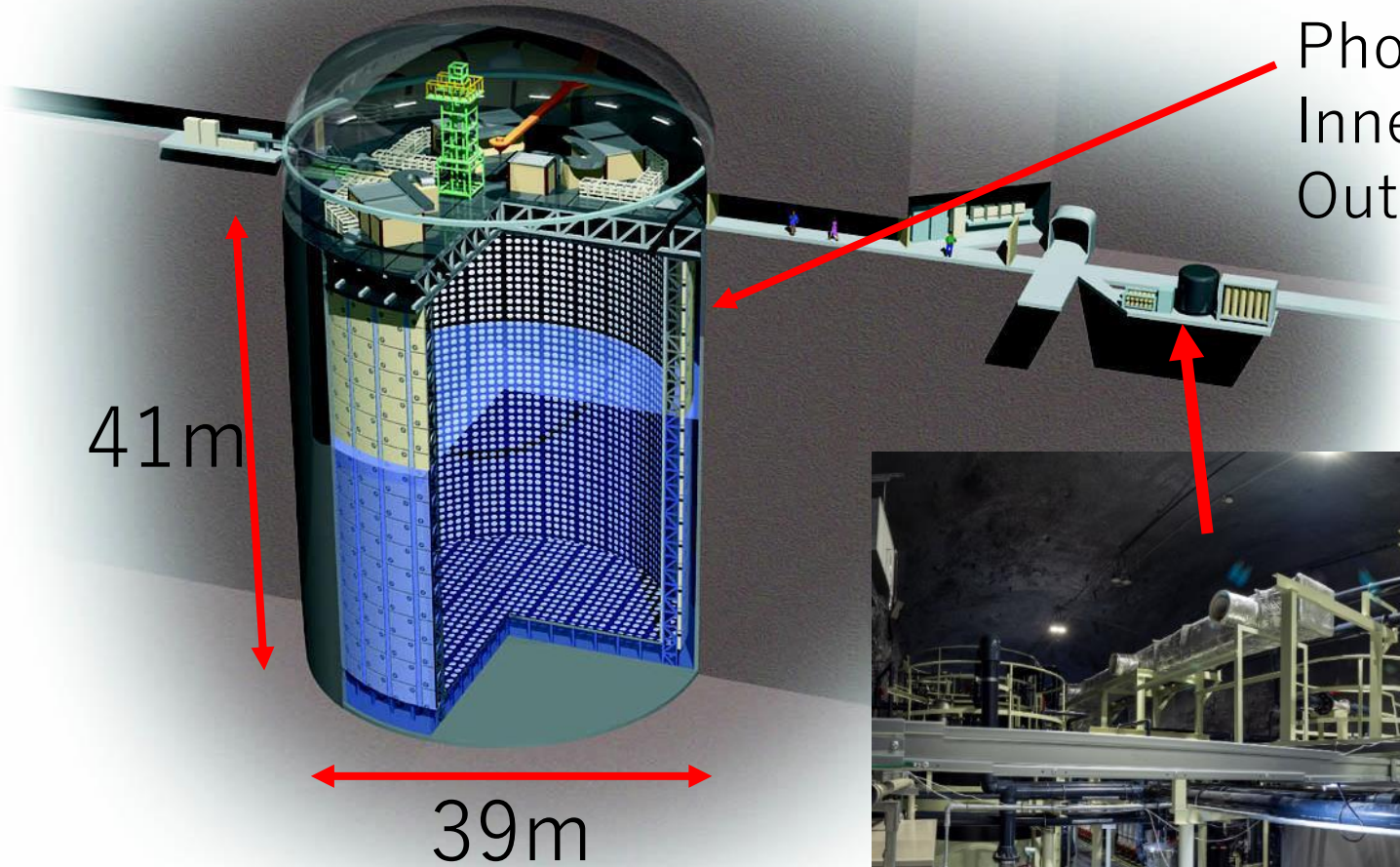


Photo sensors :

Inner detector: 11129 20inch PMTs

Outer detector: 1885 8inch PMTs

Gd water system room



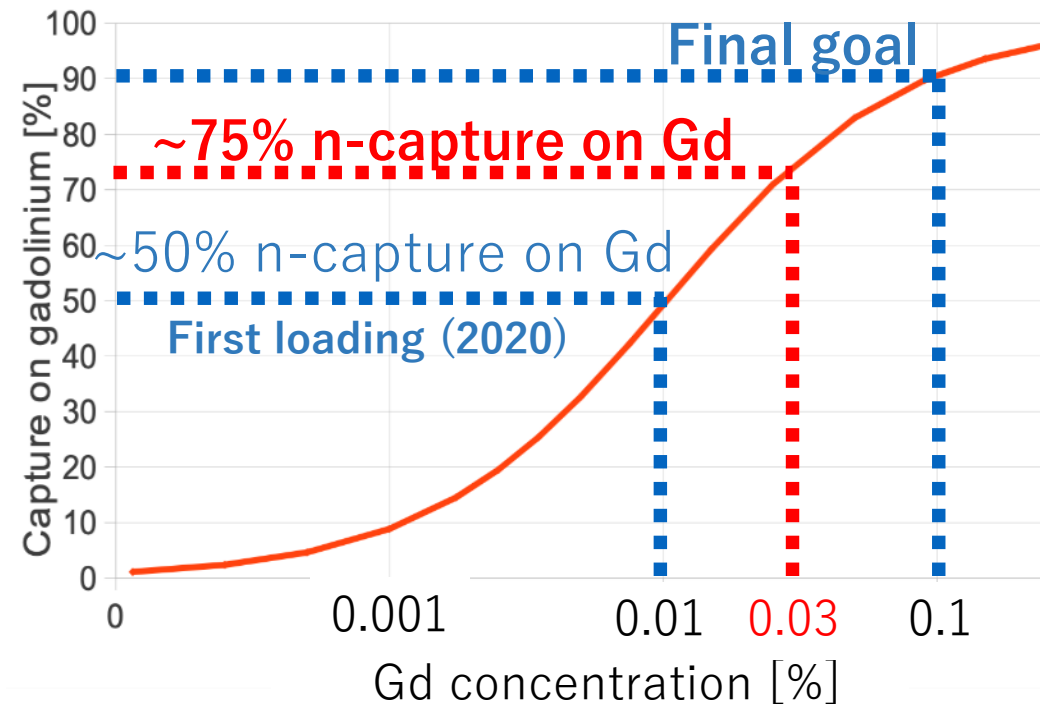
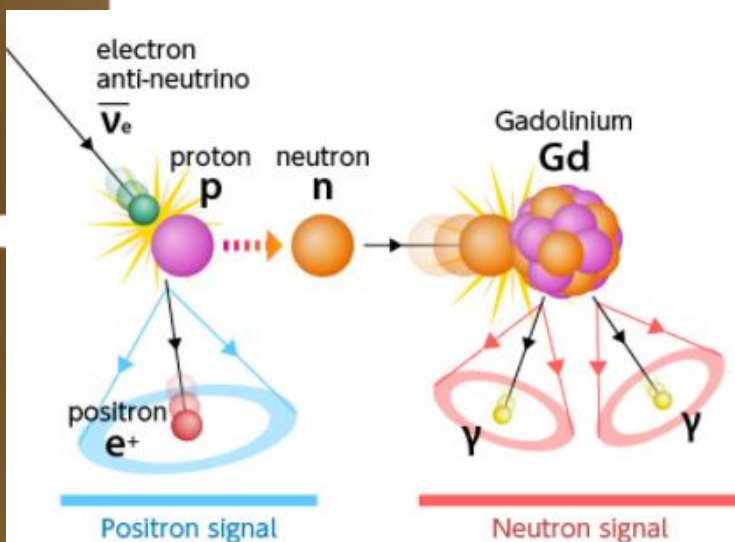
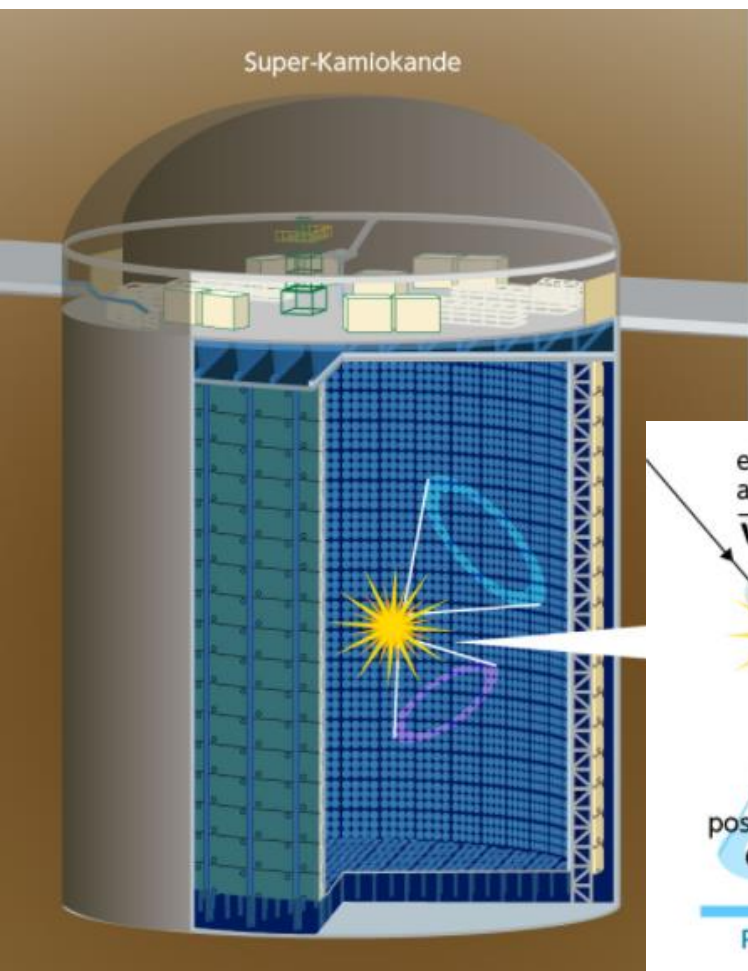
SK-Gd project

Dissolving Gd to enhance detection capability of neutrons from ν interactions

Phys.Rev.Lett. 93 (2004) 171101

Physics targets:

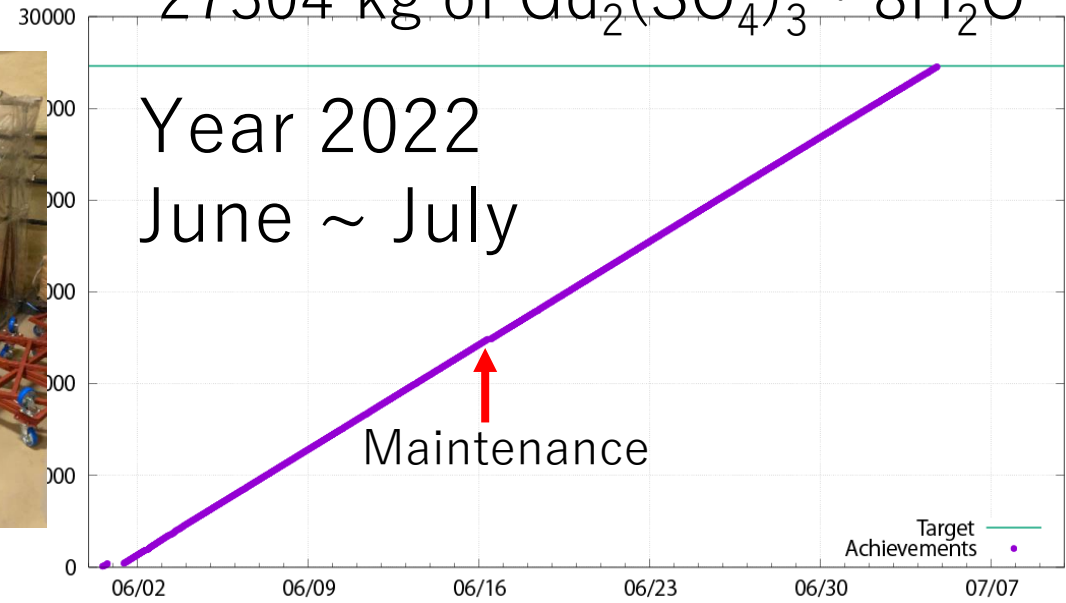
- (1) Discovery of Supernova relic neutrino (or DSNB)
- (2) Galactic supernovae (pointing accuracy, and pre-SN ν)
- (3) Reduction of BG for proton decay, solar ν , or reactor ν
- (4) Neutrino/anti-neutrino discrimination



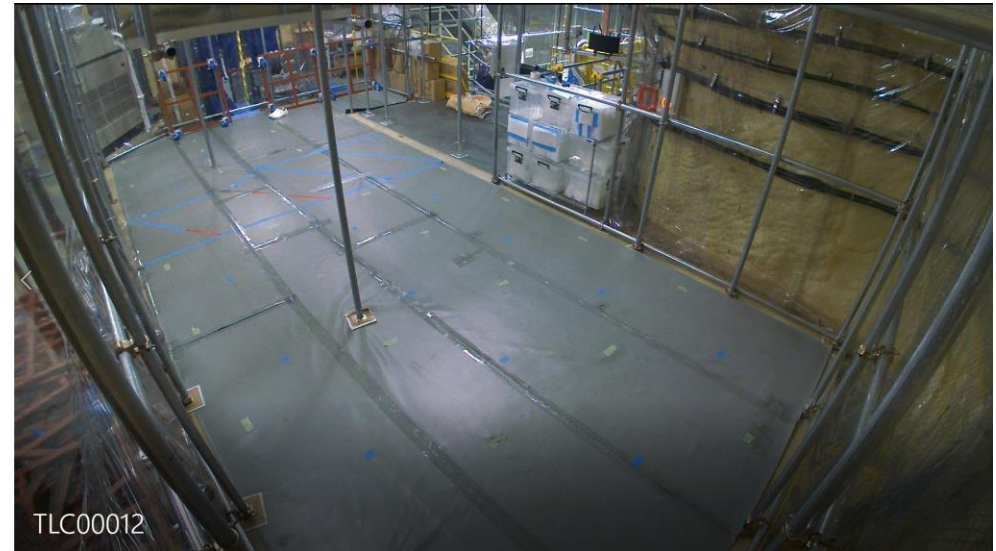
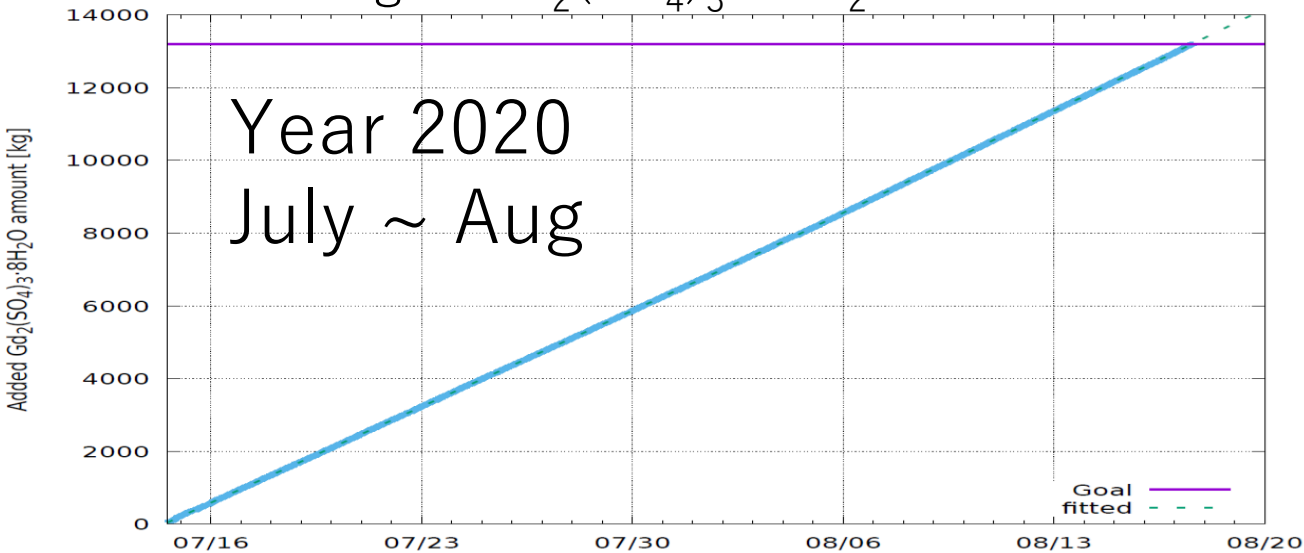
Gd loading in 2020 and 2022

Beginning of the dissolving work in 2022

27304 kg of $Gd_2(SO_4)_3 \cdot 8H_2O$



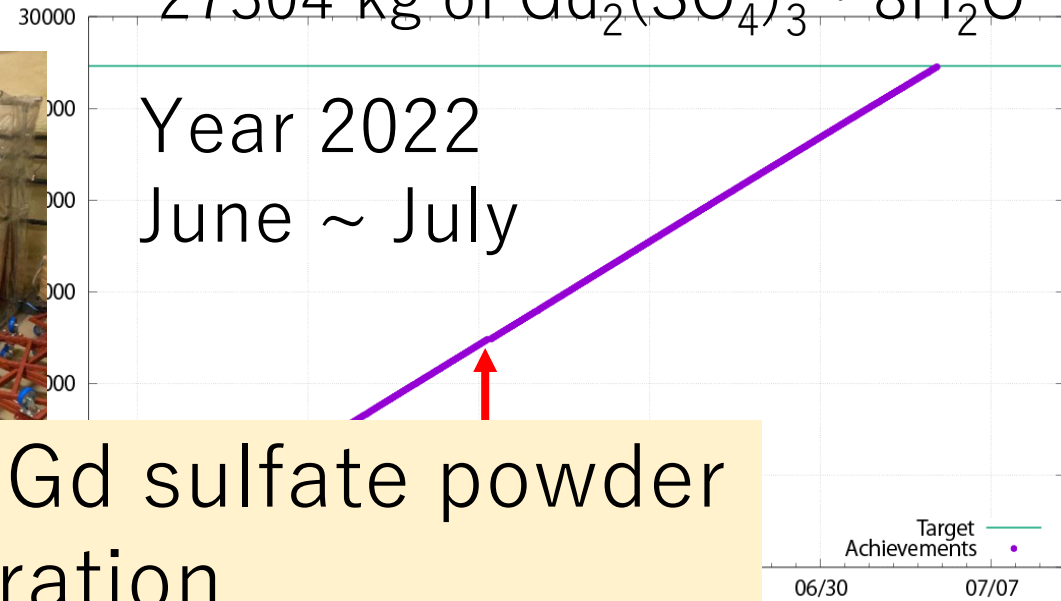
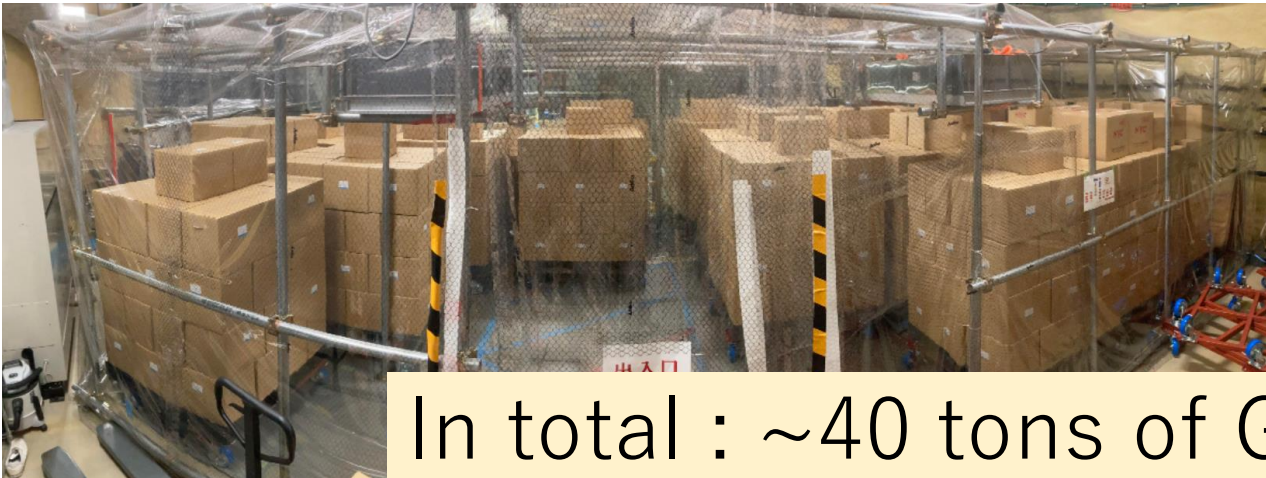
12884 kg of $Gd_2(SO_4)_3 \cdot 8H_2O$



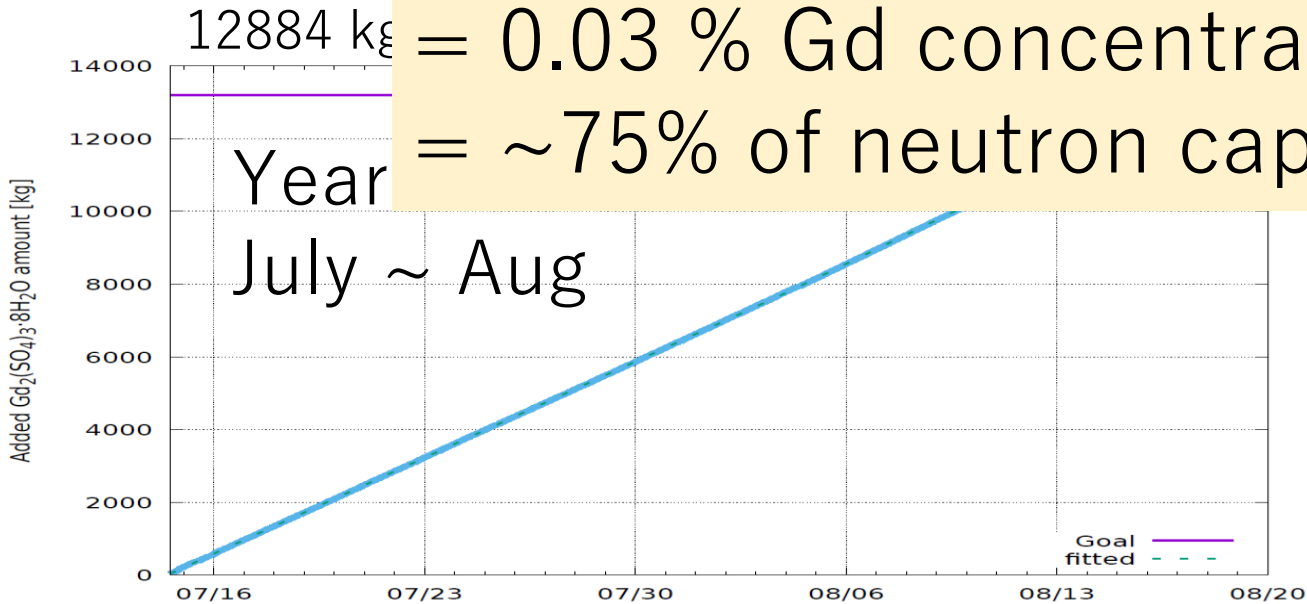
Gd loading in 2020 and 2022

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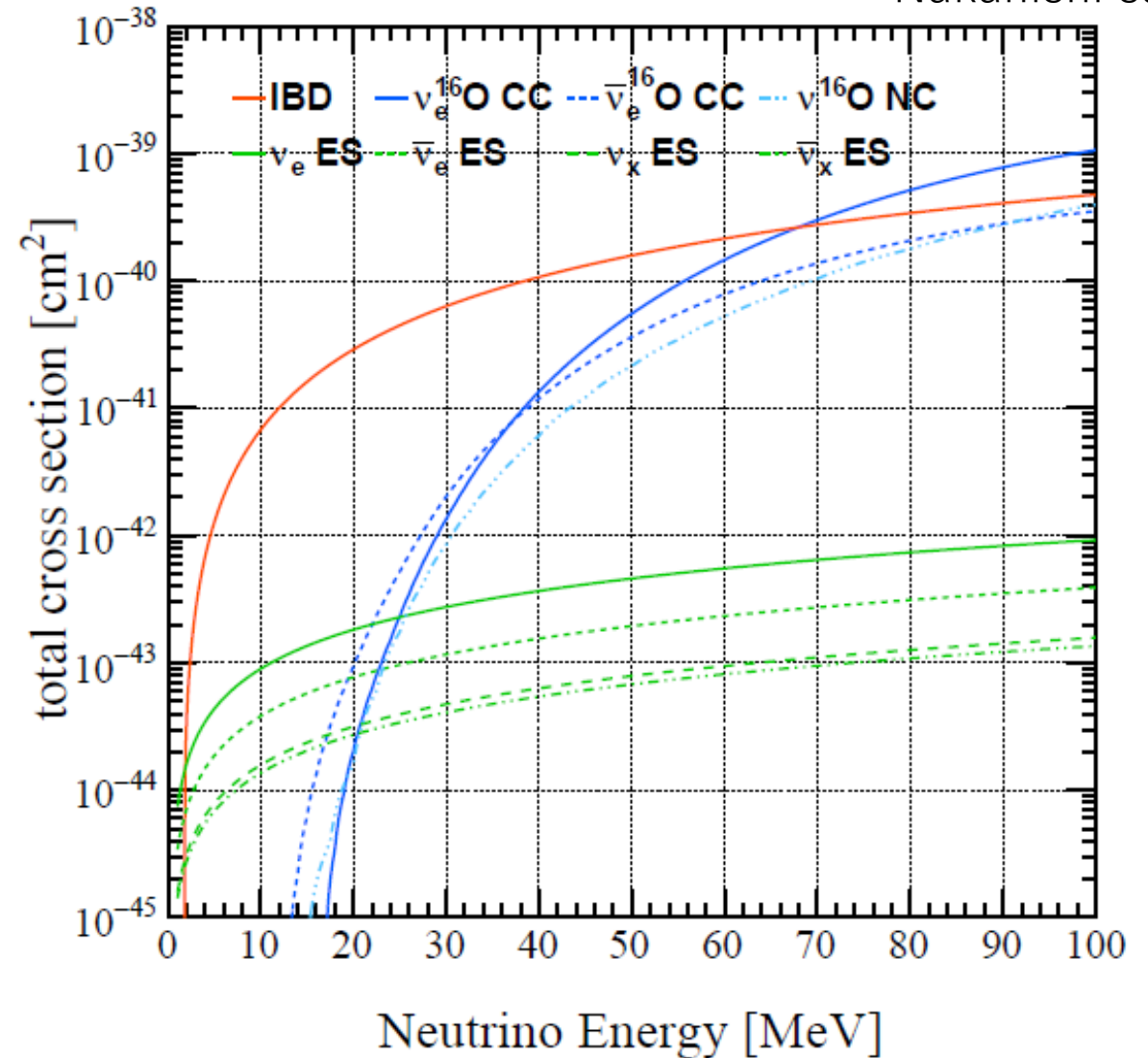
In total : ~40 tons of Gd sulfate powder
 = 0.03 % Gd concentration
 = ~75% of neutron capture by Gd



Detection of SN burst neutrinos

Nakanishi san

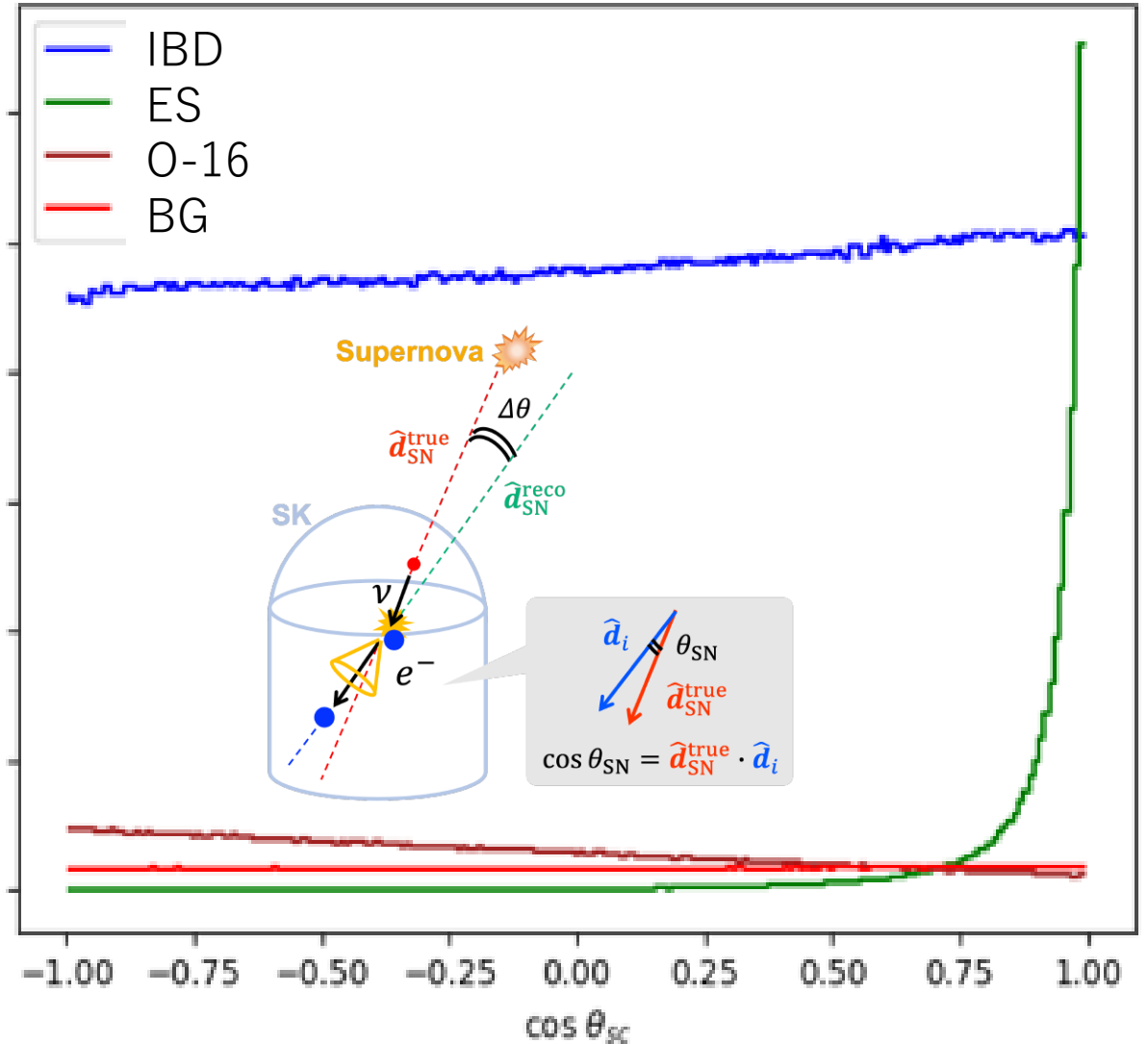
- Inverse Beta Decay (IBD)
 - $\sim 90\%$ of the expected interactions
- Electron Scattering (ES)
 - $\sim 5\%$ of the expected interactions
 - Keep the neutrino direction information
- ^{16}O interactions (CC and NC)
 - $\sim 5\%$ of the expected interactions



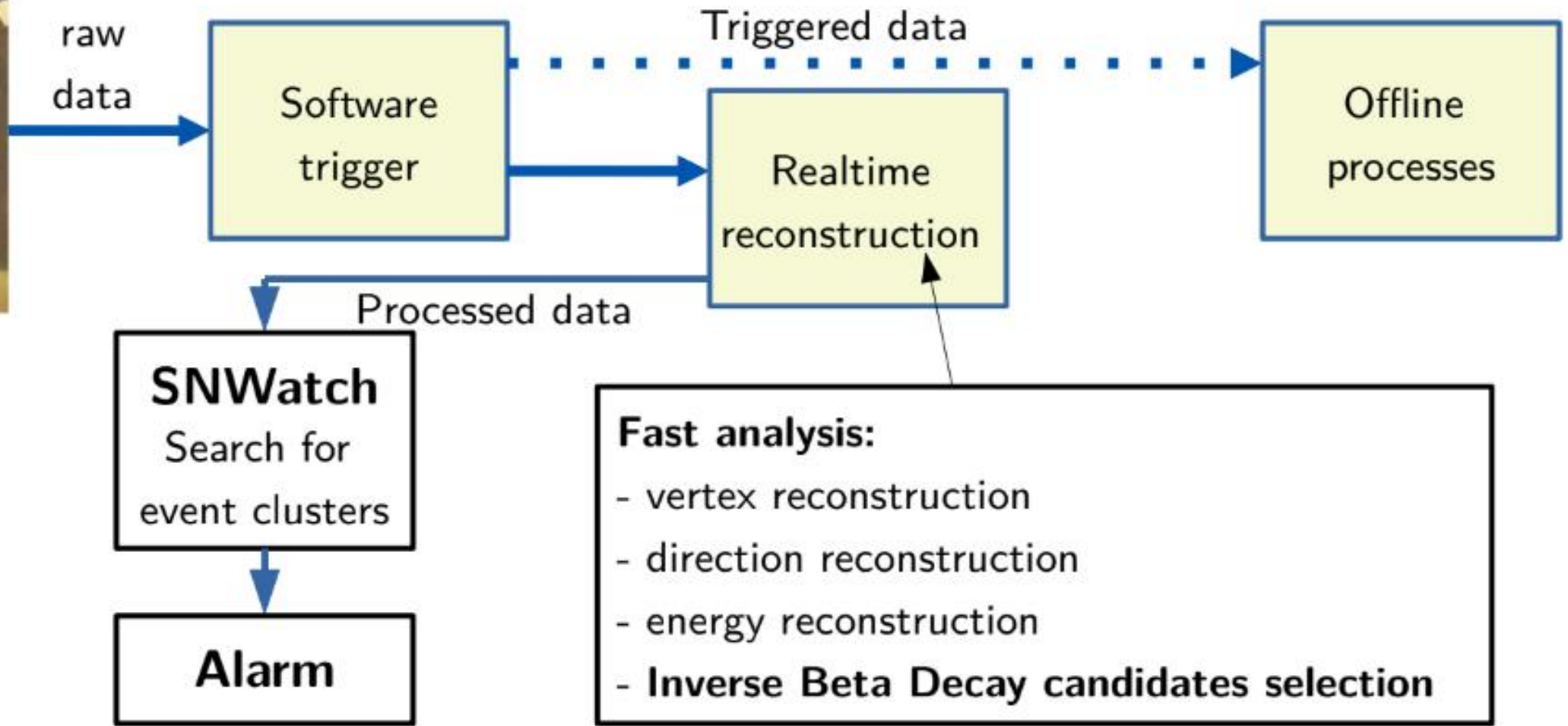
Detection of SN burst neutrinos

B.Pointon

- Inverse Beta Decay (IBD)
 - $\sim 90\%$ of the expected interactions
- Electron Scattering (ES)
 - $\sim 5\%$ of the expected interactions
- Keep the neutrino direction information
- 160 interactions (CC and NC)
 - $\sim 5\%$ of the expected interactions

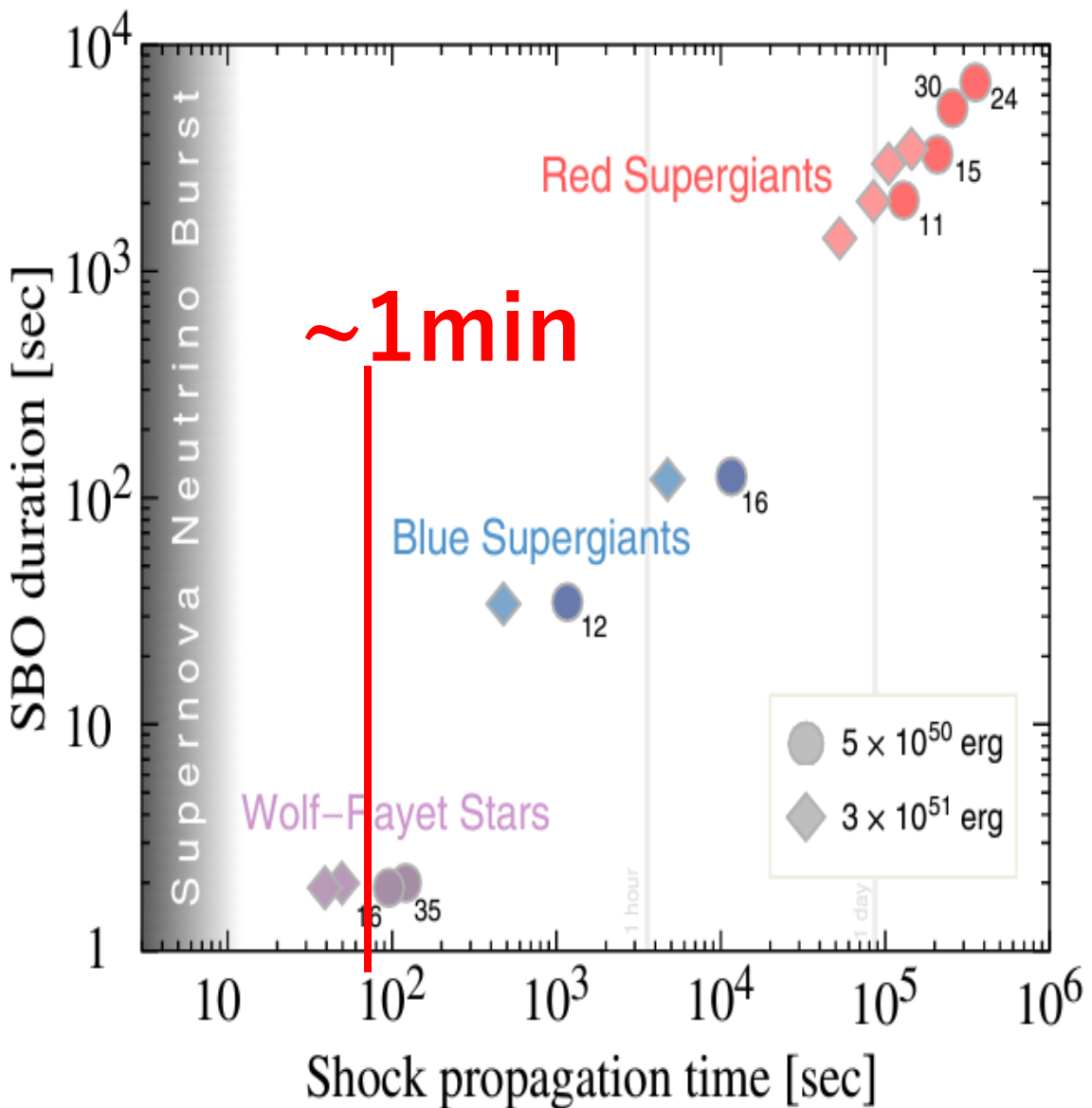


Supernova monitor at SK (Snwatch)

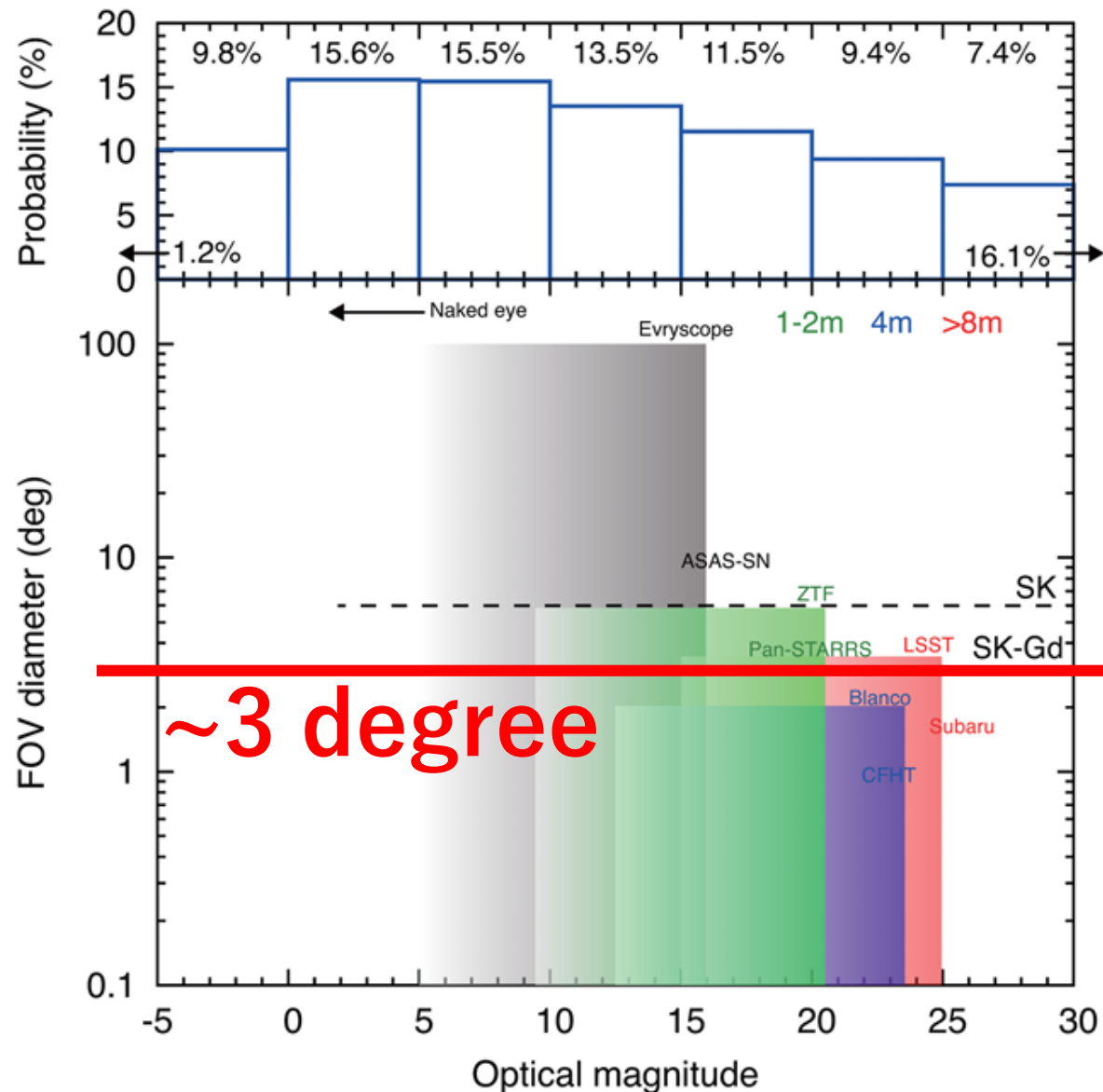


Our requirement

Kistler, M., Haxton, W., & Yuksel, H. 2013, *Astrophys. J.* 778:81, 9pp.



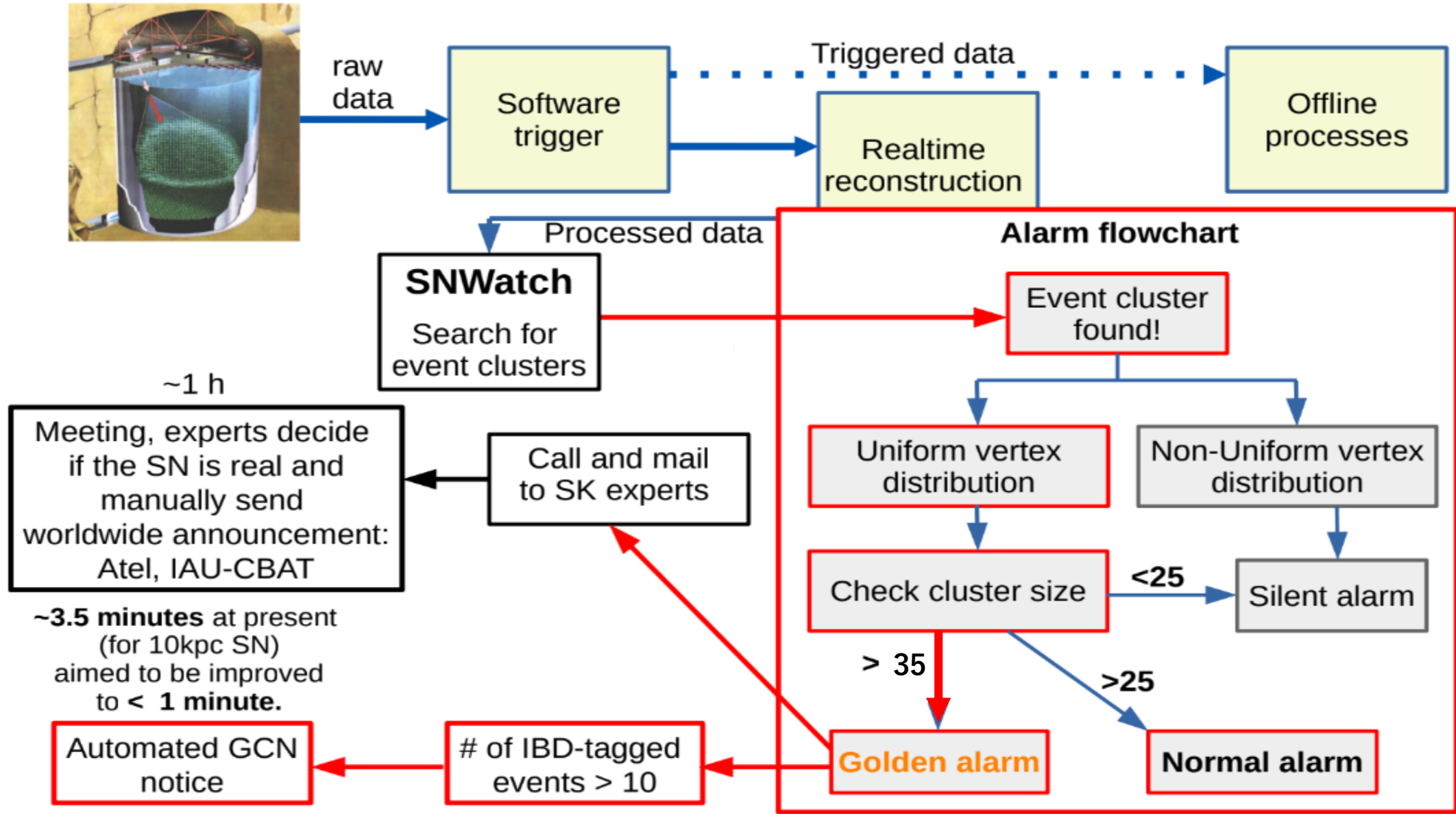
K.Nakamura et al. *MNRAS*461,3296–3313(2016)



Recent improvements of SN detection

- GCN Notice
 - G.Pronost (ILANCE/U Tokyo) et al.
- Study of SK performance for SN models
 - Kashiwagi san's (ICRR) talk at Last SN work shop
- SN direction fitter improvement
 - B.Pointon (UBTIT/TRIUMF) et al.
- Offline analysis
 - Nakanishi san (Okayama)
- Pre SN
 - SK+Kamland: Saito san (Tohoku)
- New DAQ system for very close SN
 - Mori san (NAOJ)

Automatic alert to GCN notice



Automatic GCN Notice

Alarm probability

= Can make **automatic alarm**

if we observe significant number of n-tagged IBD events.

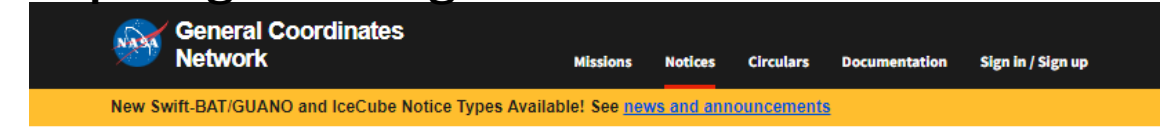
Automatic alarm:

GCN notice will be distributed automatically

~ few min after the observation.

Previously, it took ~1hour after expert checks

<https://gcn.nasa.gov/notices>



GCN Notices

GCN Notices are real-time, machine-readable alerts that are submitted by participating facilities and redistributed publicly. See the [documentation](#) for help with consuming, producing, or archiving of Notices.

Filter by tag

INTEGRAL SPI-ACS

Gamma-ray transients and light curves from the SPI-ACS instrument on INTEGRAL.

GAMMA

AGILE MCAL

GRBs detected by the MCAL instrument on AGILE.

GAMMA

AGILE SuperAGILE

GRBs detected by the SuperAGILE instrument on AGILE.

GAMMA

IPN

Light curves of GRBs detected by instruments that participate in the InterPlanetary Network (IPN).

GAMMA

Konus/WIND

GRBs detected by Konus/WIND.

GAMMA

MOA

Gravitational microlensing events detected by MOA.

OPTICAL

SNEWS

Supernova neutrinos reported by the SuperNova Early Warning System (SNEWS).

NU

Super-Kamiokande

Supernova neutrinos detected by Super-Kamiokande.

NU

GECAM

Gamma-ray transients detected by GECAM.

GAMMA

Automatic GCN Notice

Alarm probability
= Can make **automatic alarm**
if we observe significant number of n-
tagged IBD events.

Automatic alarm:
GCN notice will be distributed
automatically
~ few min after the observation.

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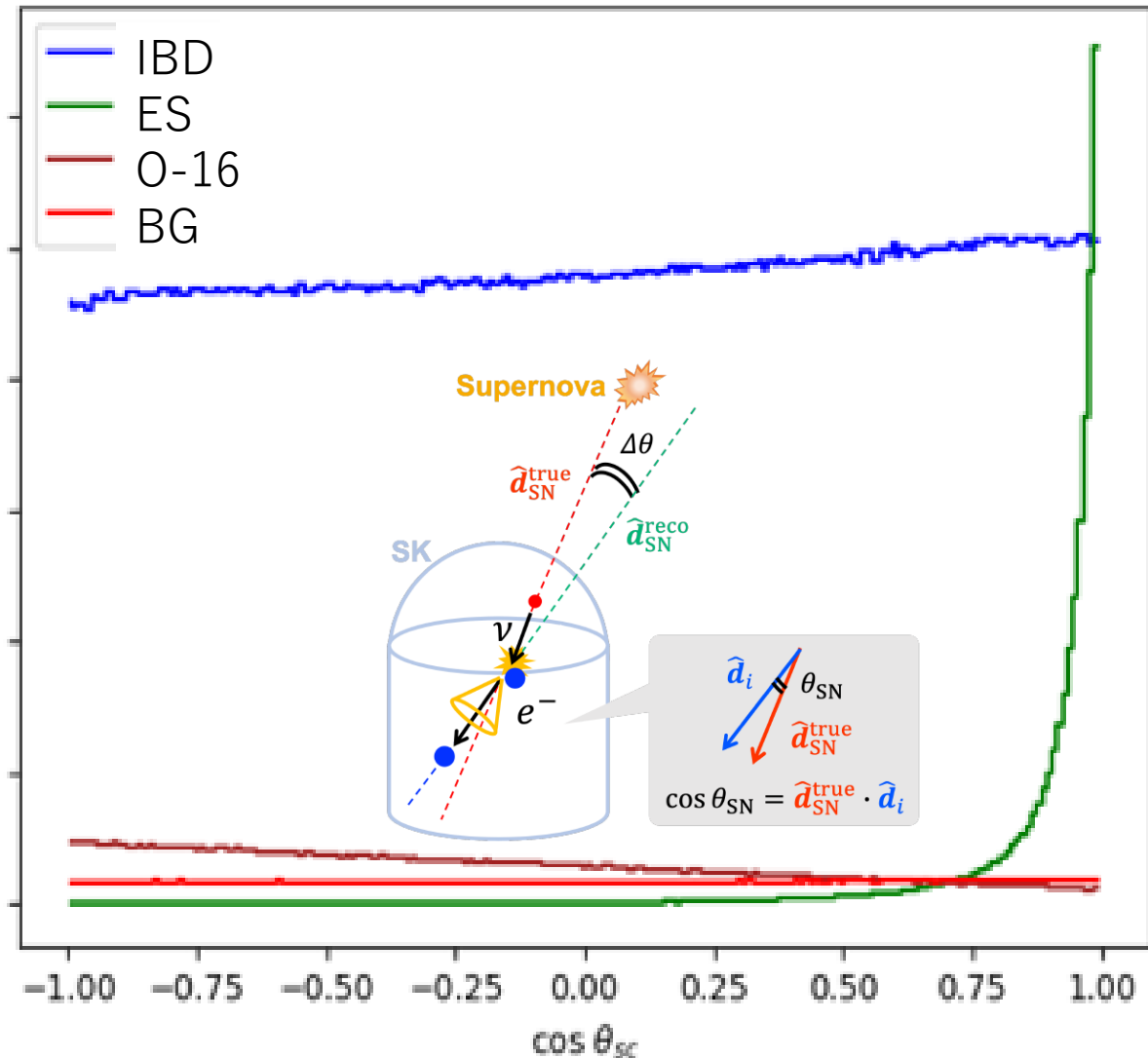
GCN Notices

```
////////////////////////////////////  
TITLE:      GCN/SK_SN NOTICE  
NOTICE_DATE:  Mon 01 Nov 21 00:00:14 UT  
NOTICE_TYPE:  SK_SN TEST  
TRIGGER_NUMBER:  SK_SN 10030  
SRC_RA:      254.4000d {+16h 57m 36s} (J2000),  
             254.6087d {+16h 58m 26s} (current),  
             253.9223d {+16h 55m 41s} (1950)  
SRC_DEC:     +31.2600d {+31d 15' 36"} (J2000),  
             +31.2275d {+31d 13' 39"} (current),  
             +31.3360d {+31d 20' 10"} (1950)  
SRC_ERROR68: 0.64 [deg radius, stat-only, 68% containment]  
SRC_ERROR90: 0.91 [deg radius, stat-only, 90% containment]  
SRC_ERROR95: 1.04 [deg radius, stat-only, 95% containment]  
DISCOVERY_DATE: 19518 TJD; 304 DOY; 21/10/31 (yy/mm/dd)  
DISCOVERY_TIME: 82816 SOD {23:00:16.74} UT  
N_EVENTS:    64124 (Number of detected neutrino events)  
ENERGY_LIMIT: 7.00 [MeV] (Minimum energy of the neutrinos)  
DURATION:    10.0 [sec] (Collection duration of the neutrinos)  
DISTANCE:    2.16 - 2.95 [kpc] (low - high as SN1987A like SNe)  
COMMENTS:    The position error is statistical only, there is no systematic added.  
COMMENTS:    All numbers are preliminary.  
COMMENTS:    NOTE: This is a TEST Notice.  
COMMENTS:
```


Direction fitter using Gd signals

B.Pointon

- Now, SK can tag IBD event with Gd
 - N-tag eff:
 - N-Capture eff \times Tagging eff $\sim 50\%$
 - Trying to improve more
- $>10\text{kpc}$, the statistics is very important.
- We should not just treat IBD events as background of ES
 - IBD also has slight directionality
- Solution:
 - If IBD like (= tagged by Gd signal)
 - Use IBD pdf (Blue)
 - If ES like
 - Set weight for IBD pdf as N-tag eff

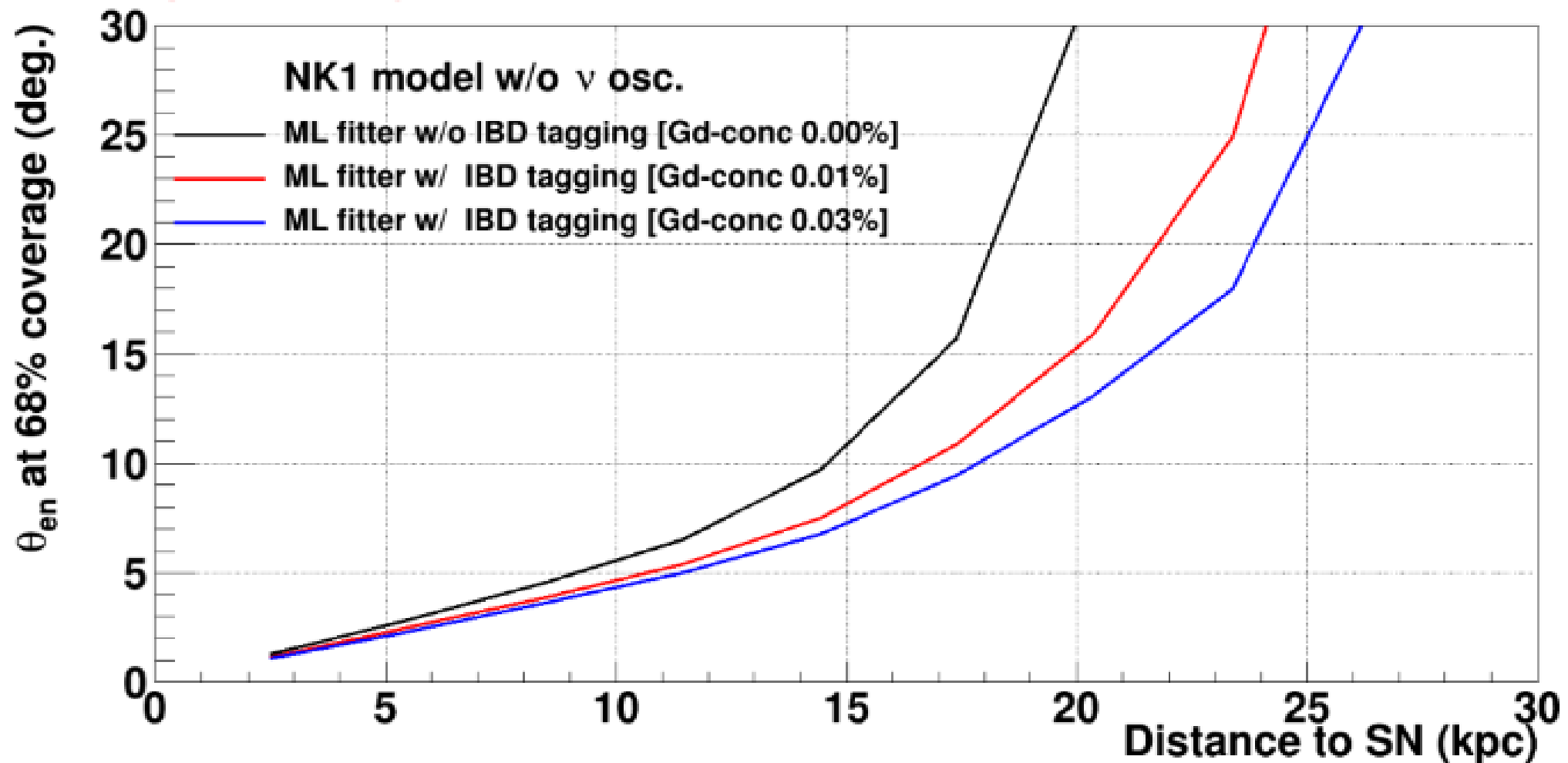


Improvement by introducing Gd

SN pointing accuracy

preliminary

G.Pronost



Performance of SK for SN models

Result
Pointing
Accuracy
at 1σ [$^\circ$]

Wilson

NMO	IMO
2.51 ± 0.08	2.81 ± 0.09

Mori

NMO	IMO
4.55 ± 0.14	4.55 ± 0.14

Fischer

NMO	IMO
6.07 ± 0.19	6.93 ± 0.22

Nakazato

NMO	IMO
4.01 ± 0.13	4.27 ± 0.14

Hüdepohl

NMO	IMO
5.11 ± 0.16	5.01 ± 0.16

Tamborra

NMO	IMO
5.09 ± 0.16	4.67 ± 0.15

- For SN at 10 kpc: 3-7 $^\circ$

- 3 $^\circ$ accuracy is achieved in Wilson model (SK-Gd's goal)

ES/IBD
Ratio

Wilson

NMO	IMO
0.035 ± 0.003	0.028 ± 0.002

Mori

NMO	IMO
0.037 ± 0.004	0.034 ± 0.004

Fischer

NMO	IMO
0.040 ± 0.006	0.033 ± 0.005

Nakazato

NMO	IMO
0.040 ± 0.004	0.033 ± 0.003

Hüdepohl

NMO	IMO
0.036 ± 0.004	0.036 ± 0.004

Tamborra

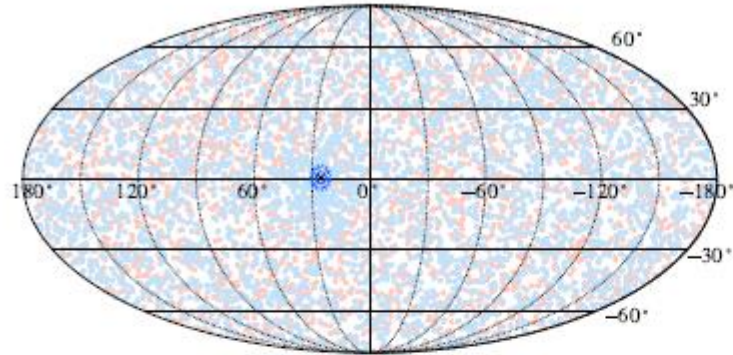
NMO	IMO
0.028 ± 0.004	0.038 ± 0.005

- Better pointing accuracy with Higher ES/IBD ratio.

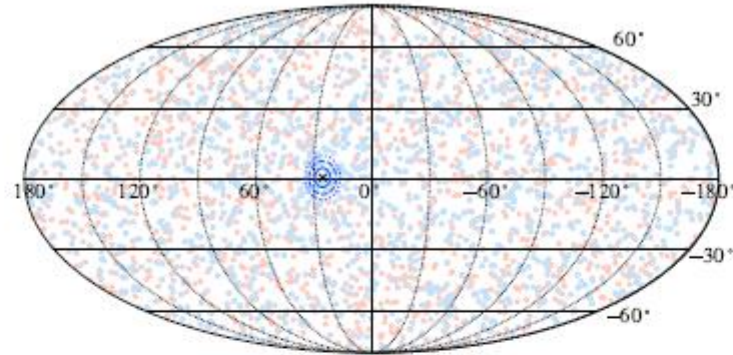
Pointing accuracy for different models

Blue : ES like, Red: IBD like

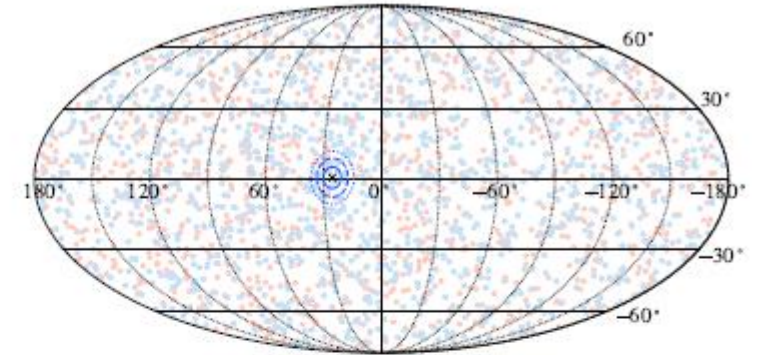
Y.Kashiwagi



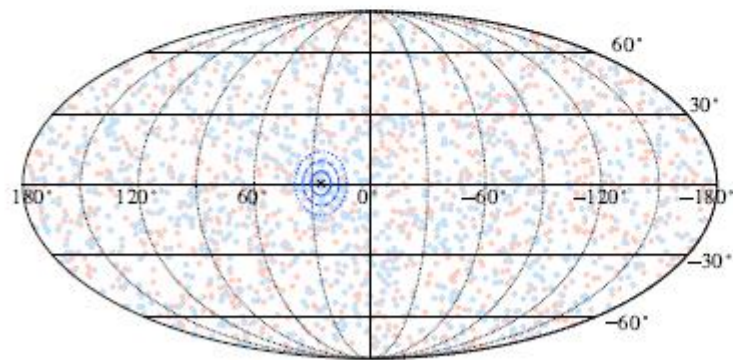
(a) the Wilson model



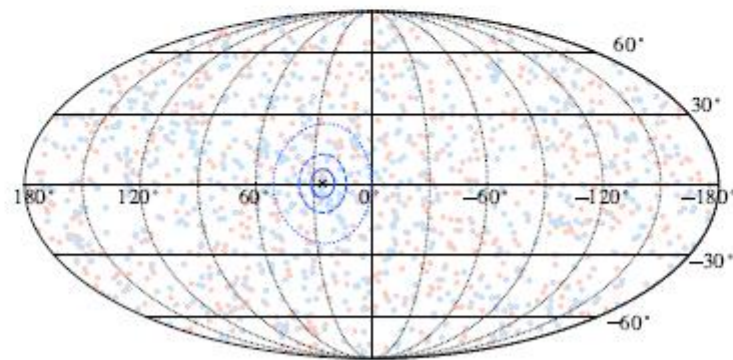
(b) the Nakazato model



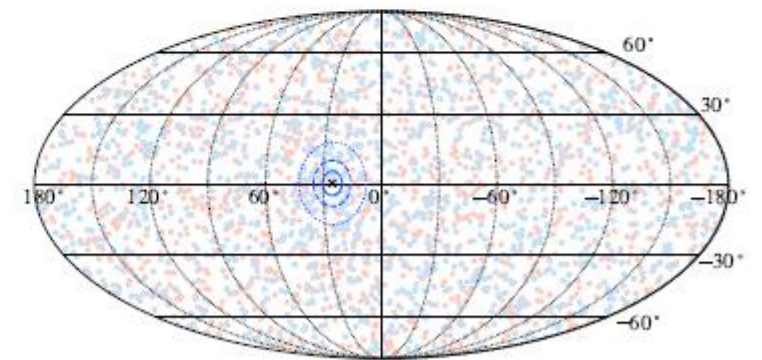
(c) the Mori model



(d) the Hüdepohl model



(e) the Fischer model

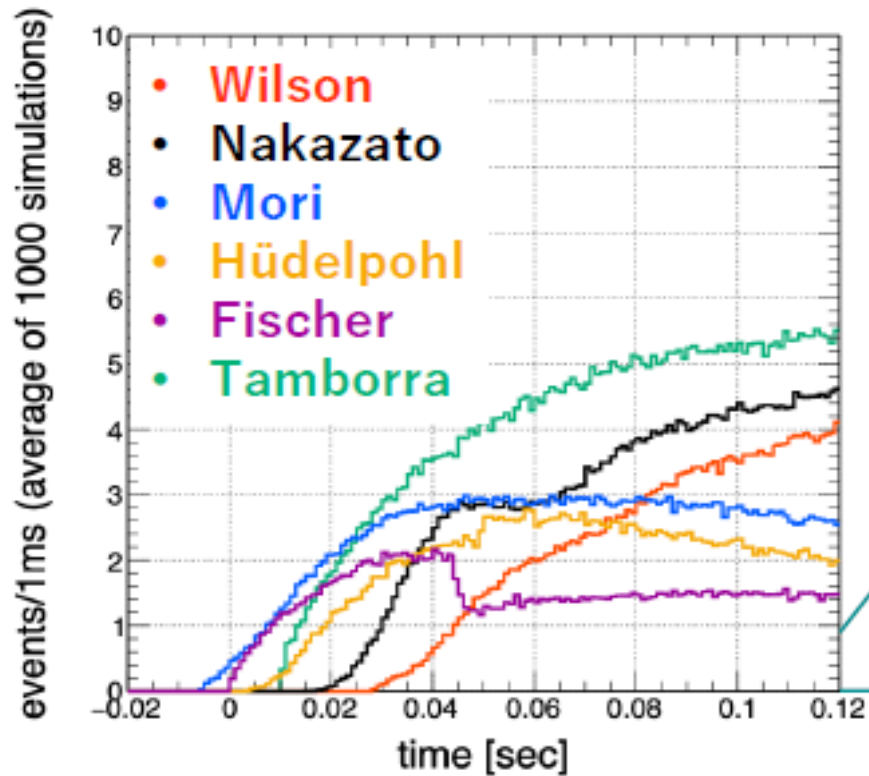


(f) the Tamborra model

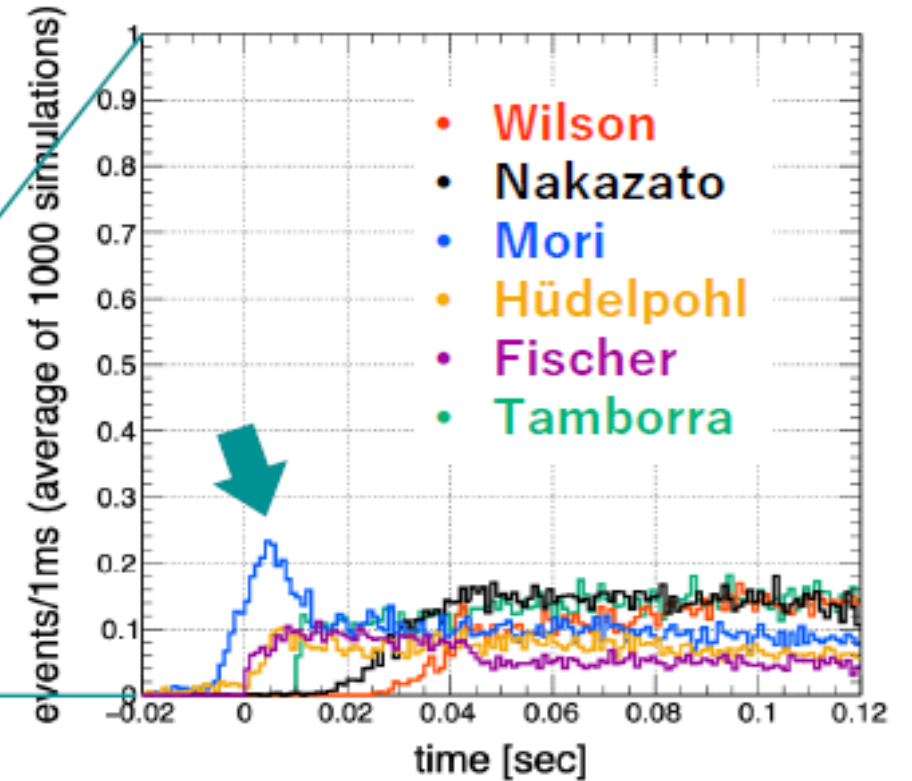
Performance of SK for SN models

- We are about submitting a new paper about SK performance to various SN models

IBD



ES

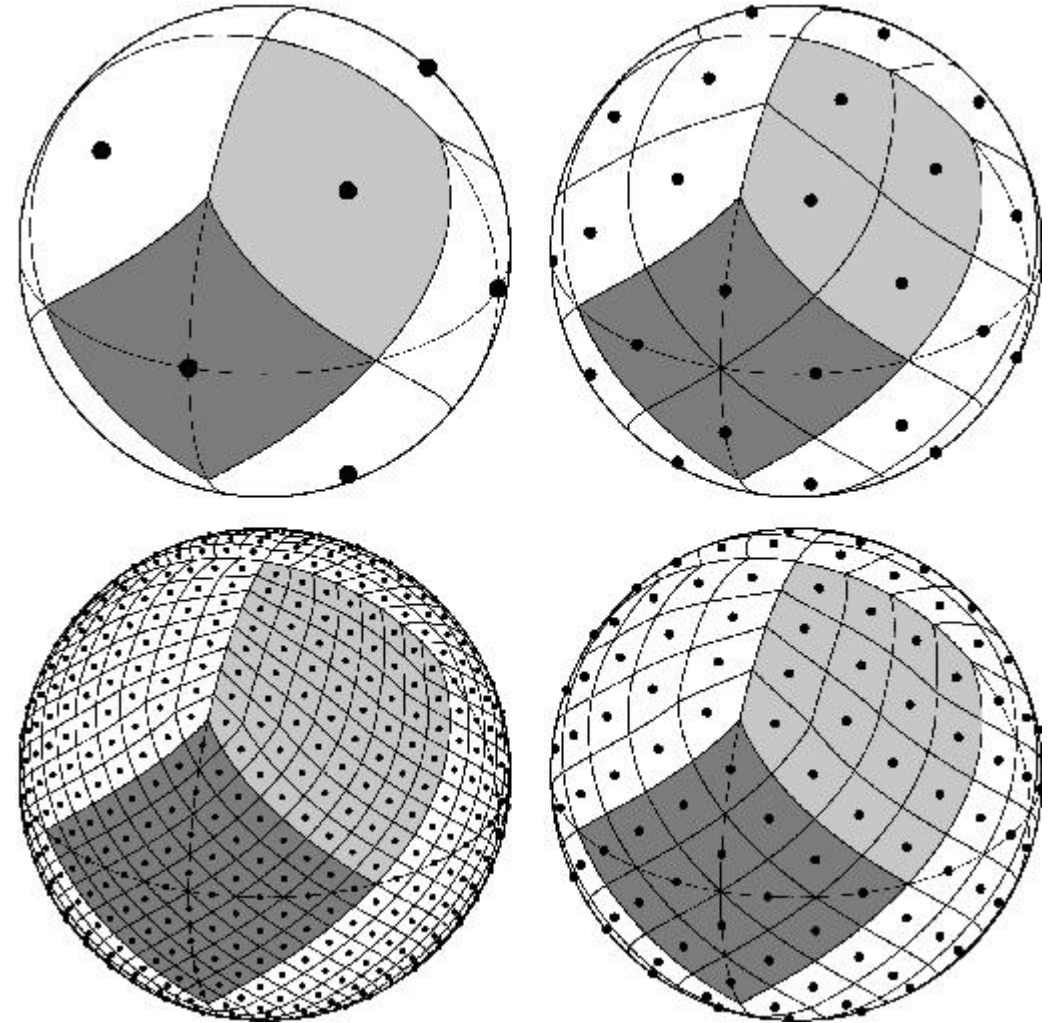


Difference in time structure can be used in model discrimination

- Neutronization burst feature appeared as an ES event peak in Mori Model
 - possibly detectable for a closer SN

Faster and more accurate!

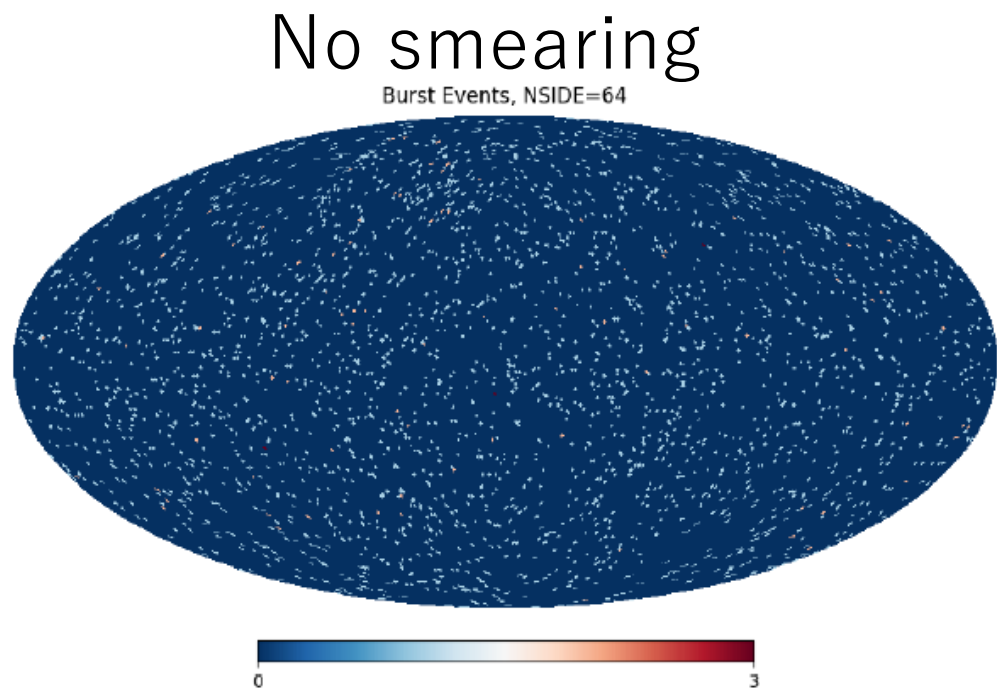
- Original fitter has 2 steps;
 - Initial grid search
 - Maximum Likelihood fit
- In both steps, we needed many loops which runs all burst events to get difference between a trial SN direction and each event direction.
 - Takes ~ 5 min for 10kpc burst
- New fitter
 - Grid search \rightarrow HEALPix spheres
 - Event loops \rightarrow put them in to vectors
 - To implement them, Python is used since it has many useful packages



<https://healpix.sourceforge.io/>

HEALPix fitter

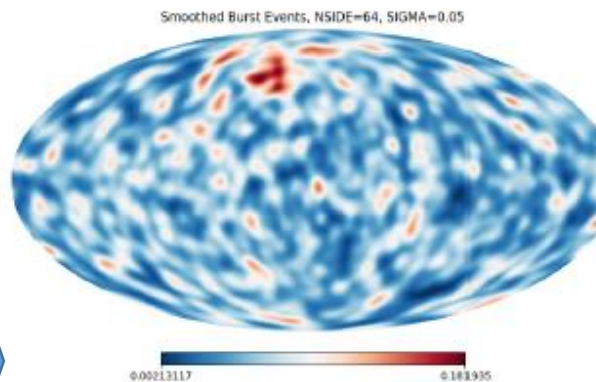
- Very simple!
 - Plot event direction to HEALPix sphere with Gaussian smearing
 - Find a pixel with maximum content



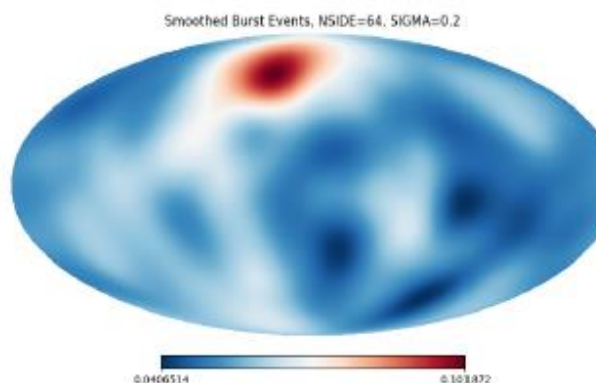
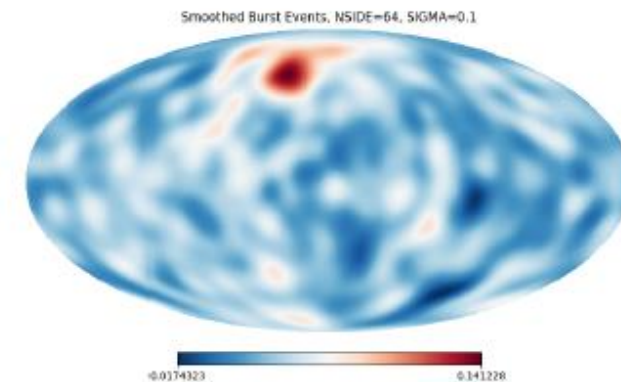
Number of pixels : 49152



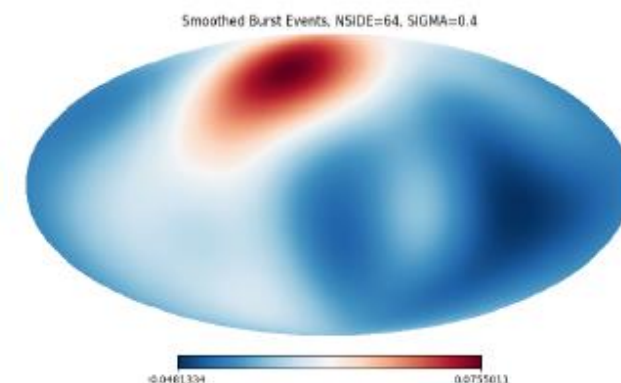
Smearing width: 0.05rad



Smearing width: 0.1rad



Smearing width: 0.2rad

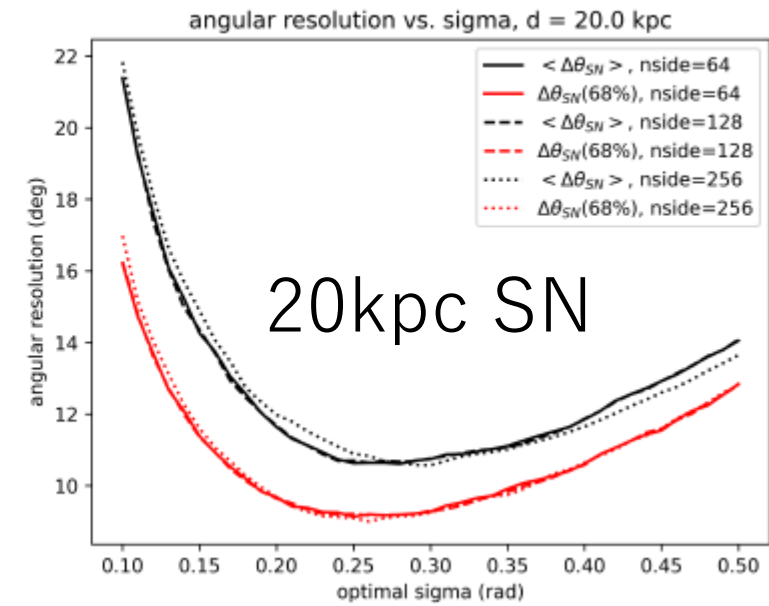
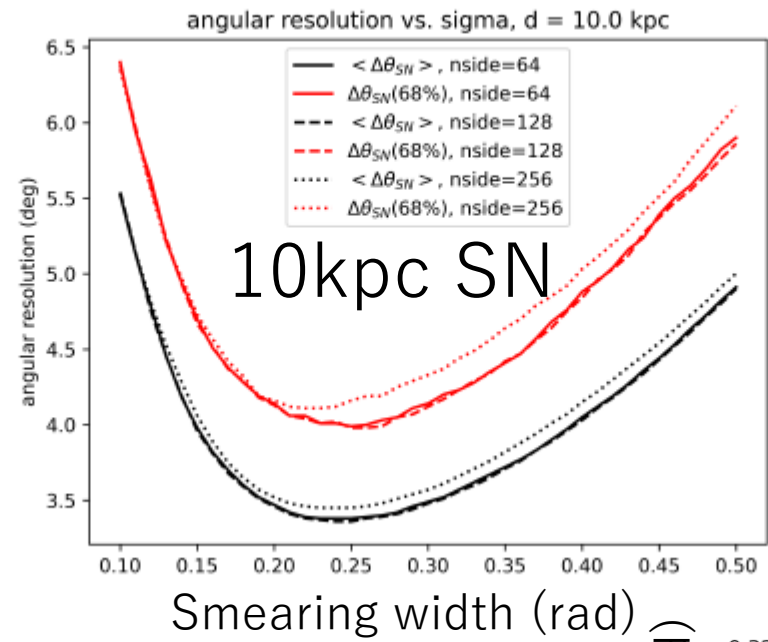
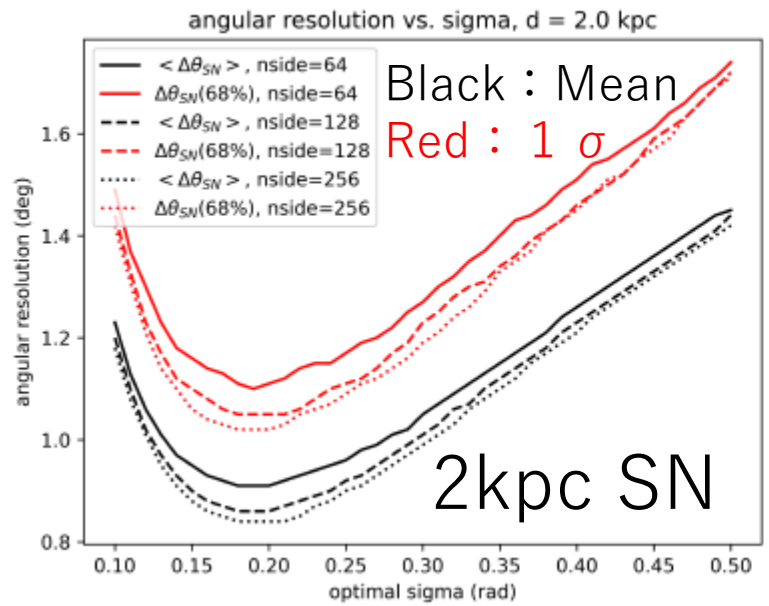


Smearing width: 0.4rad

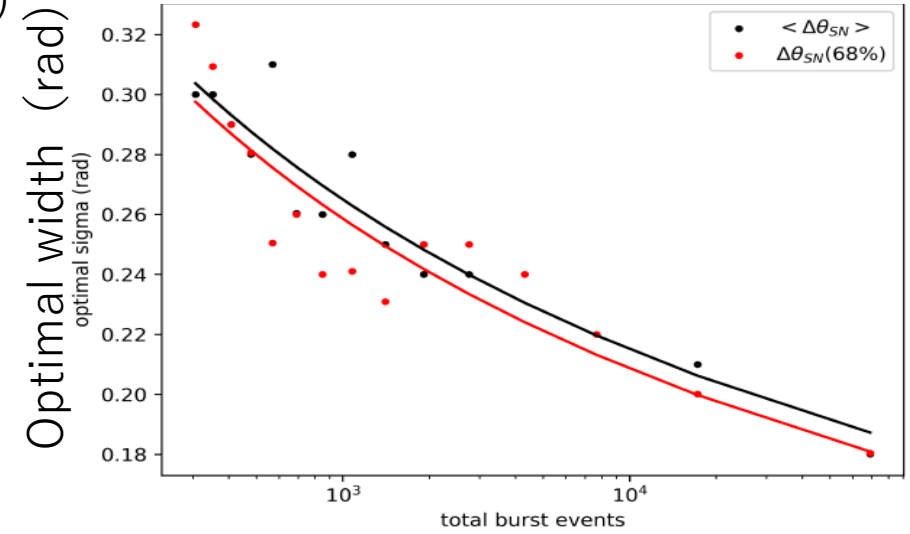
B.Pointon

Set best parameters for HP fitter

Angular resolution (deg)



NSIDE	# of Pixels	Angular Res (deg)	Pixel Area (deg ²)
8	768	7.3	53.7
16	3072	3.7	13.4
32	12,288	1.8	3.4
64	49,152	0.92	0.84
128	196,608	0.45	0.21
256	786,432	0.23	0.052



Number of events

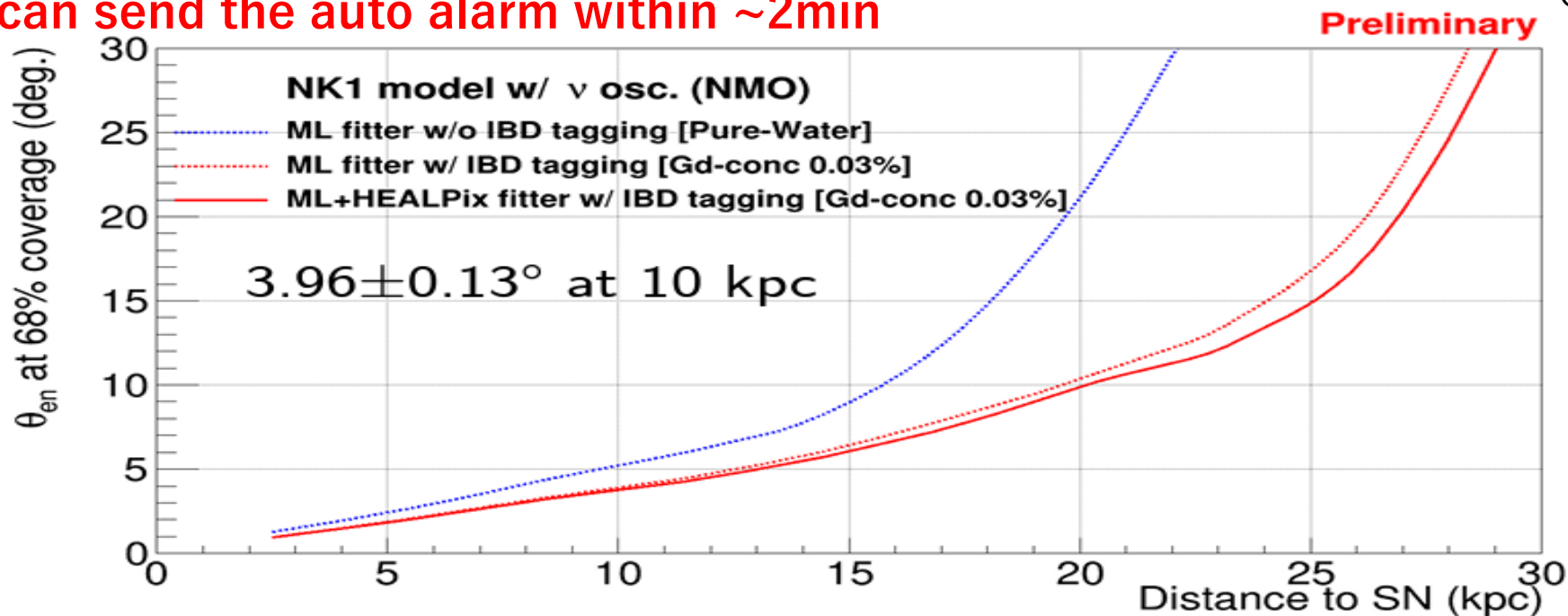
Great improve!

Paper in preparation

method	Direction Reconstruction Times (sec)		
	3000 events	10,000 events	60,000 events
<u>ML-fitter (2021)</u>	<u>~300</u>	<u>~600</u>	<u>~1320</u>
HP-fitter	0.36	0.36	0.35
<u>ML-fitter (2022)</u>	<u>1.00</u>	<u>2.21</u>	<u>11.12</u>

Including data processing and reconstruction,
we can send the auto alarm within ~2min

G.Pronost



Summary

- SK Gd status:
 - Start observation with 0.03% Gd since 2022
- Many improvements of SN burst detection
 - **Automatic GCN Notice** has been installed
 - You can register SK_SN notice
 - Study of SK performance for SN models
 - Now we have tools to compare models and data quickly!
SN direction accuracy : 3-6 degree (depending on models)
- SN direction fitter improvement
 - HP fitter and new ML fitter enable to send auto alarm within 2min.