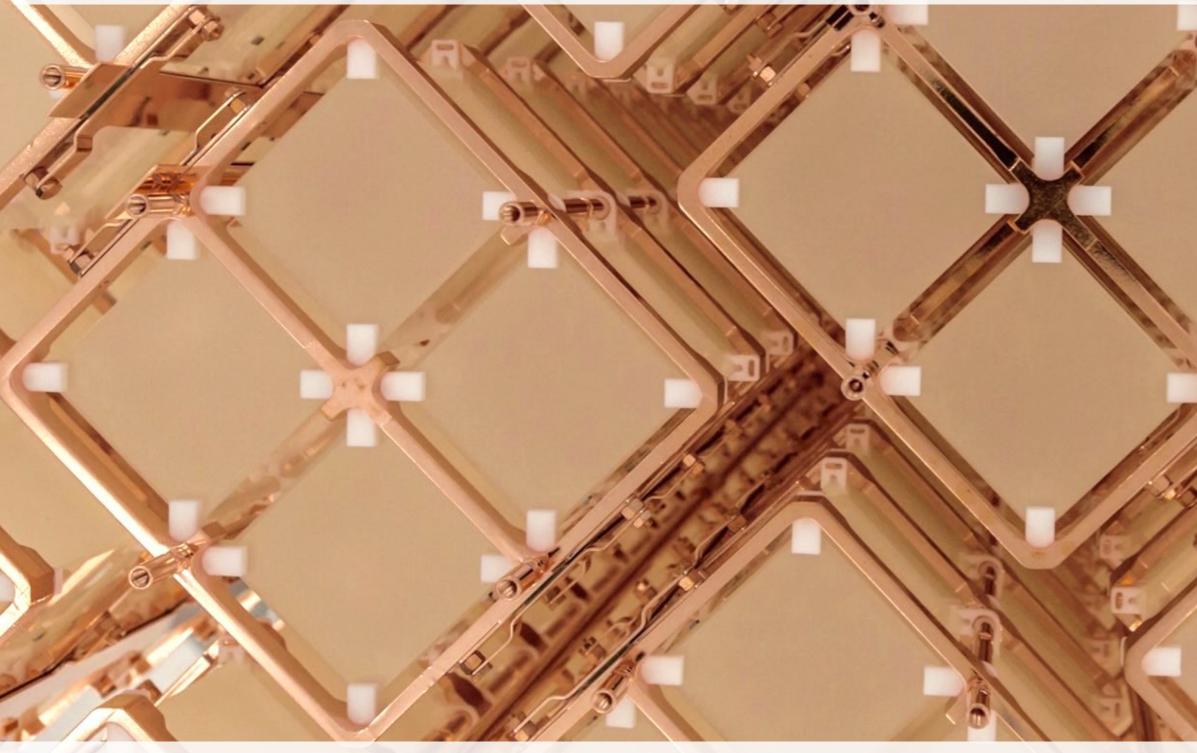
# Results from the CUORE experiment

Matteo Biassoni on behalf of the CUORE Collaboration **INFN - Sez. Milano Bicocca** 



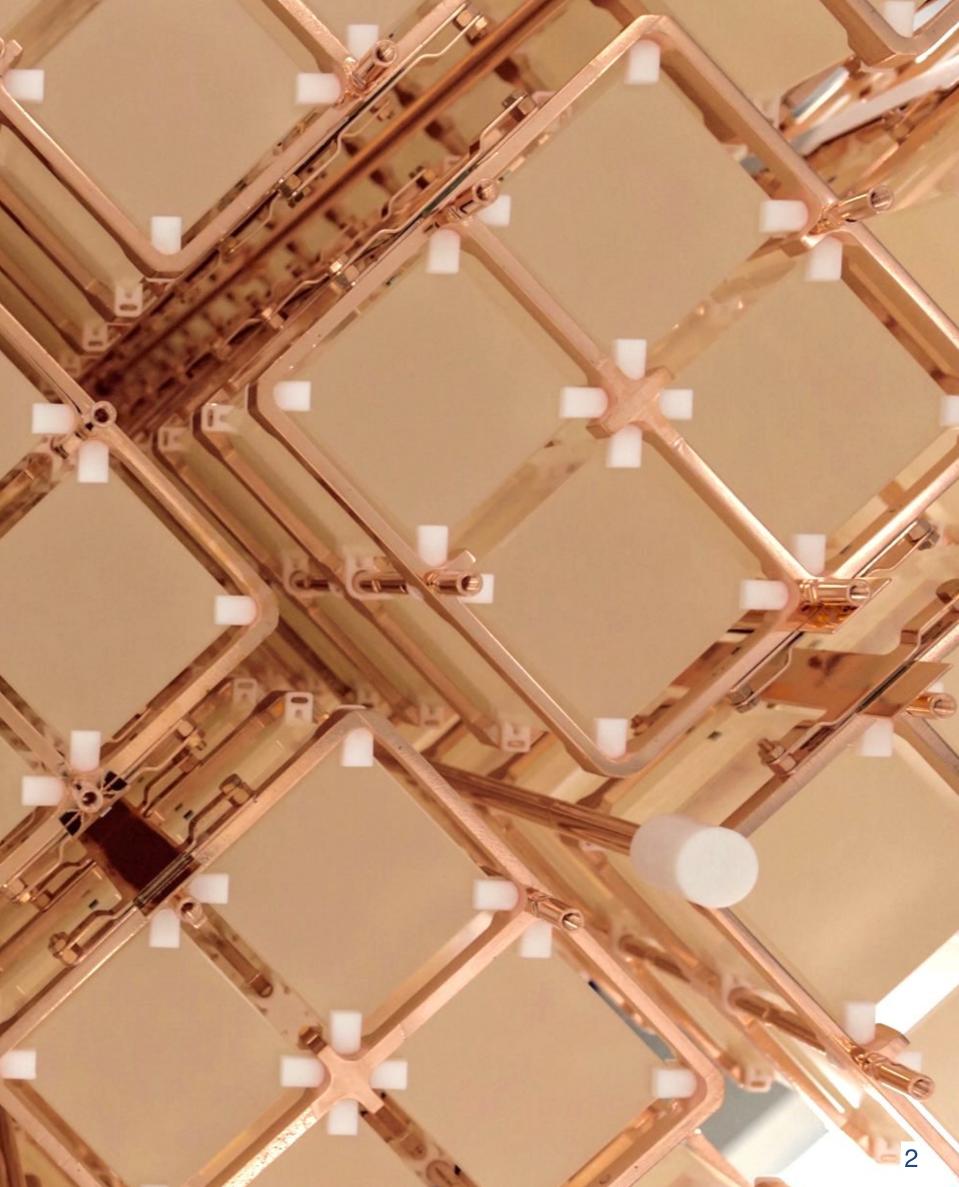
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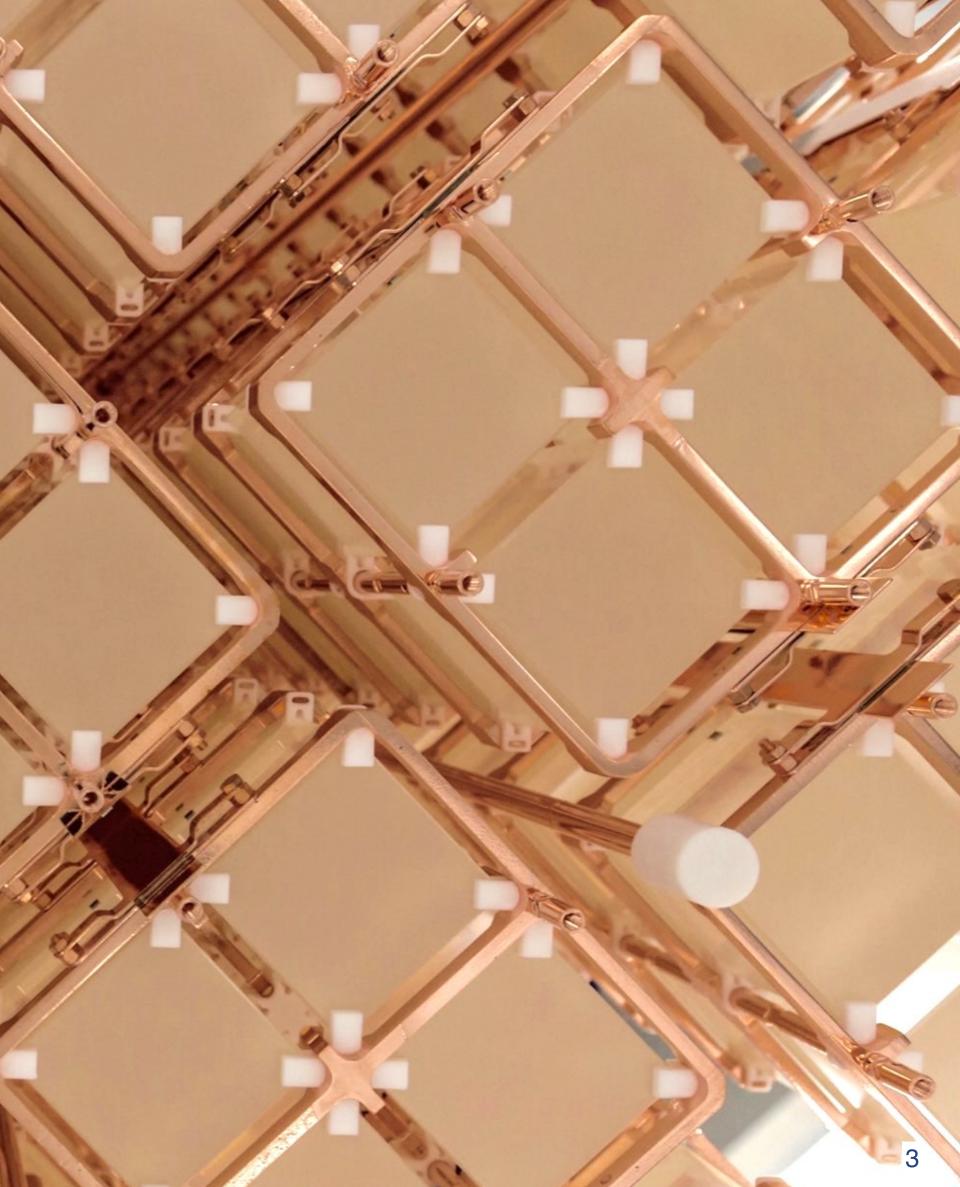


Tohoku University, Japan - March 7-9, 2019

- TeO<sub>2</sub> and thermal detectors for neutrino-less DBD
- CUORE setup
  - Location
  - Cryogenic system
  - Calibration system
  - Shielding
- CUORE roadmap
- Analysis procedures
  - Calibration and detector response
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- Physics results
  - 0vββ and lepton number violation
  - background model
  - 2vββ
  - Conclusions and outlook



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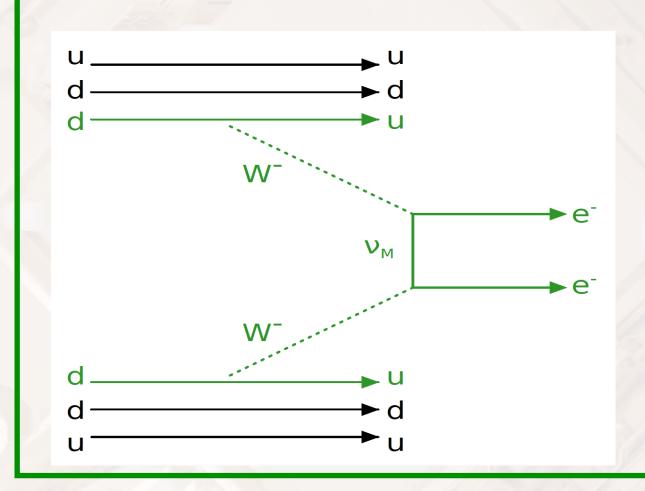


## TeO<sub>2</sub> and thermal detectors for 0vDBD

Second order nuclear process, alternative to beta decay forbidden by mass difference for some even-even nuclei

 $(A, Z) \to (A, Z+2) + 2e^- + 2\bar{\nu}_e$ 2nd order SM process,  $T_{1/2} \sim 10^{18 \sim 24}$  years

#### $(A,Z) \rightarrow (A,Z+2) + 2e^{-}$



- SM forbidden,  $\Delta L = 2$
- if observed, then neutrino is a Majorana particle
- underlying mechanism can give insight into beyond SM physics
  - light neutrino mass scale and hierarchy
  - heavy neutrino

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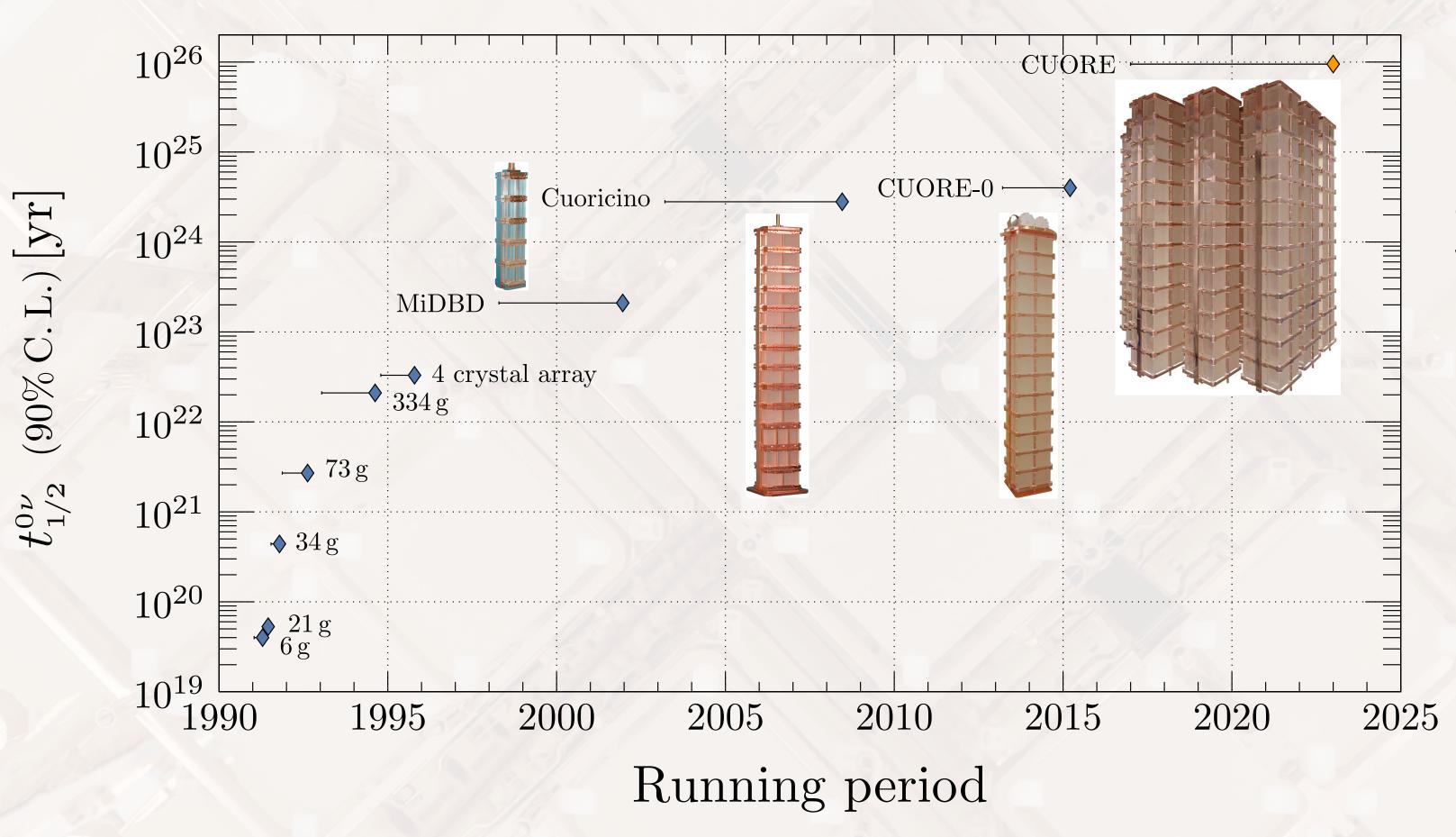
<sup>130</sup>**Te** is a good candidate source for 0vDBD search:

- high natural isotopic abundance (~34%)
- NME and phase space on average
- Q-value (2528 keV) above most of the natural radioactivity
- easy to mix in convenient chemical compounds (TeO<sub>2</sub>)

Thermal detectors are a good choice for 0vDBD search:

- excellent energy resolution
- large active mass and efficiency/unit cost
- fully active source and sensitive volume, no dead-layer

### TeO<sub>2</sub> arrays: state of the art



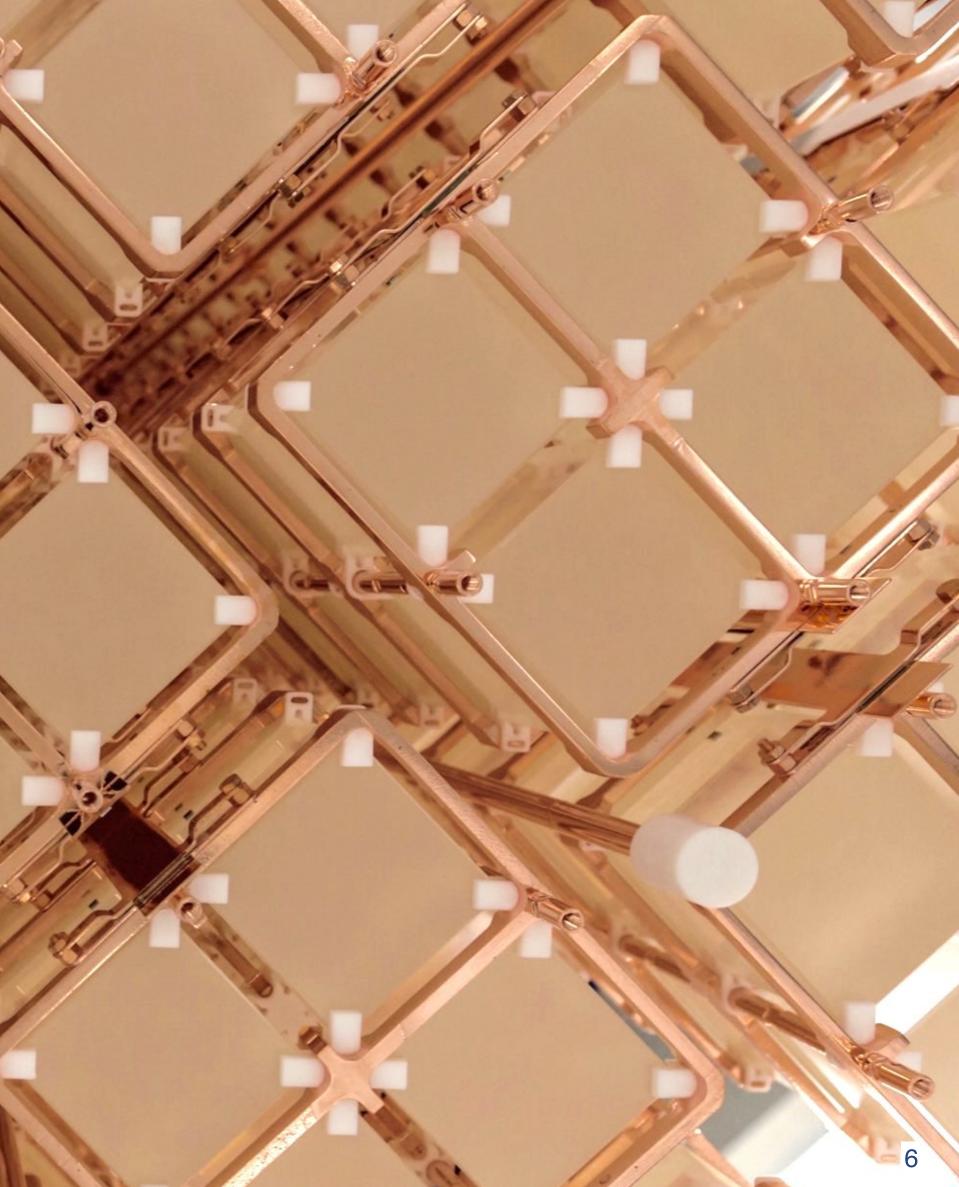
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CUORE is the latest in a long progression of TeO<sub>2</sub> detectors which included two large demonstrators:

- Cuoricino (2.8x10<sup>24</sup> y)
- CUORE-0 (4.0x10<sup>24</sup> y combined)

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## CUORE

Cryogenic Underground Observatory for Rare Events **Primary goal:** search for 0vββ decay in <sup>130</sup>Te

#### **Detector design:**

closely packed array of 988 TeO<sub>2</sub> crystals arranged in 19 towers

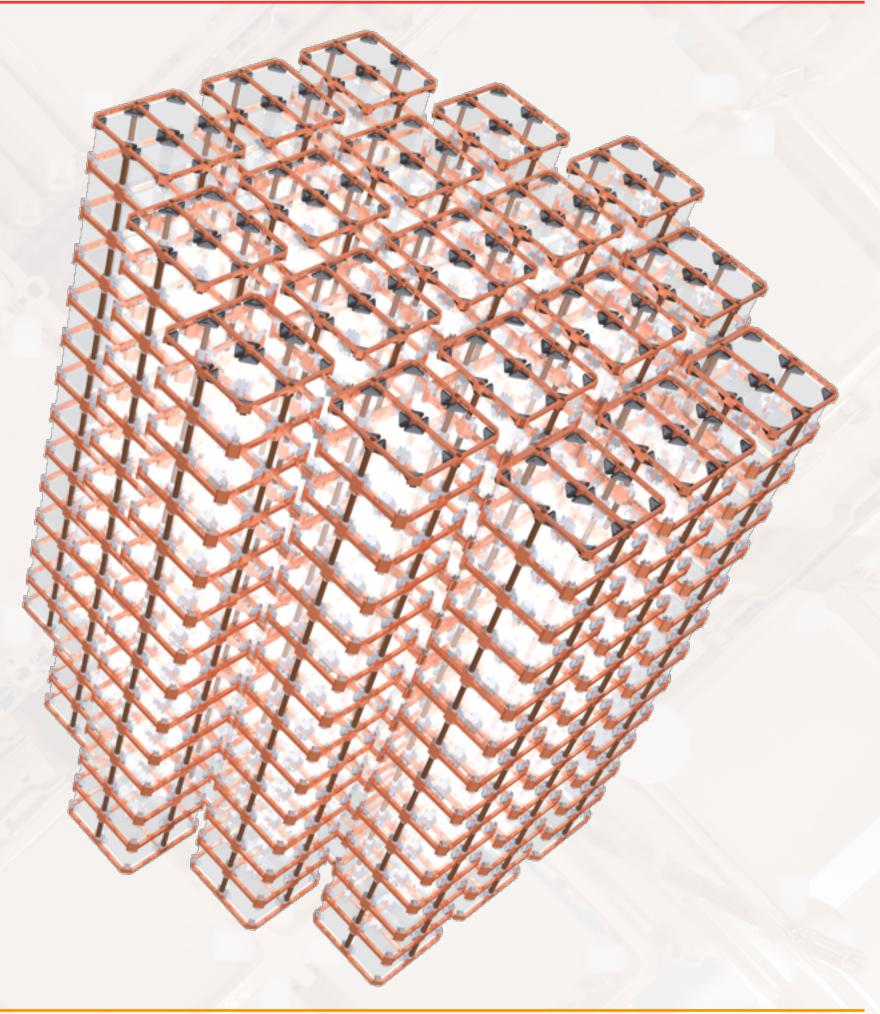
#### **Design parameters:**

- mass of TeO2: 742 kg (206 kg of <sup>130</sup>Te)
- low background aim: 10-2 c/(keV · kg · yr)
- target energy resolution: 5 keV FWHM in the Region Of Interest (ROI)
- high granularity
- deep underground location
- strict radio-purity controls on materials and assembly

#### CUORE projected sensitivity (5 years, 90% C.L.): $T_{1/2} > 9 \times 10^{25}$ yr

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# The CUORE Collaboration



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#### CUORE @ LNGS





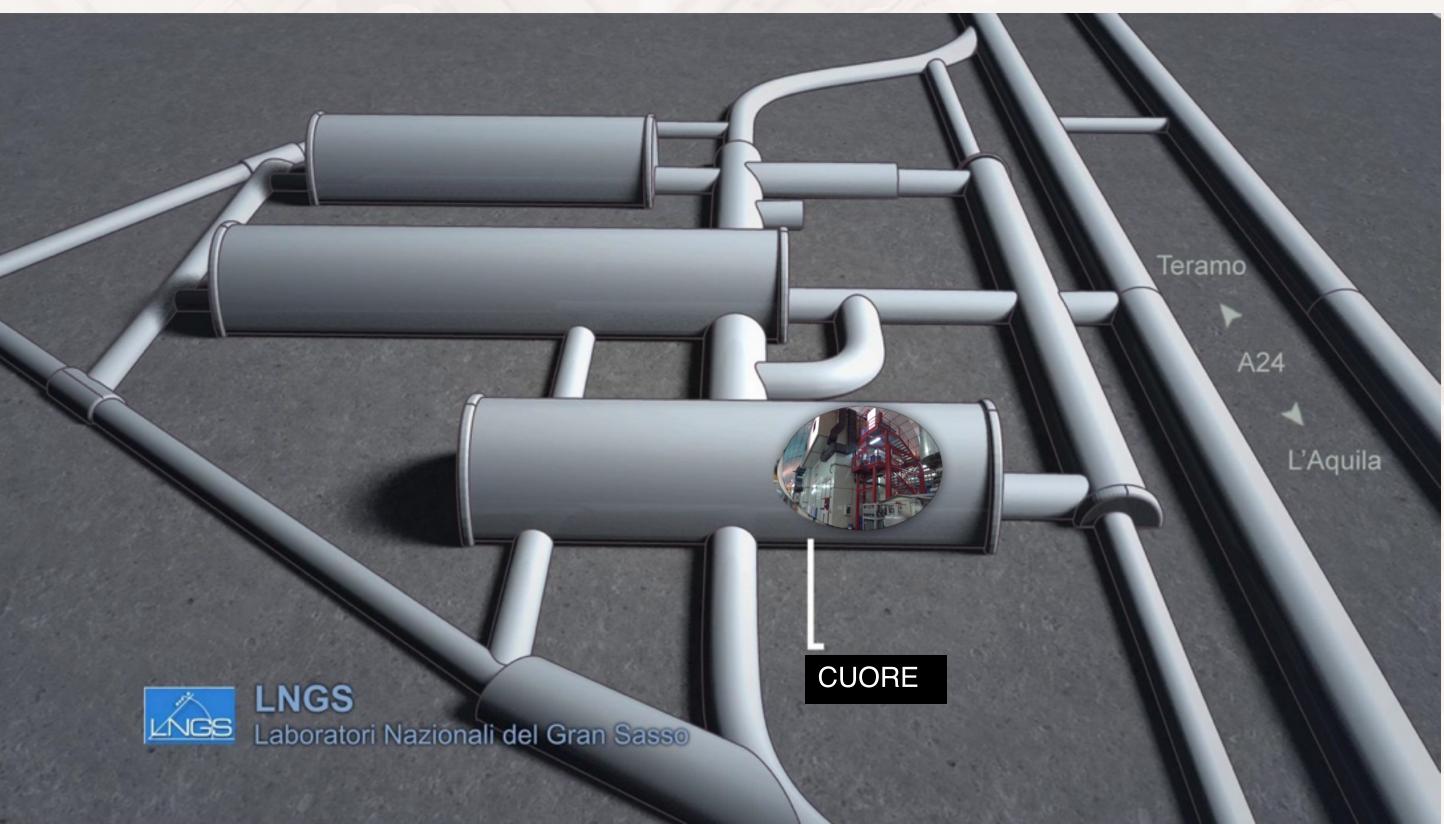
1400 m of rock (~3600 m.w.e.) deep •  $\mu$ 's:  $\sim 3 \times 10^{-8} / (s \cdot cm^2)$  γ's: ~0.73 / (s · cm<sup>2</sup>) • neutrons:  $4 \times 10^{-6} / (s \cdot cm^2)$  below 10 MeV

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### CUORE @ LNGS





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## **Underground Laboratory**

Hosting the cryostat supporting structure

Three-story building

Y beam

Main Support Plate

Cryostat

H<sub>3</sub>BO<sub>3</sub> panels

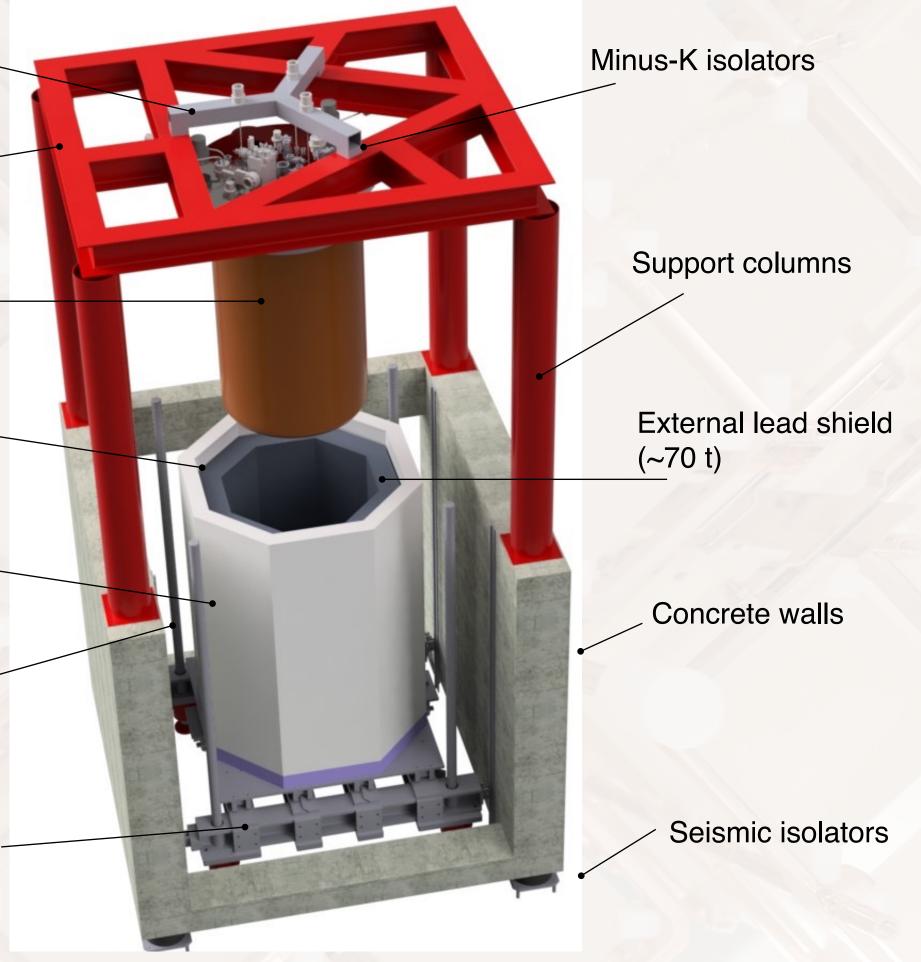
Polyethylene

Screw jacks

Movable platform

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# The CUORE cryostat

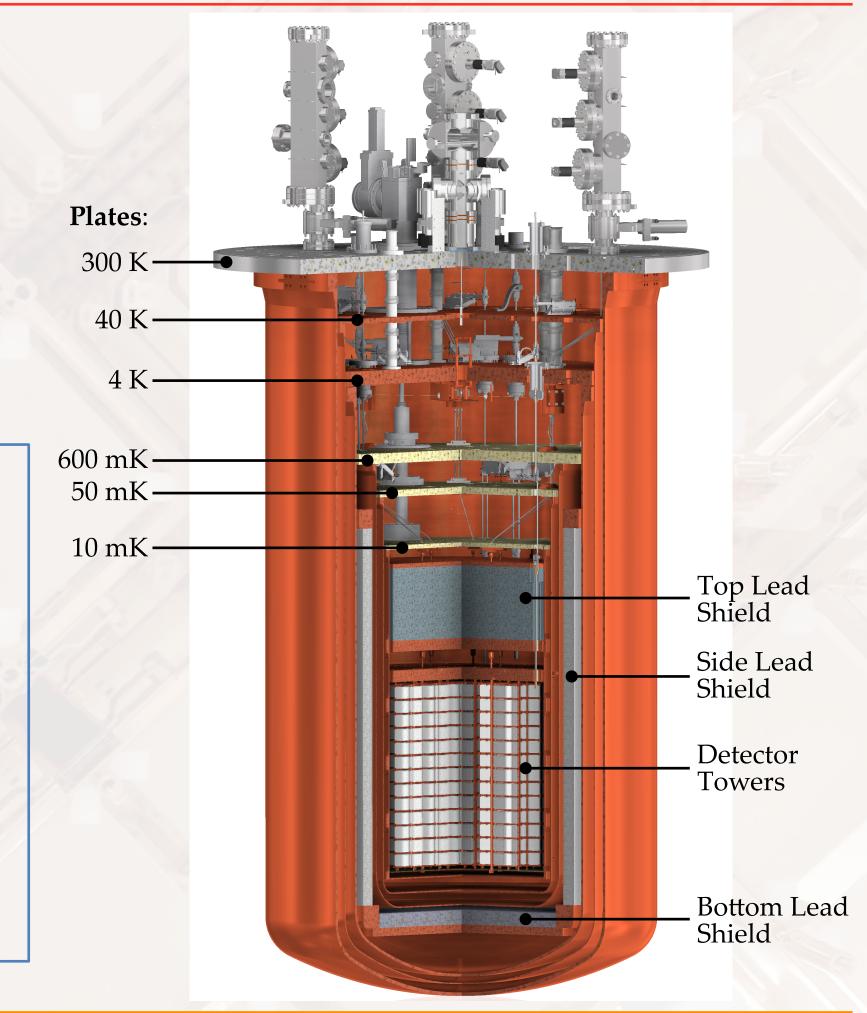
#### **Challenges:**

- Cool down ~1 ton detector to ~10 mK
- Mechanically decoupled for extremely low vibrations
- Low background environment
- Large duty cycle and long term stability

- Cryogen-free cryostat
- Fast Cooling System (<sup>4</sup>He gas) down to ~50K
- 5 pulse tubes cryocooler down to ~4K
- Dilution refrigerator down to operating temperature ~10 mK
- Nominal cooling power: 3 μW @ 10mK
- Cryostat total mass ~30 tons
- Mass to be cooled < 4K: ~15 tons</li>
- Mass to be cooled < 50 mK: ~3 tons (Pb, Cu and TeO<sub>2</sub>)

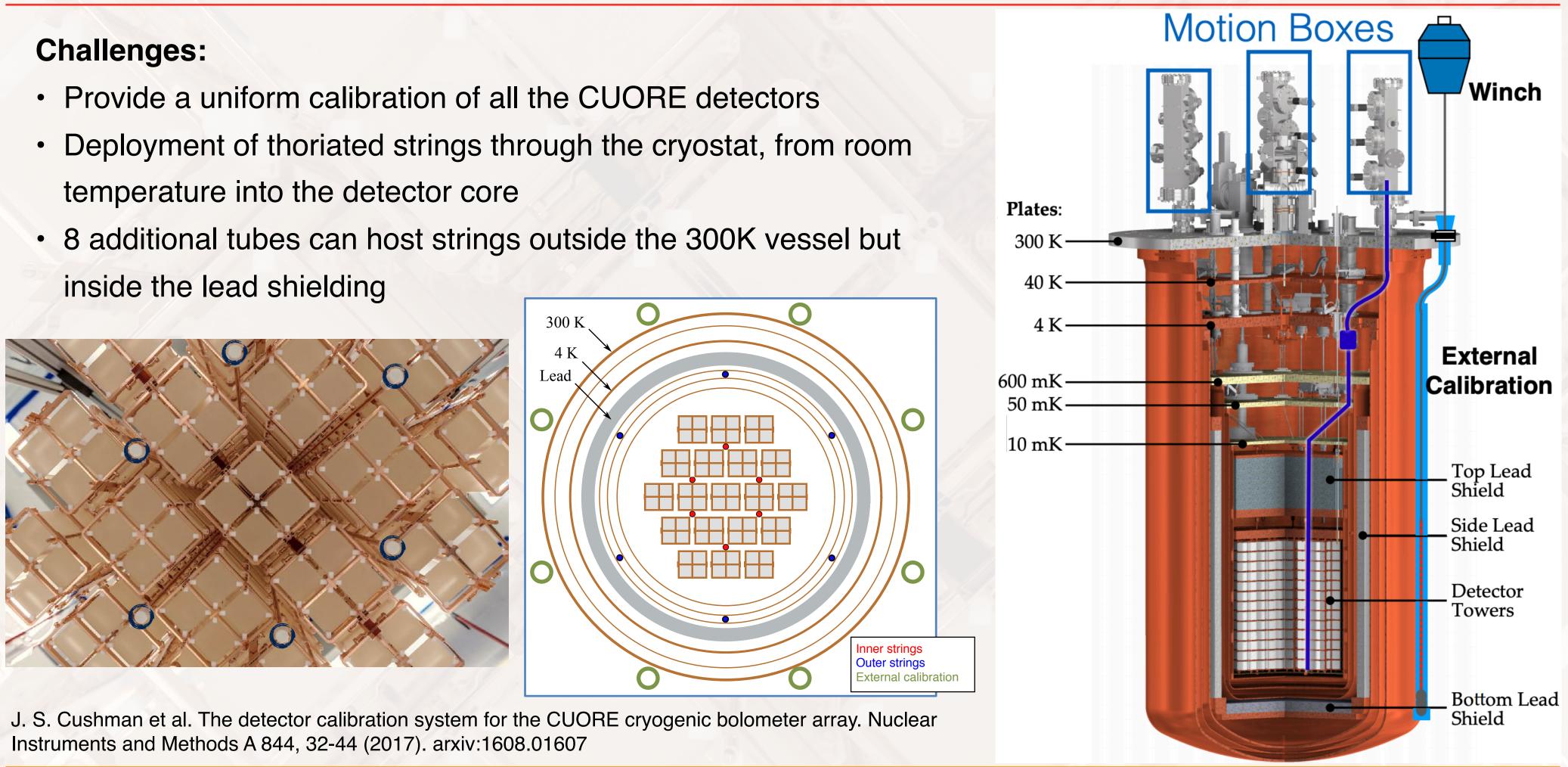
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## **Detector calibration system**

- temperature into the detector core
- inside the lead shielding



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## **Passive shielding**

#### **Challenges:**

- Protect the detectors with a heavy shield against gamma and neutron activity from external sources (~70 tonnes lead + H<sub>3</sub>BO<sub>3</sub>)
- Select materials that don't contribute themselves to the background level (ancient roman lead and selected NOSV copper)
- Cool down inner layers of the shielding to the correct temperature (2.5 tonnes @ 50mK + 5.5 tonnes @ 4K)

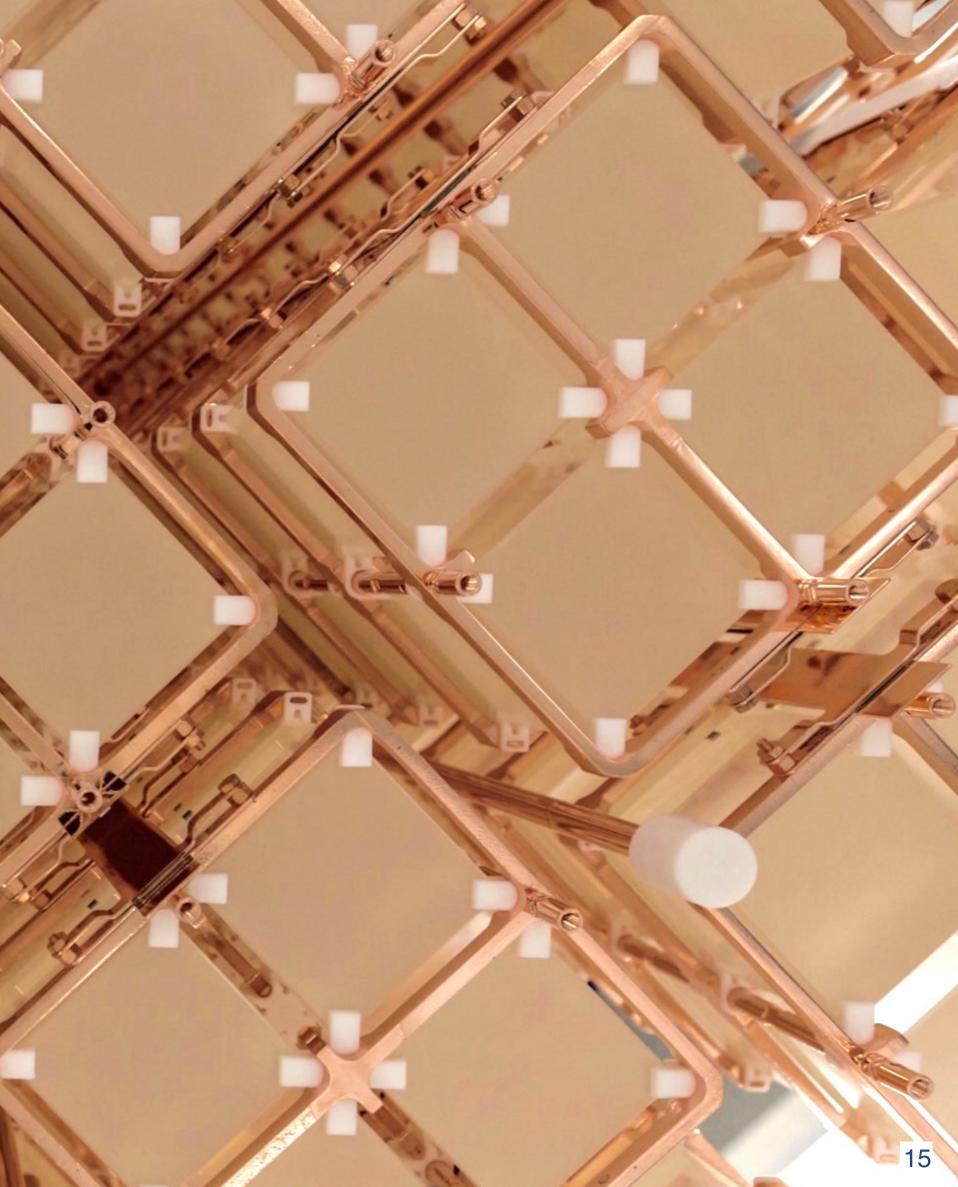


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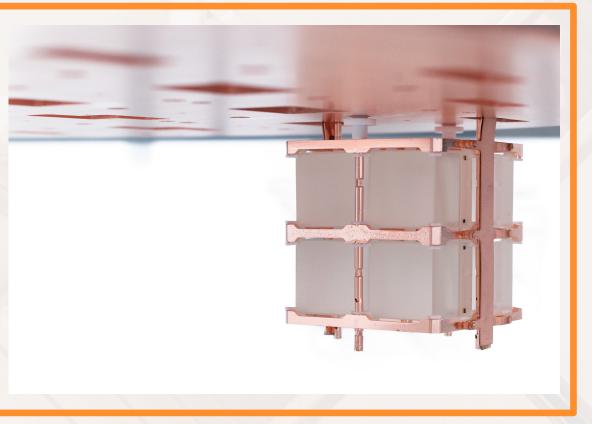


# **CUORE** Roadmap

#### Cryostat Commissioning

- 8-channels functioning mock-up detector
- Noise study and mitigation
- Stable base temperature < 7 mK
- Calibration sources deployment

Feb 2016



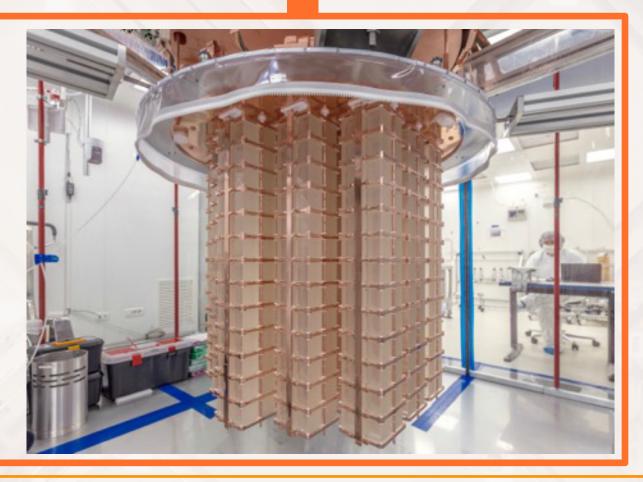
- 22 days

- down

#### Aug - Oct 2016

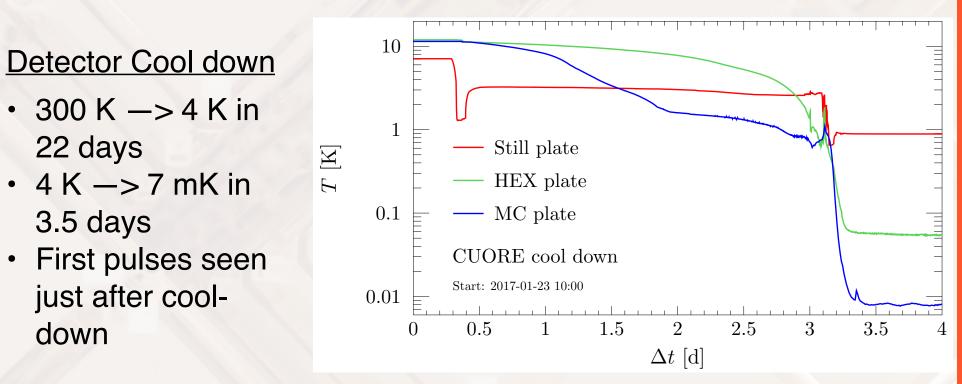
#### **Detector Installation**

- Radon-free environment
- 1 tower/day, 3 operators
- Read-out testing
- Cryostat interfaces
- Inner radiation shields



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Jan 2017

#### Mar - Apr 2017

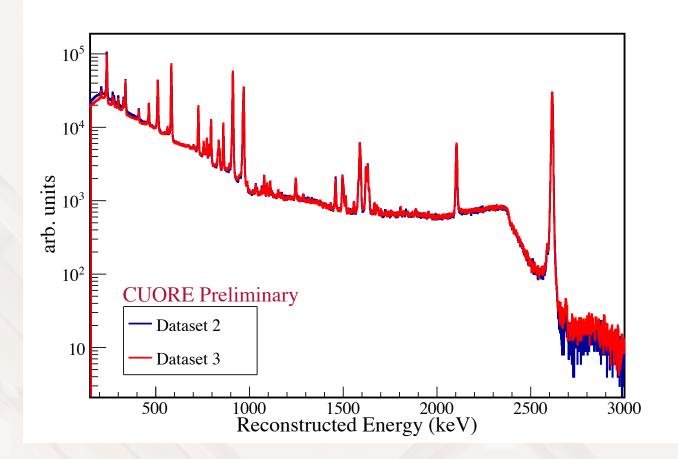
#### **Detector pre-operation**

- Optimisation of all sub-systems
- Working temperature and working point selection
- Noise reduction

# **CUORE** Roadmap

#### Second period of Science Data taking

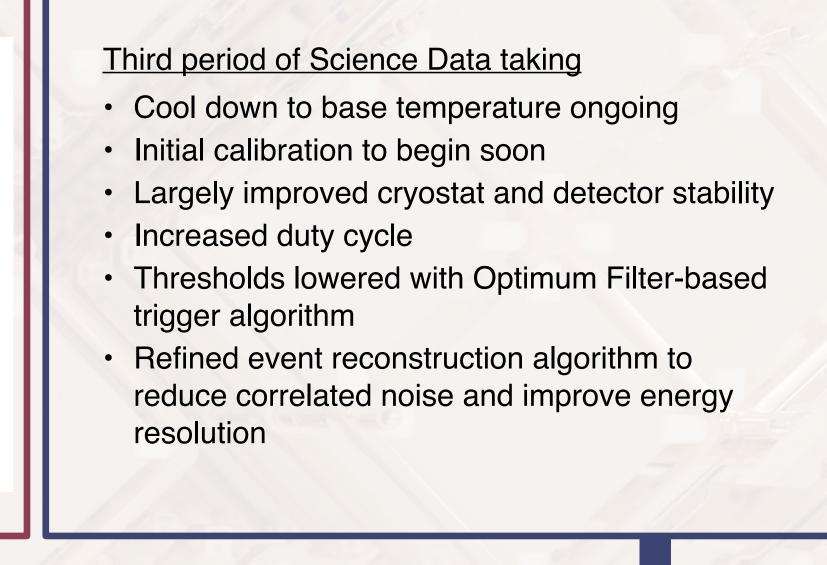
- 2 Datasets
- 984/988 operational channels
- New optimised working temperature
- New working points
- PT noise reduction
- Resolution in new data unchanged
- More then doubled the statistics
- Updated physics results coming soon



#### Jul - Fall 2018 Nov - Mar 2017 May - Sep 2017 First period of Science Data taking Partial warm up Warmed up to 100 K 2 Datasets Fix small leak in 984/988 operational channels the cryostat region 86.3 kg\*y exposure Improved mixture circuit to Warmed up to Most sensitive search for $0\nu\beta\beta$ 100 K in <sup>130</sup>Te to date term stability

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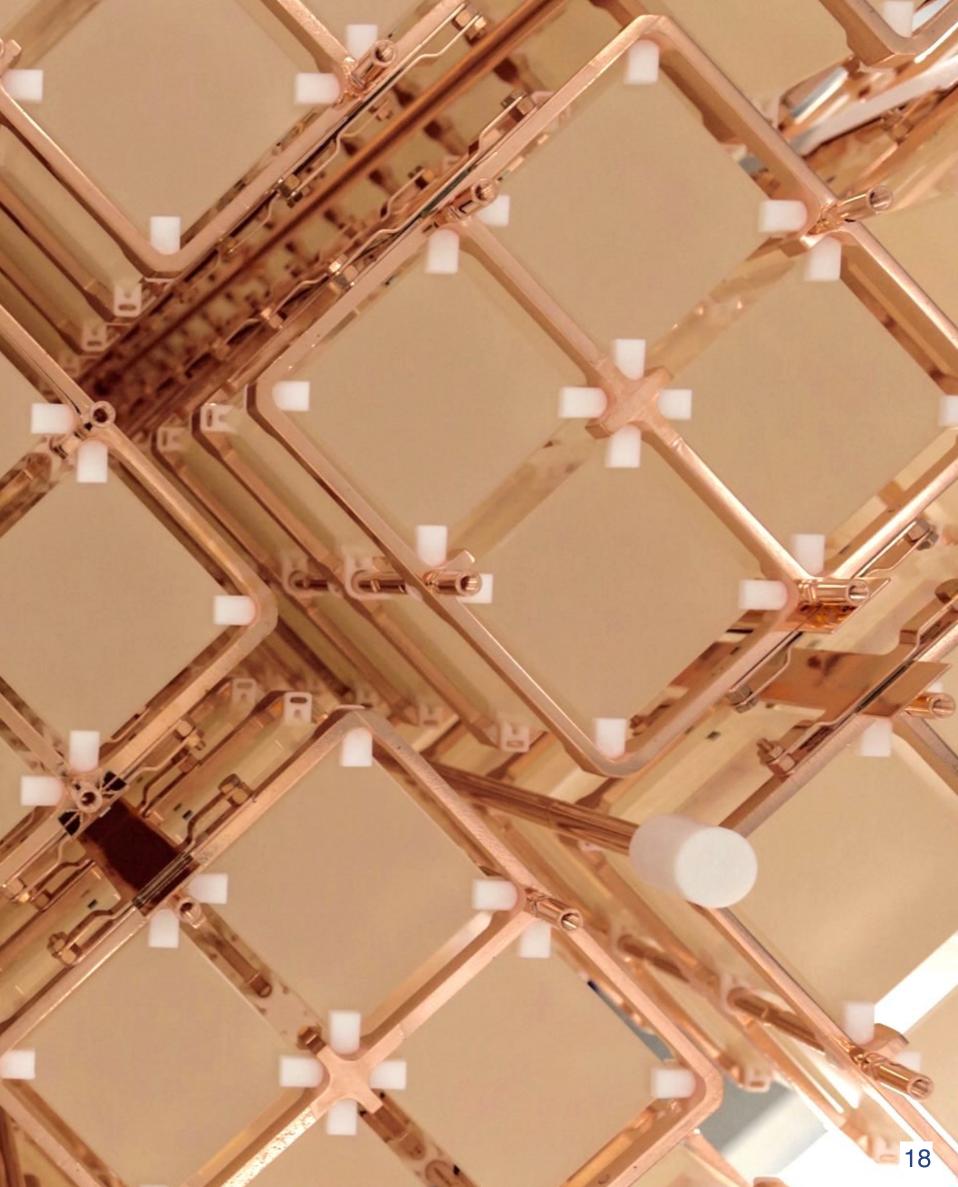
#### Winter 2018/2019

#### March 2019

#### Warm up and major maintenance

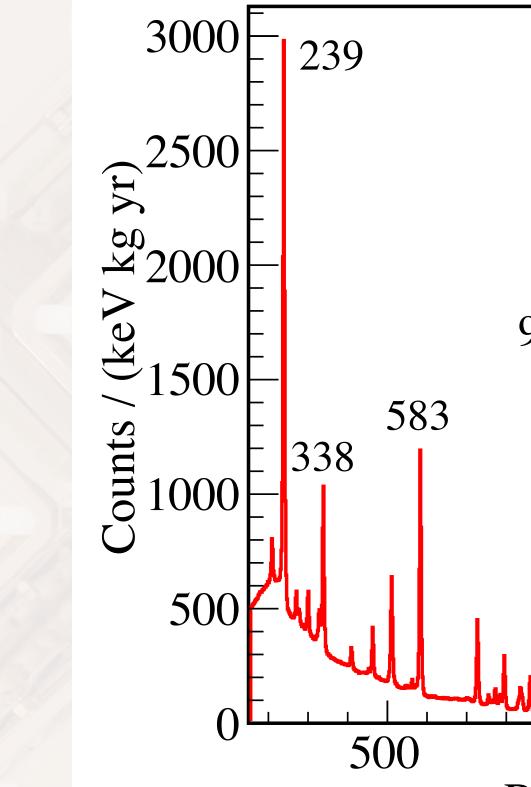
- Cleaned <sup>3</sup>He-<sup>4</sup>He line from air
- that prevented stable circulation
- reduce leaks and increase long
- Added external calibration system to reduce dead time due to strings deployment
- Fixed leaking pulse tubes to reestablish redundancy
- Ordinary maintenance of pumps

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### **Calibration spectrum**

- Calibration strings deployed inside and outside the CUORE detector (Th and Co)
- Summed energy spectrum of all the CUORE detectorsdatasets
- Calibration data used for:
  - energy scale calibration
  - thermal gain stabilisation
  - detector response (line shape) study

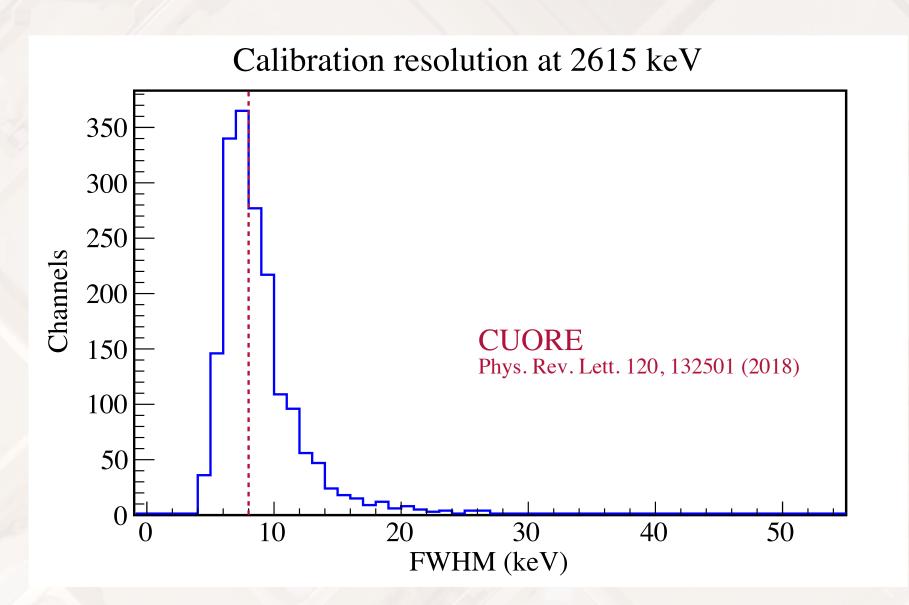




239 keV - <sup>212</sup>Pb 338, 911, 969 keV - <sup>228</sup>Ac 583, 2615 keV - 208TI CUORE Phys. Rev. Lett. 120, 132501 (2018) 911 969 2615 2000 2500 1000 1500 Reconstructed Energy (keV)

#### **Energy resolution**

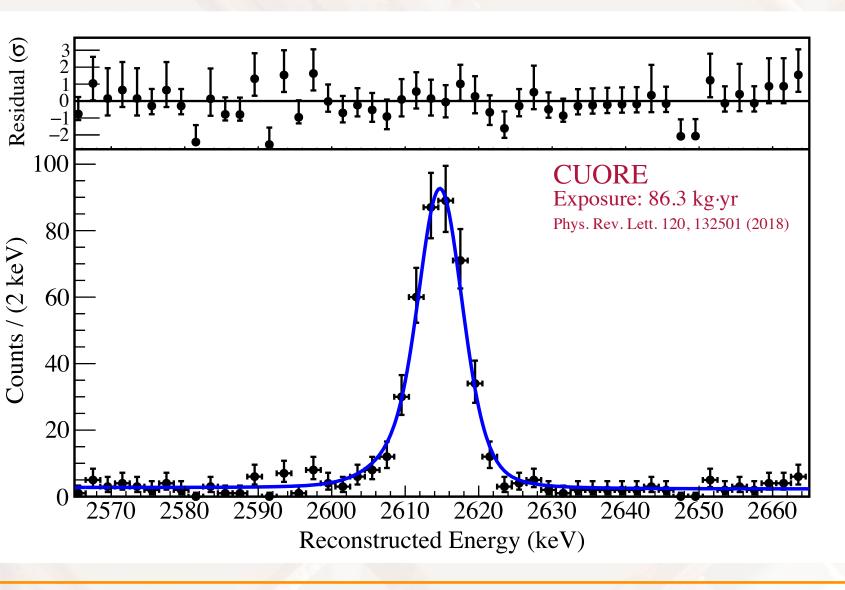
@ 2615 keV ds3018: 9.0 keV FWHM ds3021: 7.4 keV FWHM effective (exposure-weighted): 8.0 keV FWHM



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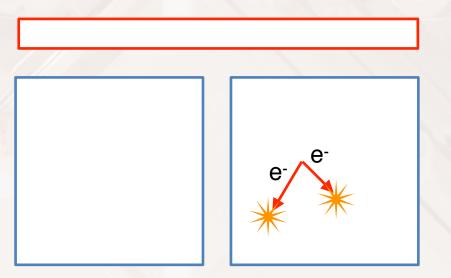
@ Q-value ds3018: (8.3 ± 0.4) keV FWHM ds3021: (7.4 ± 0.7) keV FWHM effective (exposure-weighted): (7.7 ± 0.5) keV FWHM



#### **Event selection**

Event selection occurs after periods of low-quality data (~1% of the total live time) are removed.

- Base quality cuts (number of pulses in the window, baseline stability, etc...)
- Anti-coincidence ullet



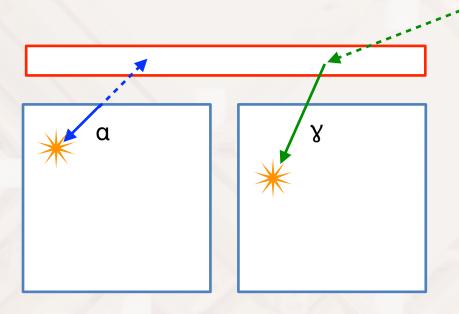
M1 -  $0\nu\beta\beta/2\nu\beta\beta$  signal like

M2 - rejected background and important information for bkg studies

Pulse shape analysis: deformed events are not used to build the final spectra to avoid • spectral shape distortions

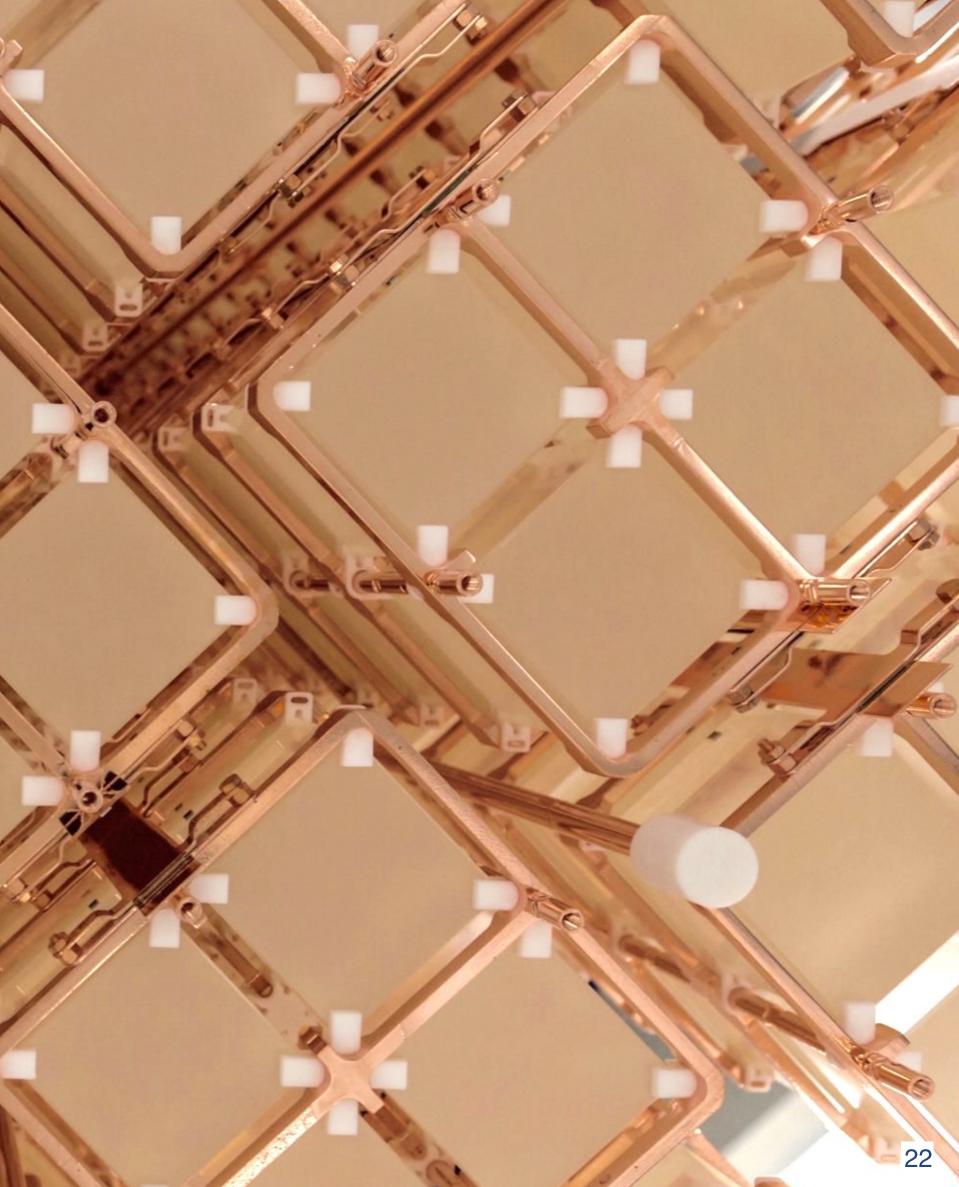
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M1 - dangerous background

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## Fit in the 0vDBD ROI

Region of interest: 2465 to 2575 keV

Overall signal efficiency: (75.7 ± 3.0)% - ds3018

(83.0 ± 2.6)% - ds3021

Events in the region of interest: 155

Residual (0) 16 CUORE Exposure: 86.3 kg·yr 14 Phys. Rev. Lett. 120, 132501 (2018)  $\mathbf{Q}_{\beta\beta}$ Counts / (2.5 keV) 12 **O** 2500 2520 2540 2560 2480 Reconstructed Energy (keV)

ROI background index:  $(1.49_{-0.17}^{+0.18}) \times 10^{-2} \text{ c/(keV \cdot kg \cdot yr)}$ 

Best fit for 60Co mean: (2506.4 ± 1.2) keV Best fit decay rate: (-1.0<sub>-0.3</sub>+0.4 (stat.) ± 0.1 (syst.))×10<sup>-25</sup> / yr

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 $(1.35_{-0.18}^{+0.20}) \times 10^{-2} \text{ c/(keV \cdot kg \cdot yr)}$ 

#### No evidence of signal

Decay rate limit (90% CL, including systematics): 0.51 × 10<sup>-25</sup> / yr Half-life limit (90% CL, including systematics): 1.3 × 10<sup>25</sup> yr Median expected sensitivity: 7.0 × 10<sup>24</sup> yr

> **CUORE, CUORE-0 and CUORICINO** combined 90% C.L. limit is  $T_{0v} > 1.5 \times 10^{25} \text{ yr}$

> > Phys. Rev. Lett. 120, 132501 (2018)

## Combination with previous 0vDBD results

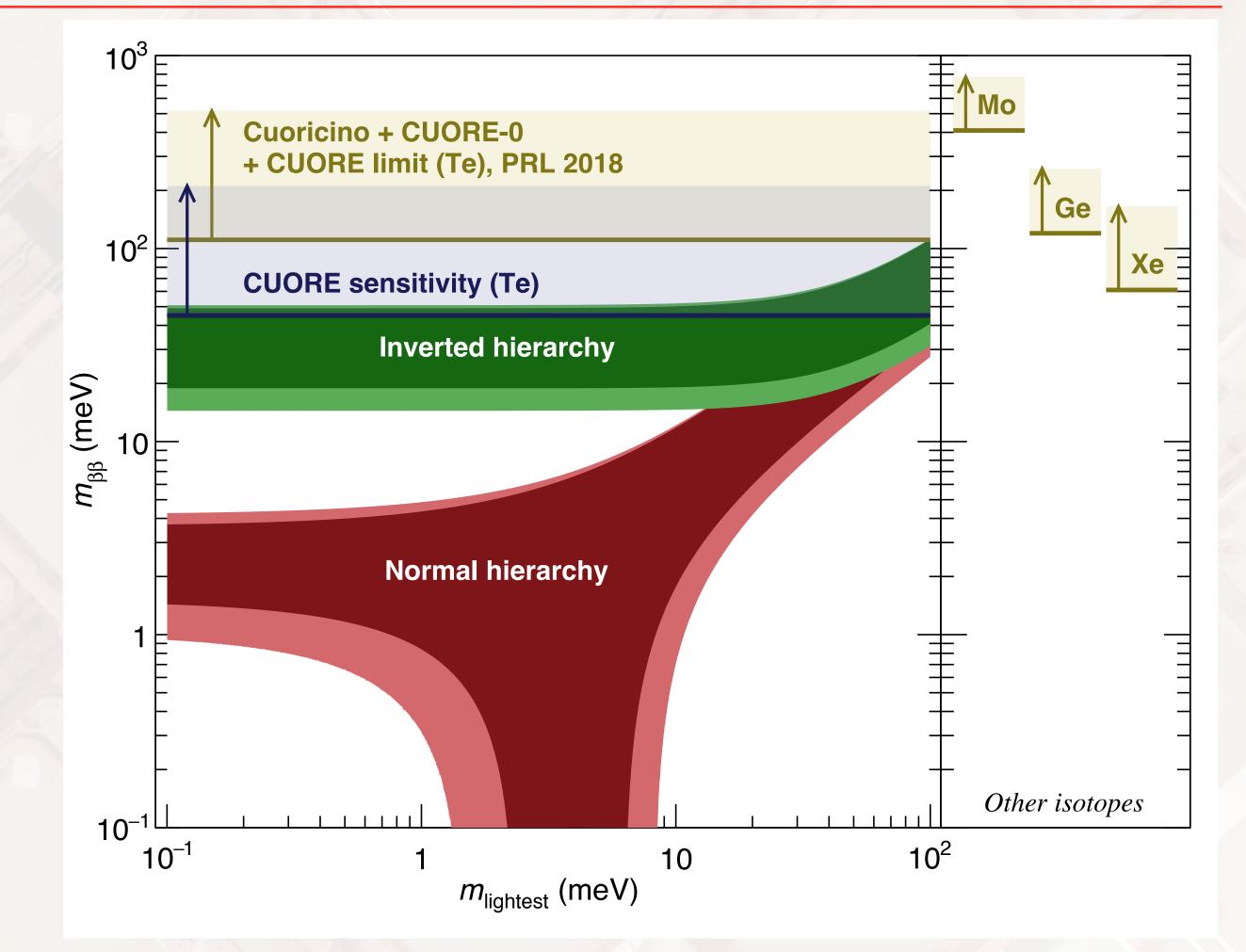
- Total <sup>130</sup>Te exposure
  - 86.3 kg·yr of CUORE
  - 19.75 kg·yr of Cuoricino
  - 9.8 kg·yr of CUORE-0
- The combined 90% C.L. limit is  $T_{0v} > 1.5 \times 10^{25} \text{ yr}$  $m_{\beta\beta} < 110-520 \text{ meV}$

#### NME:

JHEP02 (2013) 025 Nucl. Phys. A 818, 139 (2009) Phys. Rev. C 87, 045501 (2013) Phys. Rev. C 87, 064302 (2014) Phys. Rev. C 91, 034304 (2015) Phys. Rev. C 91, 024613 (2015) Phys. Rev. C 91, 024309 (2015) Phys. Rev. C 91, 024316 (2015) Phys. Rev. Lett. 105, 252503 (2010) Phys. Rev. Lett. 111, 142501 (2013)

Experiments:

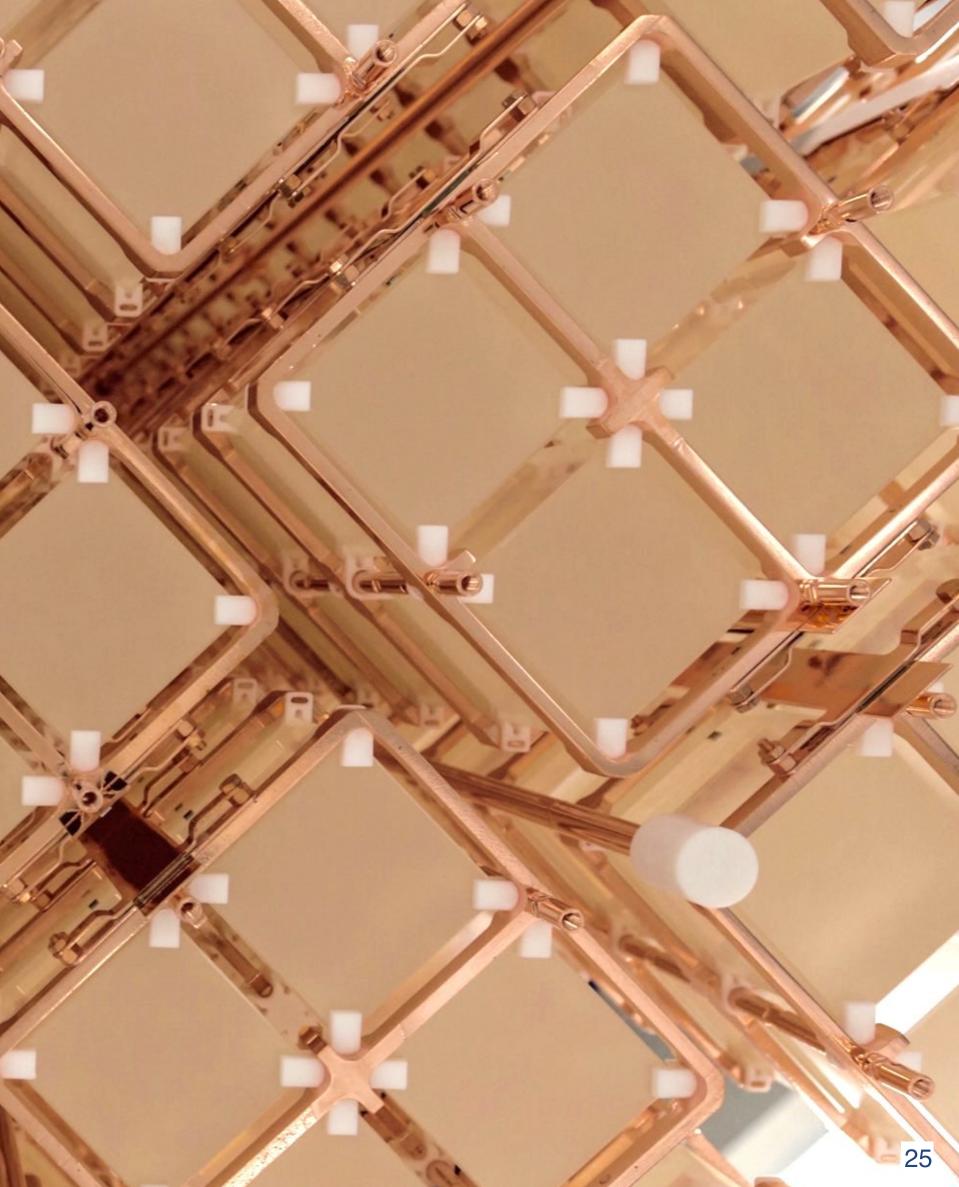
<sup>130</sup>Te:  $1.5 \times 10^{25}$  yr from PRL 120, 132501 (2018) <sup>76</sup>Ge:  $8.0 \times 10^{25}$  yr from PRL 120, 132503 (2018) <sup>136</sup>Xe:  $1.1 \times 10^{26}$  yr from Phys. Rev. Lett. 117, 082503 (2016) <sup>100</sup>Mo:  $1.1 \times 10^{24}$  yr from Phys. Rev. D 89, 111101 (2014) CUORE sensitivity:  $9.0 \times 10^{25}$  yr

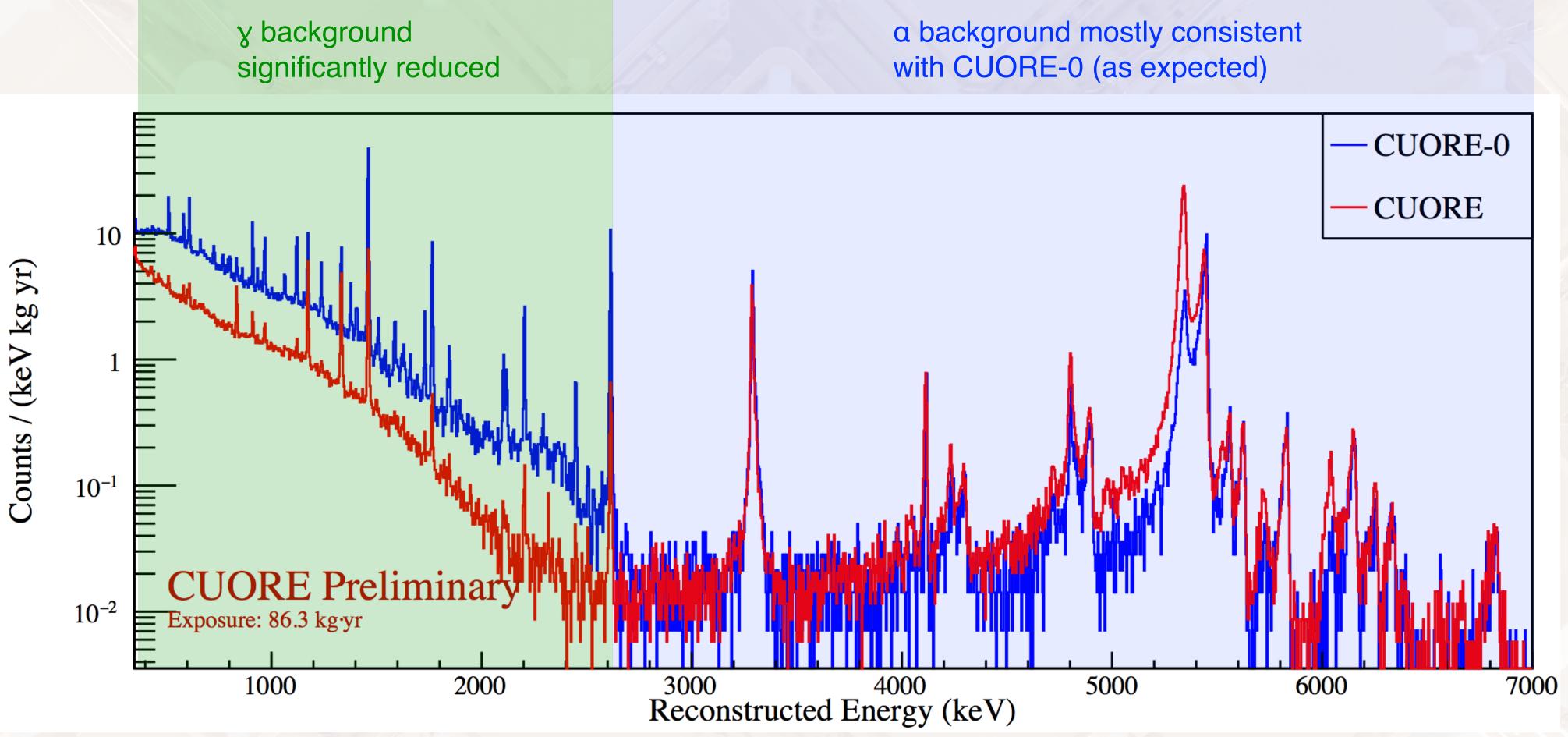


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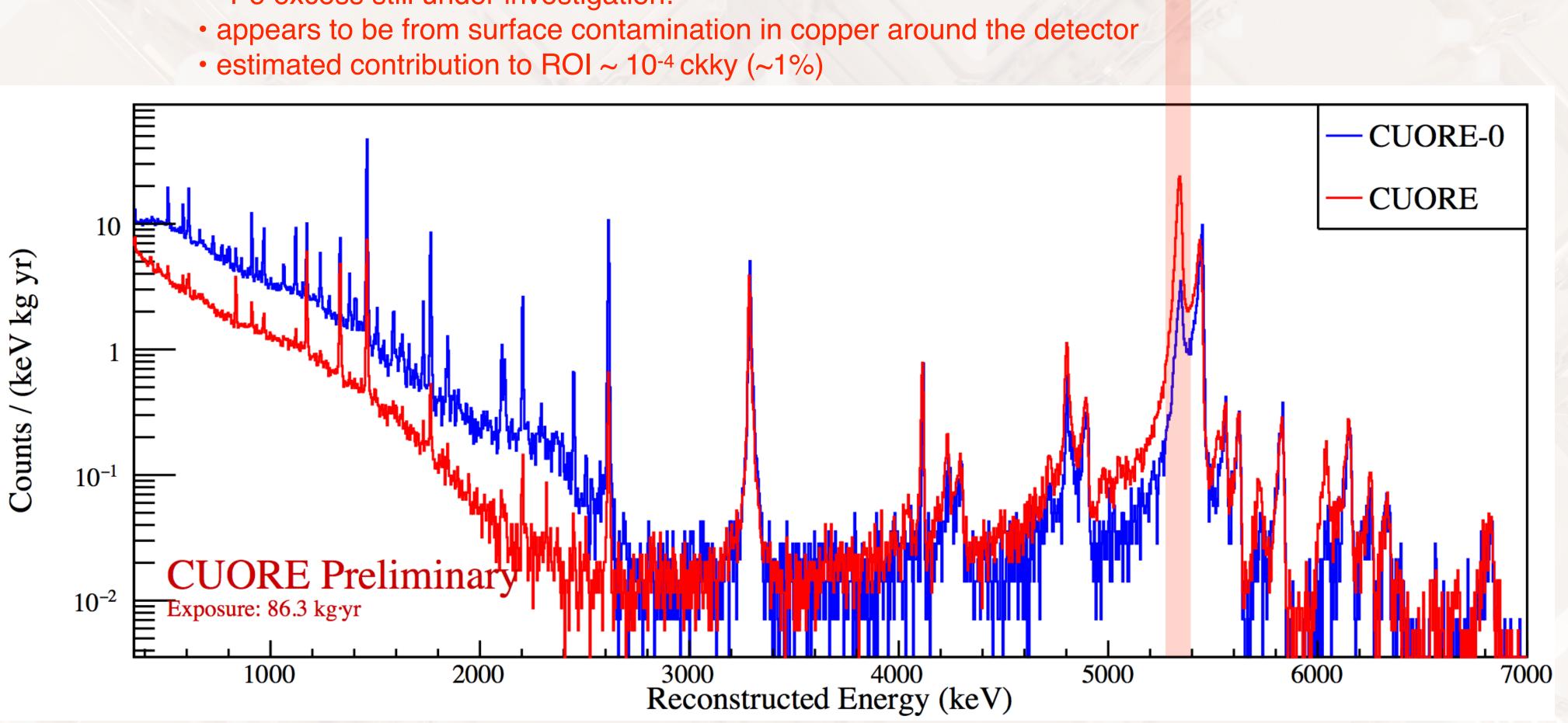




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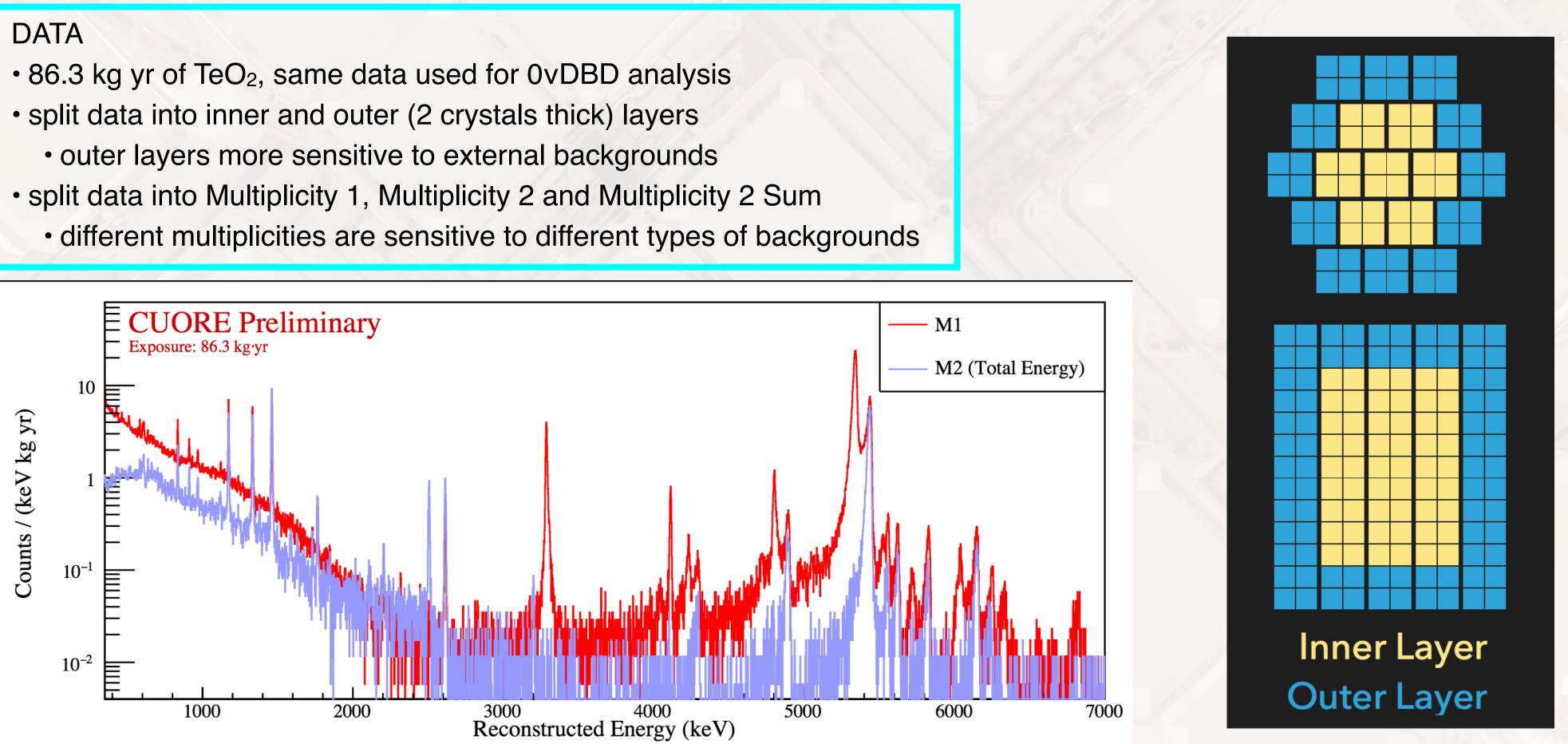


<sup>210</sup>Po excess still under investigation:



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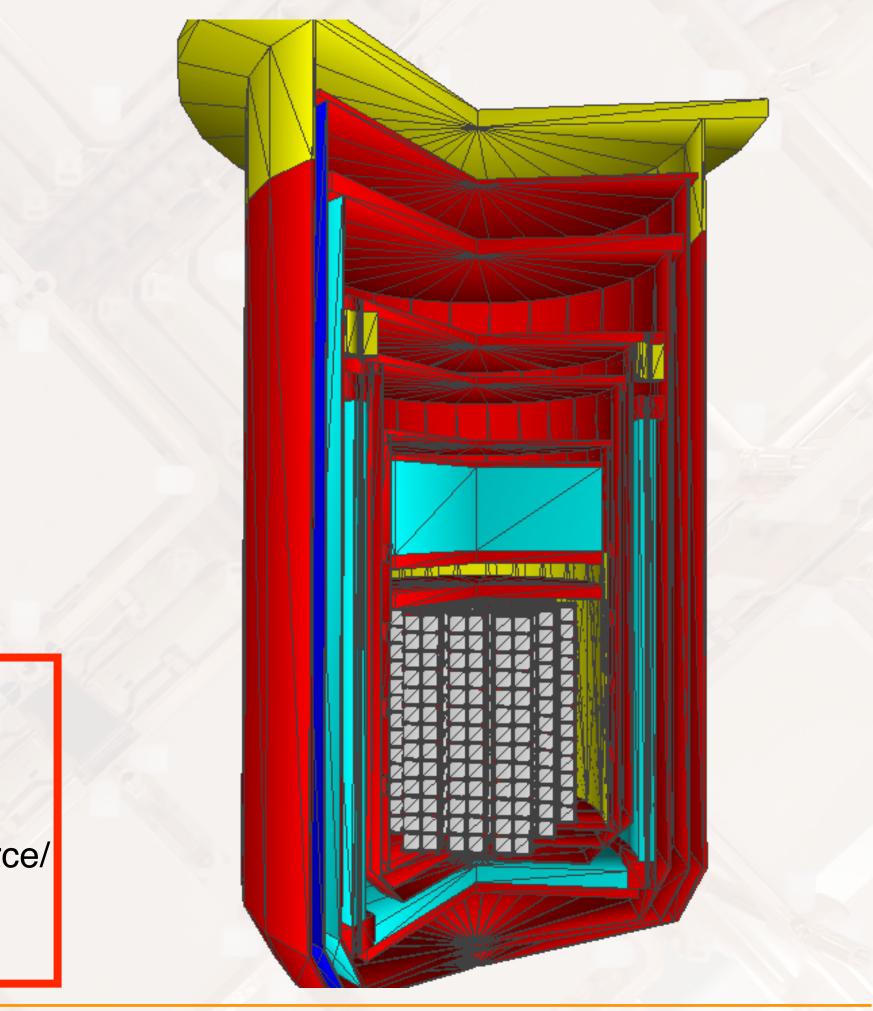
Volume	Туре	Components
TeO <sub>2</sub>	Bulk	$2 uetaeta$ , $^{210}$ Pb, $^{232}$ Th, $^{228}$ Ra- $^{208}$ Pb, $^{238}$ U- $^{230}$ Th, $^{230}$ Th $^{226}$ Ra- $^{210}$ Pb, $^{40}$ K, $^{60}$ Co, $^{125}$ Sb, $^{190}$ Pt
TeO <sub>2</sub>	Surface (0.01 $\mu$ m)	<sup>232</sup> Th, <sup>228</sup> Ra- <sup>208</sup> Pb, <sup>238</sup> U- <sup>230</sup> Th, <sup>226</sup> Ra- <sup>210</sup> Pb, <sup>210</sup> Pb
TeO <sub>2</sub>	Surface (1 $\mu$ m)	<sup>210</sup> Pb
TeO <sub>2</sub>	Surface (10 $\mu$ m)	<sup>210</sup> Pb, <sup>232</sup> Th, <sup>238</sup> U
CuNOSV	Bulk	<sup>232</sup> Th, <sup>238</sup> U, <sup>40</sup> K, <sup>60</sup> Co, <sup>54</sup> Mn
CuNOSV	Surface (0.01 $\mu$ m)	<sup>210</sup> Pb, <sup>232</sup> Th, <sup>238</sup> U
CuNOSV	Surface (1 $\mu$ m)	<sup>210</sup> Pb, <sup>232</sup> Th, <sup>238</sup> U
CuNOSV	Surface (10 $\mu$ m)	<sup>210</sup> Pb, <sup>232</sup> Th, <sup>238</sup> U
Roman lead	Bulk	<sup>232</sup> Th, <sup>238</sup> U, <sup>108</sup> <i>m</i> Ag
Top lead	Bulk	<sup>232</sup> Th, <sup>238</sup> U, <sup>210</sup> Bi
Ext. lead	Bulk	<sup>210</sup> Bi
CuOFE	Bulk	<sup>232</sup> Th, <sup>238</sup> U, <sup>60</sup> Co
External	-	Cosmic muons

#### MONTE CARLO

- ~60 independent simulations of sources/location in the setup
- full radioactive chains and single isotopes
- the different energy spectra (inner/outer, M1/M2,M2sum) of each source/ location are generated with Geant4 based simulation implementing a detailed geometry of the setup (detector, cryostat, shields)

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#### DATA

- 86.3 kg yr of TeO<sub>2</sub>, same data used for 0vDBD analysis
- split data into inner and outer (2 crystals thick) layers
  - outer layers more sensitive to external backgrounds
- split data into Multiplicity 1, Multiplicity 2 and Multiplicity 2 Sum
  - different multiplicities are sensitive to different types of backgrounds

## Assign to each ingredient (source/location) a normalisation factor Assign a prior to each normalisation factor **BAYESIAN FIT** from material screening, assays and cosmogenic analysis • Fit the model to the data and sample the posteriors with MCMC Gibbs sampler

#### **MONTE CARLO**

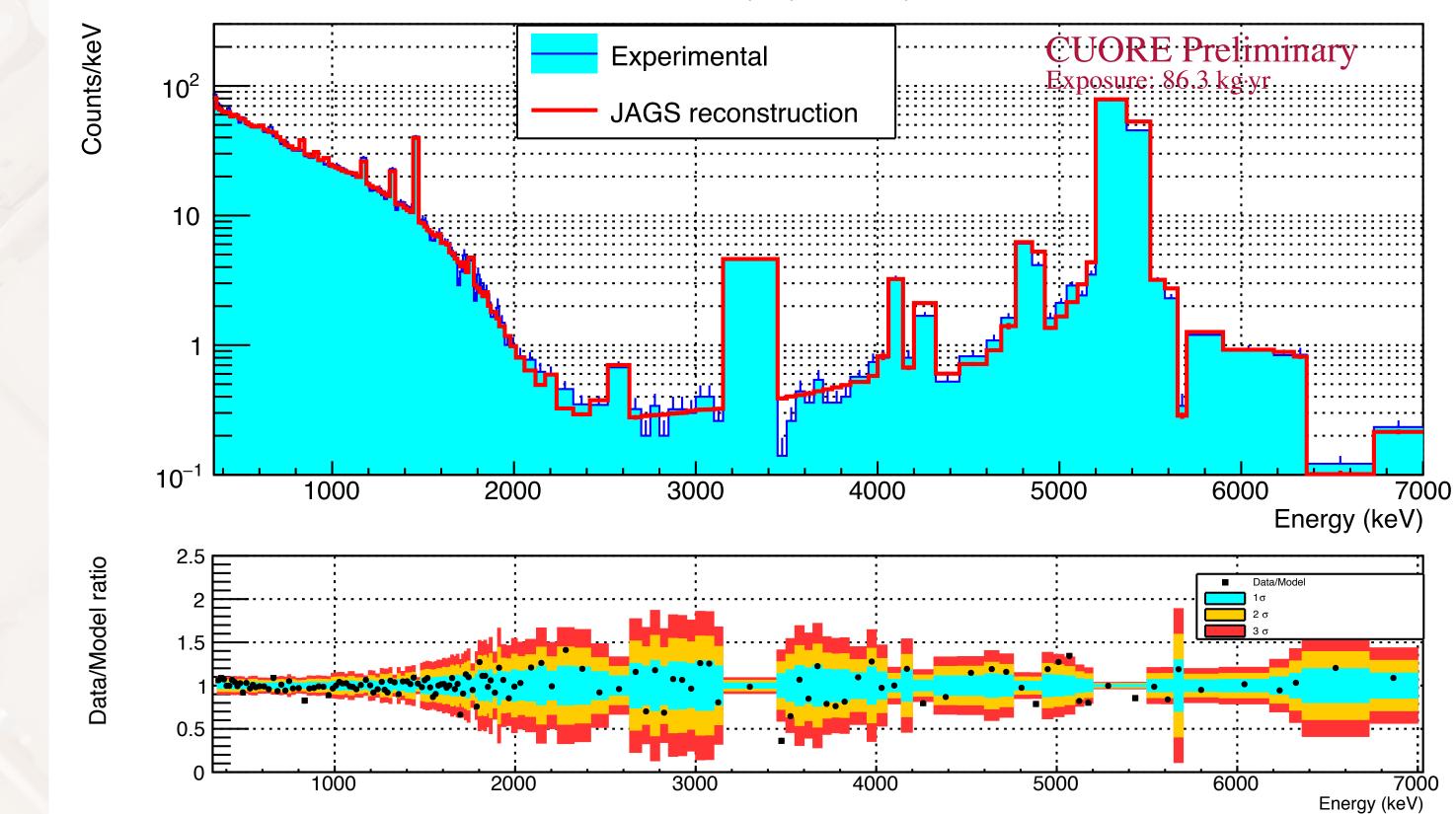
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#### **BAYESIAN FIT**

Multiplicity 1 - Inner layer

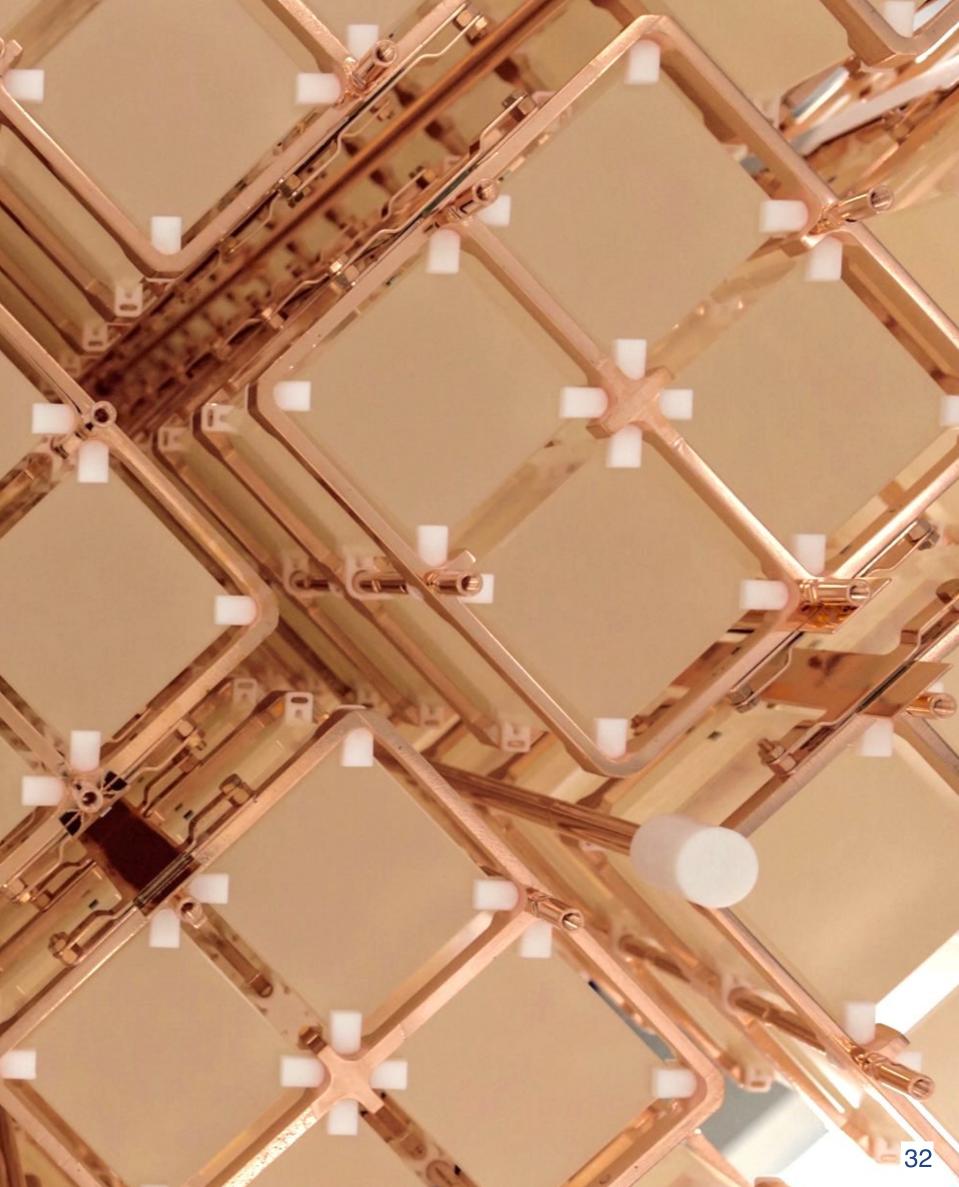


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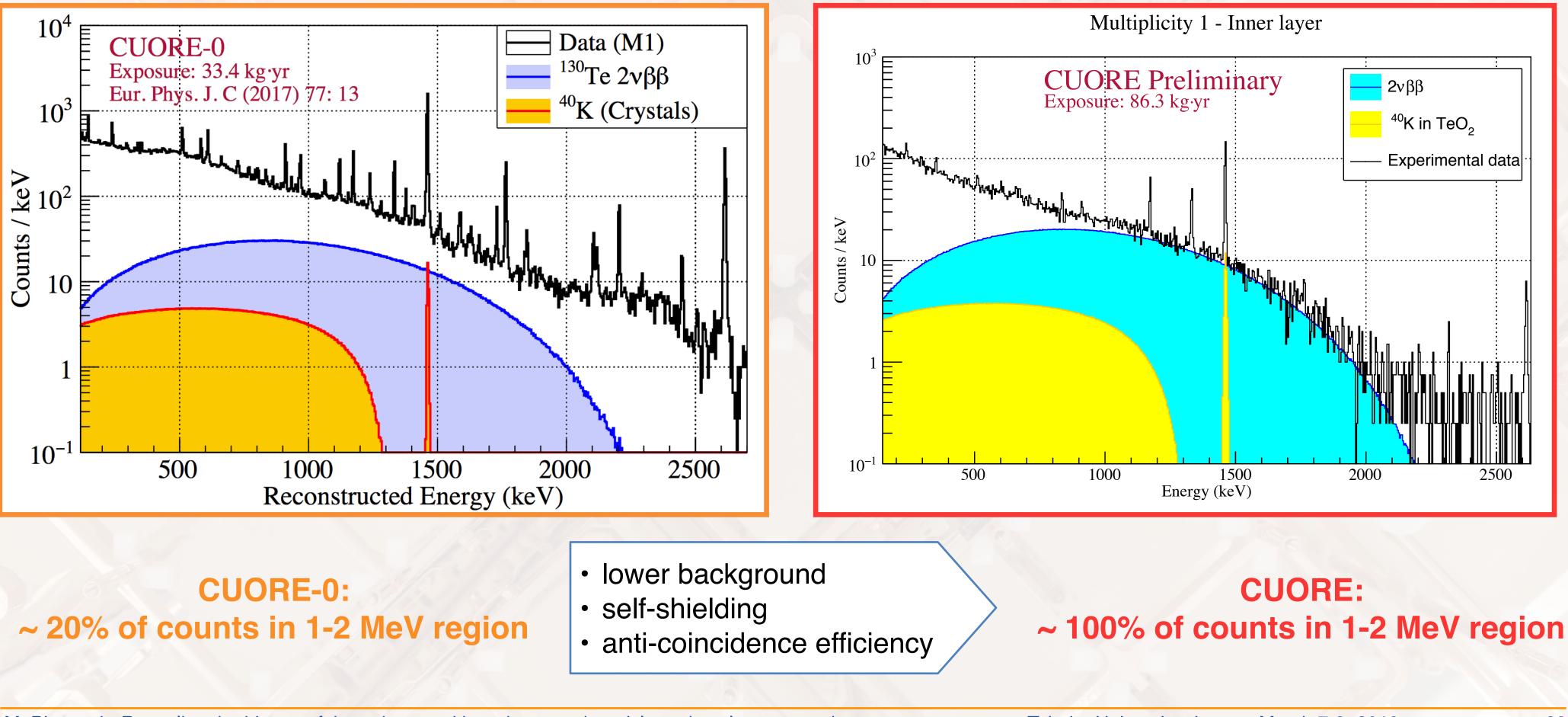


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### 2vDBD

2vDBD spectrum is one of the ingredients in the background model

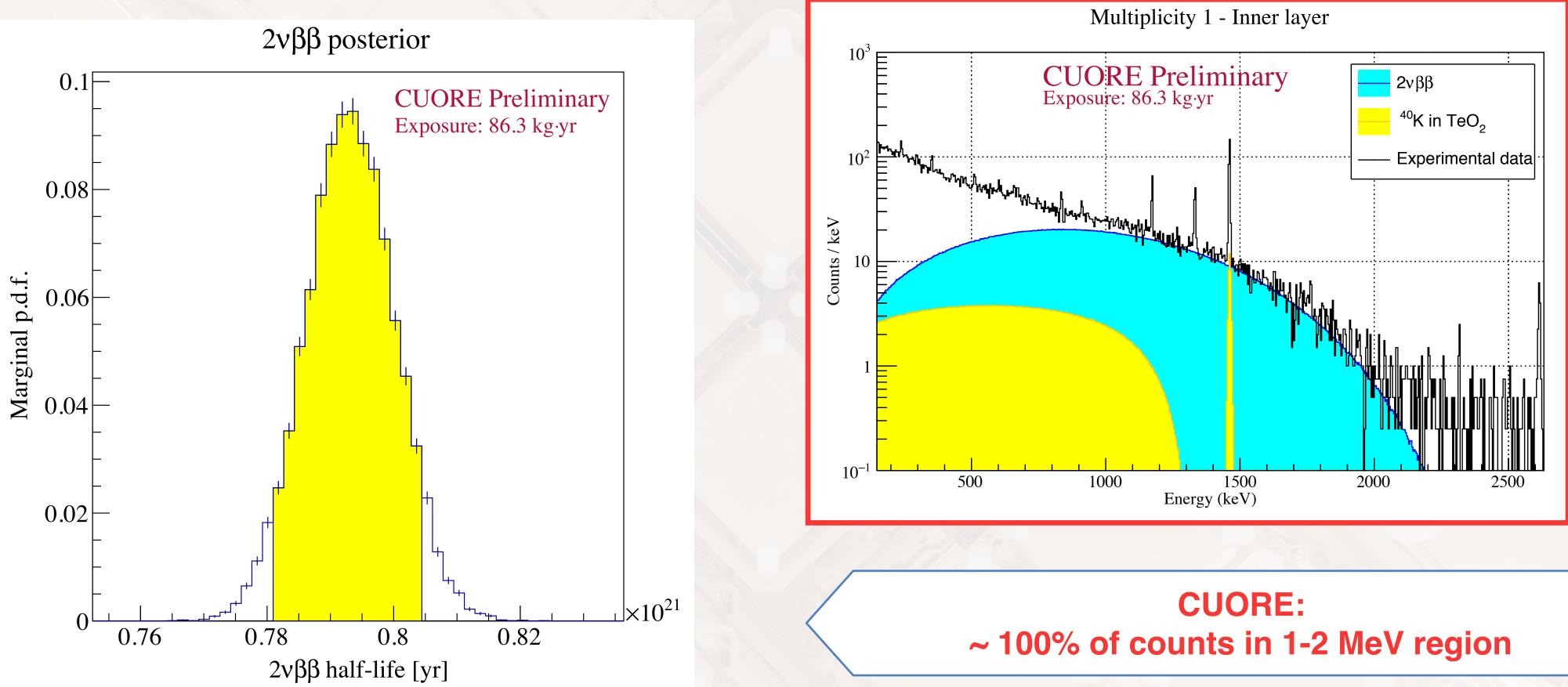


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#### 2vDBD

#### 2vDBD spectrum is one of the ingredients in the background model



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### 2vDBD

#### 2vDBD spectrum is one of the ingredients in the background model

Counts / keV

 $10^{-10}$ 

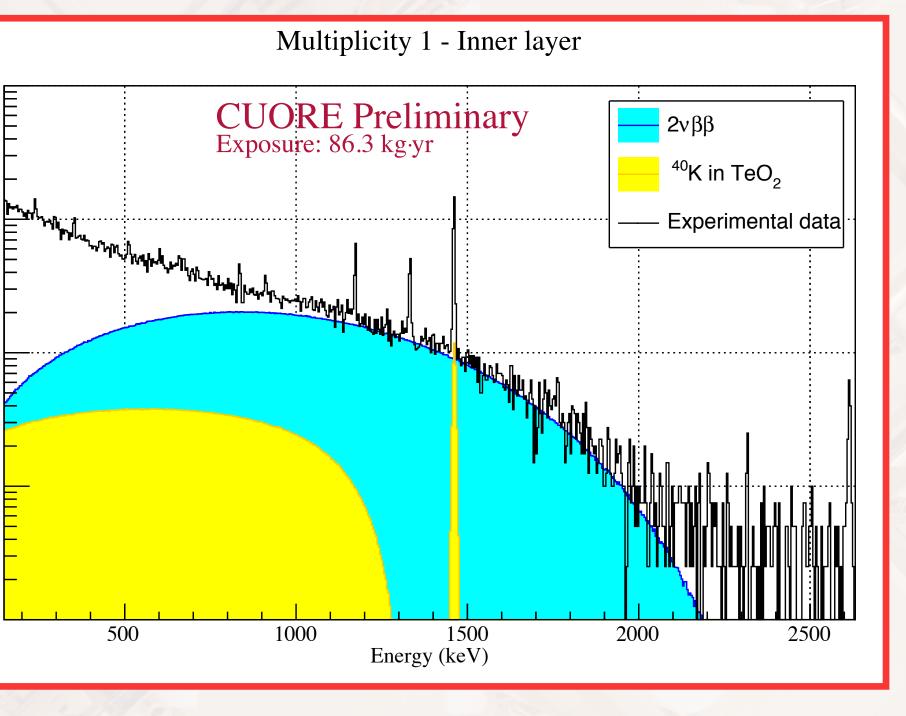
$$\Gamma_{1/2}^{2\nu} = [8.7 \pm 0.1 (\text{stat.}) \pm 0.2 (\text{syst.})] \times 10^{-22} \text{yr}^{-1}$$
$$T_{1/2}^{2\nu} = [7.9 \pm 0.1 (\text{stat.}) \pm 0.2 (\text{syst.})] \times 10^{20} \text{yr}^{-1}$$

CUORE - 0 :  $T_{1/2}^{2\nu} = [8.2 \pm 0.2 (\text{stat.}) \pm 0.6 (\text{syst.})] \times 10^{20} \text{yr}$ NEMO - 3 :  $T_{1/2}^{2\nu} = [7.0 \pm 0.9 (\text{stat.}) \pm 1.1 (\text{syst.})] \times 10^{20} \text{yr}$ 

- Systematic uncertainty dominated by uncertainty on contaminant location and uniformity
  - studied by repeating the fit with different geometric splitting of the data
  - more data —> finer geometric splitting of the array —> reduce this component
- Fit is independent of energy threshold over the range 100 750 keV
- Used M2 data (purer sample of particle events) to improve the systematic uncertainty on the selection efficiency by one order of magnitude w.r.t. PRL 2018 data release

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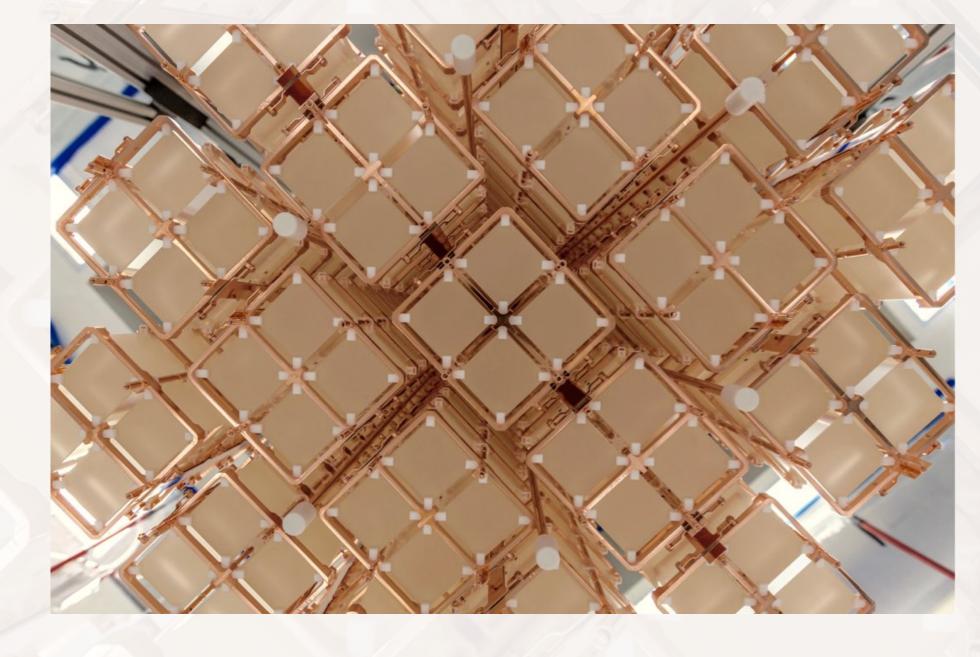




# CUORE: ~ 100% of counts in 1-2 MeV region

### **Conclusions and Outlook**

- With the first two datasets CUORE have:
  - accumulated a total exposure of almost 100 kg·y •
  - Invaluable operational experience •
  - collected important information on detector performance, noise, resolutions, background levels •
  - pushed for the first time the limit on neutrino-less double beta decay half life of <sup>130</sup>Te beyond 10<sup>25</sup> years
  - performed the most precise measure of two-neutrino half life to date



- physics data
- - ٠
- experiments
- coming soon



The largest and most complex cryogenic experiment is taking

The first analysis efforts were focused on the neutrino-less and two-neutrino double beta decay of <sup>130</sup>Te to GS

Physics results on more processes are on their way:

Majoron emission, CPTV

Dark matter, axions

 $\beta\beta$  to excited states,  $\beta$ +/EC decays

With an unprecedented amount of data, CUORE is the best tool to

study and model the backgrounds for the next generation

Data release and physics results with more then doubled statistics