

# Introduction

• From experiments neutrinos experience flavor oscillations • This leads to massive neutrinos, which are beyond the SM

 $m_{\nu 2}^2 - m_{\nu 1}^2 \simeq 7.42 \times 10^{-5} \text{eV}^2$   $|m_{\nu 3}^2 - m_{\nu 2}^2| \simeq 2.517 \times 10^{-3} \text{eV}^2$ 

• Massive neutrinos are allowed two types of mass • Determination of the mass type is an important goal in physics

Dirac mass or Majorana mass

$$\overline{\nu}_R m_D \nu_L \quad \frac{1}{2} m_M \overline{(\nu_L)^c} \nu_L$$

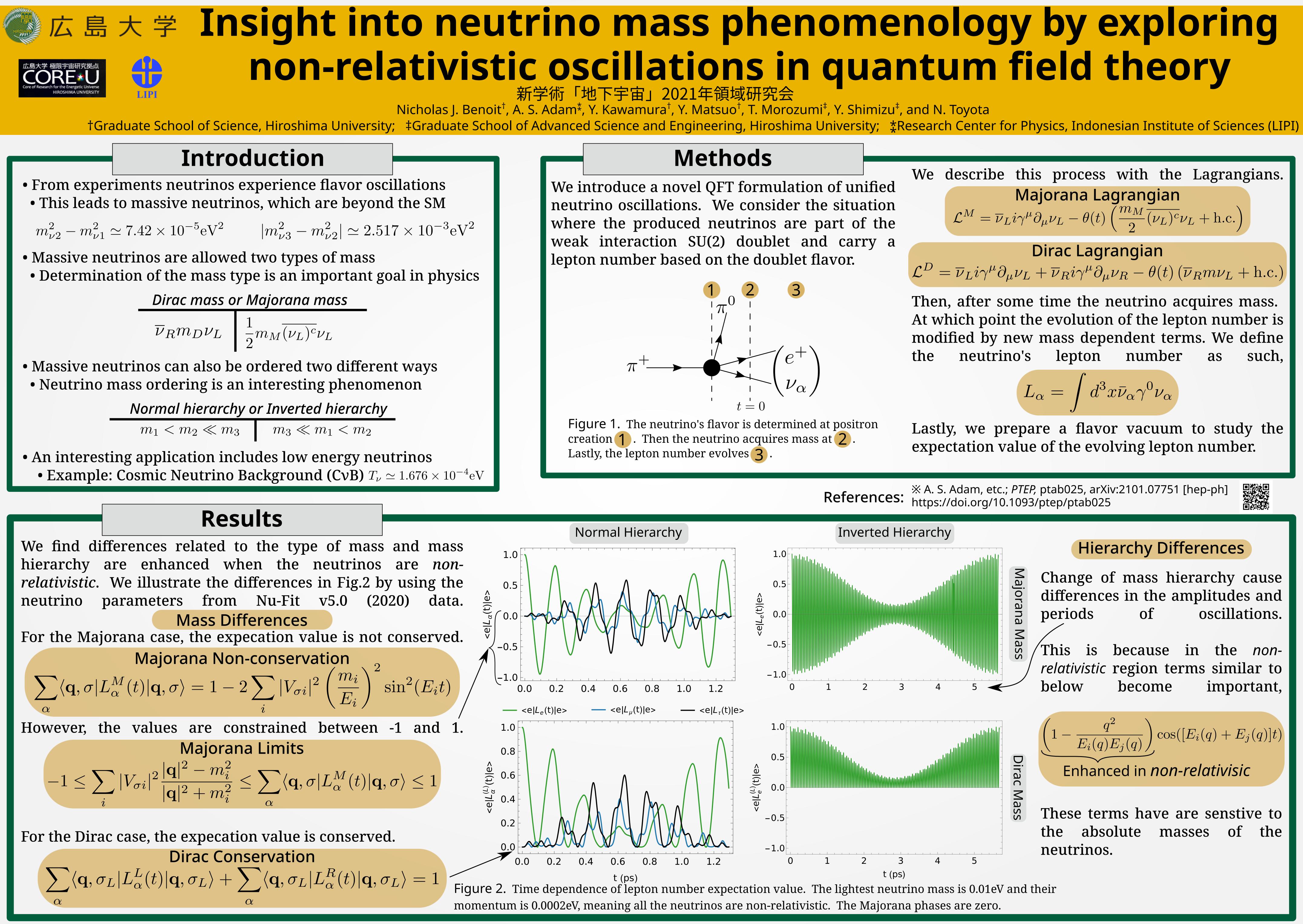
• Massive neutrinos can also be ordered two different ways • Neutrino mass ordering is an interesting phenomenon

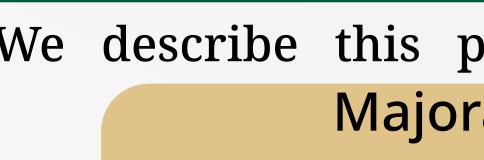
Normal hierarchy or Inverted hierarchy

• An interesting application includes low energy neutrinos • Example: Cosmic Neutrino Background (CvB)  $T_{\nu} \simeq 1.676 \times 10^{-4} \text{eV}$ 

## Results

We find differences related to the type of mass and mass hierarchy are enhanced when the neutrinos are non*relativistic.* We illustrate the differences in Fig.2 by using the parameters from Nu-Fit v5.0 neutrino Mass Differences For the Majorana case, the expecation value is not conserved. Majorana Non-conservation  $\sum \langle \mathbf{q}, \sigma | L_{\alpha}^{M}(t) | \mathbf{q}, \sigma \rangle = 1 - 2 \sum |V_{\sigma i}|^{2}$ However, the values are constrained between -1 and 1. Majorana Limits  $-1 \leq \sum_{i} |V_{\sigma i}|^2 \frac{|\mathbf{q}|^2 - m_i^2}{|\mathbf{q}|^2 + m_i^2} \leq \sum_{i} \langle \mathbf{q}, \sigma | L_{\alpha}^M(t) | \mathbf{q}, \sigma \rangle \leq 1$ For the Dirac case, the expecation value is conserved. **Dirac Conservation**  $\sum \langle \mathbf{q}, \sigma_L | L_{\alpha}^L(t) | \mathbf{q}, \sigma_L \rangle + \sum \langle \mathbf{q}, \sigma_L | L_{\alpha}^R(t) | \mathbf{q}, \sigma_L \rangle = 1$ 





Majorana Lagrangian Dirac Lagrangian lepton number such, as  $d^3 x \bar{\nu}_{\alpha} \gamma^0 \nu_{\alpha}$  $L_{\alpha} = /$ Hierarchy Differences Change of mass hierarchy cause differences in the amplitudes and oscillations. periods of This because in the *non*-**1**S relativistic region terms similar to below become important,  $\cos([E_i(q) + E_j(q)]t)$  $E_i(q)E_j(q)$ Enhanced in *non-relativisic* These terms have are sensitve to absolute masses of the the neutrinos.



We describe this process with the Lagrangians.  $\mathcal{L}^{M} = \overline{\nu}_{L} i \gamma^{\mu} \partial_{\mu} \nu_{L} - \theta(t) \left( \frac{m_{M}}{2} \overline{(\nu_{L})^{c}} \nu_{L} + \text{h.c.} \right)$  $\mathcal{L}^{D} = \overline{\nu}_{L} i \gamma^{\mu} \partial_{\mu} \nu_{L} + \overline{\nu}_{R} i \gamma^{\mu} \partial_{\mu} \nu_{R} - \theta(t) \left( \overline{\nu}_{R} m \nu_{L} + \text{h.c.} \right)$ Then, after some time the neutrino acquires mass. At which point the evolution of the lepton number is modified by new mass dependent terms. We define Lastly, we prepare a flavor vacuum to study the expectation value of the evolving lepton number. **References:** X A. S. Adam, etc.; *PTEP*, ptab025, arXiv:2101.07751 [hep-ph] https://doi.org/10.1093/ptep/ptab025

