

XENON

# XENONnT/XLZD実験の現状と 極低放射能技術

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On behalf of the XENON Collaboration

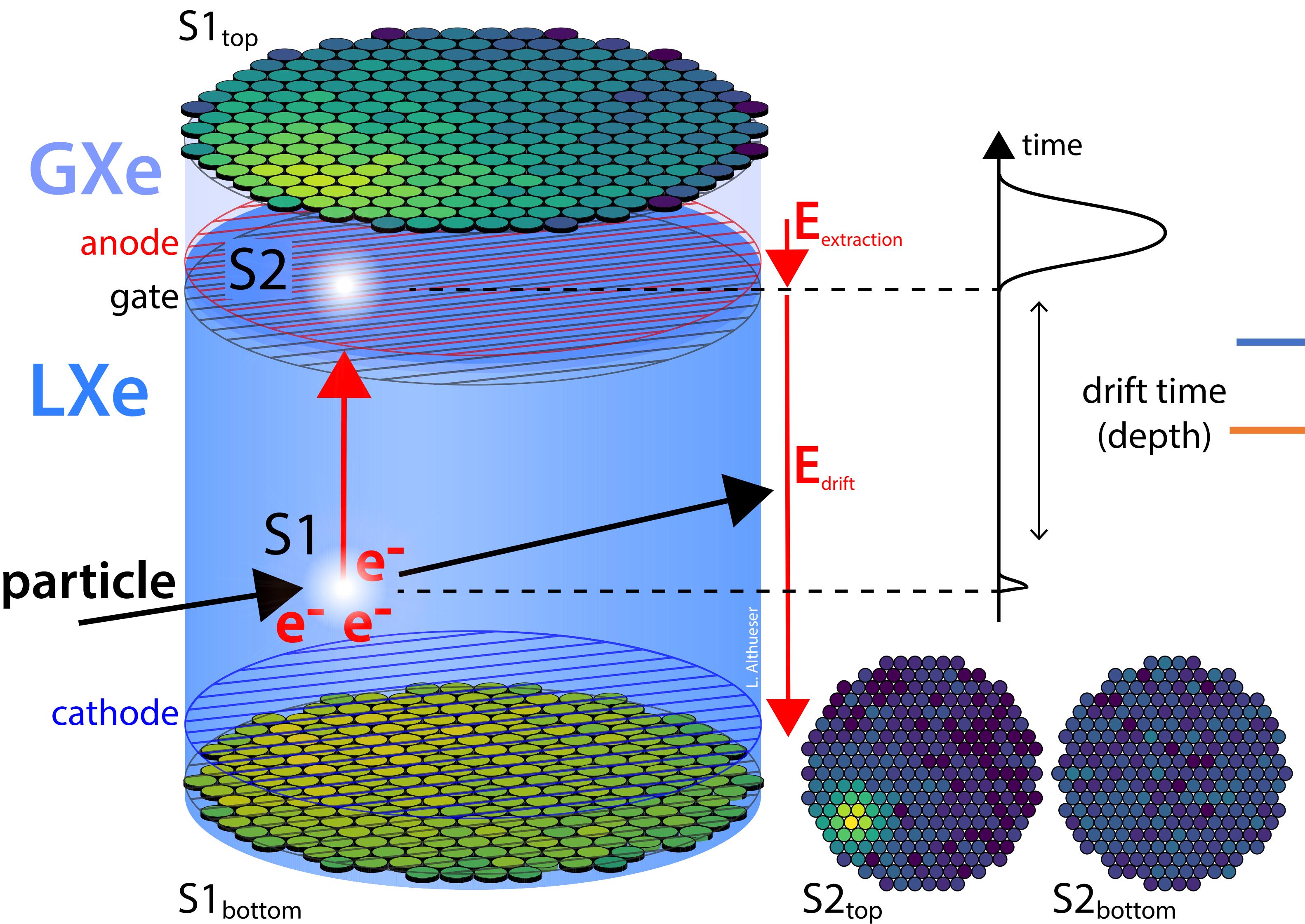
# XENONnT

# XENON Collaboration

> 200 scientists  
29 institutions  
12 countries



# XENONnT TPC

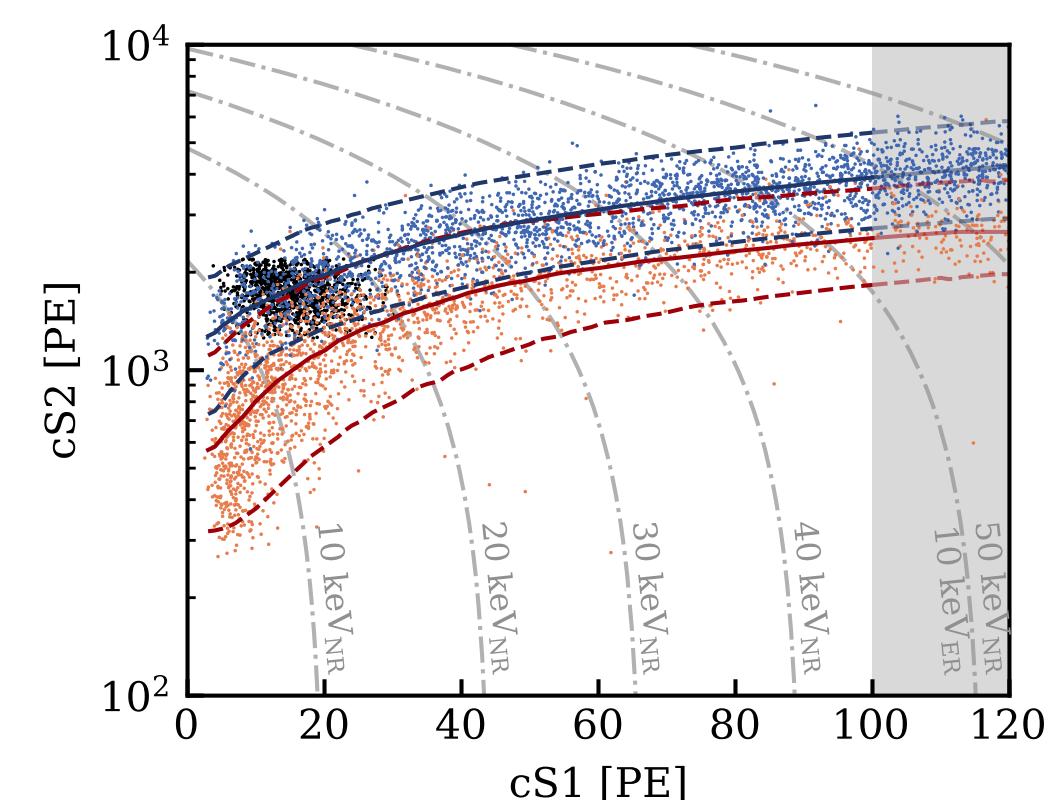


Dual phase Liq. Xe TPC w/ 5.9 t sensitive mass

- Self shielding
- 3D vertex reconstruction
  - x, y : S2 hit pattern
  - z : drift time
- ER/NR discrimination by S2/S1 ratio

electron recoil (ER)  
 $\beta, \gamma, \nu$

nuclear recoil (NR)  
 WIMPs, n, v



# XENONnT Timeline

2020

## Construction

2020.03 – 12

2021

## Commissioning

2021.01 – 06

## Science Run 0 (SR0)

2021.07 – 11

## Science Run 1 (SR1)

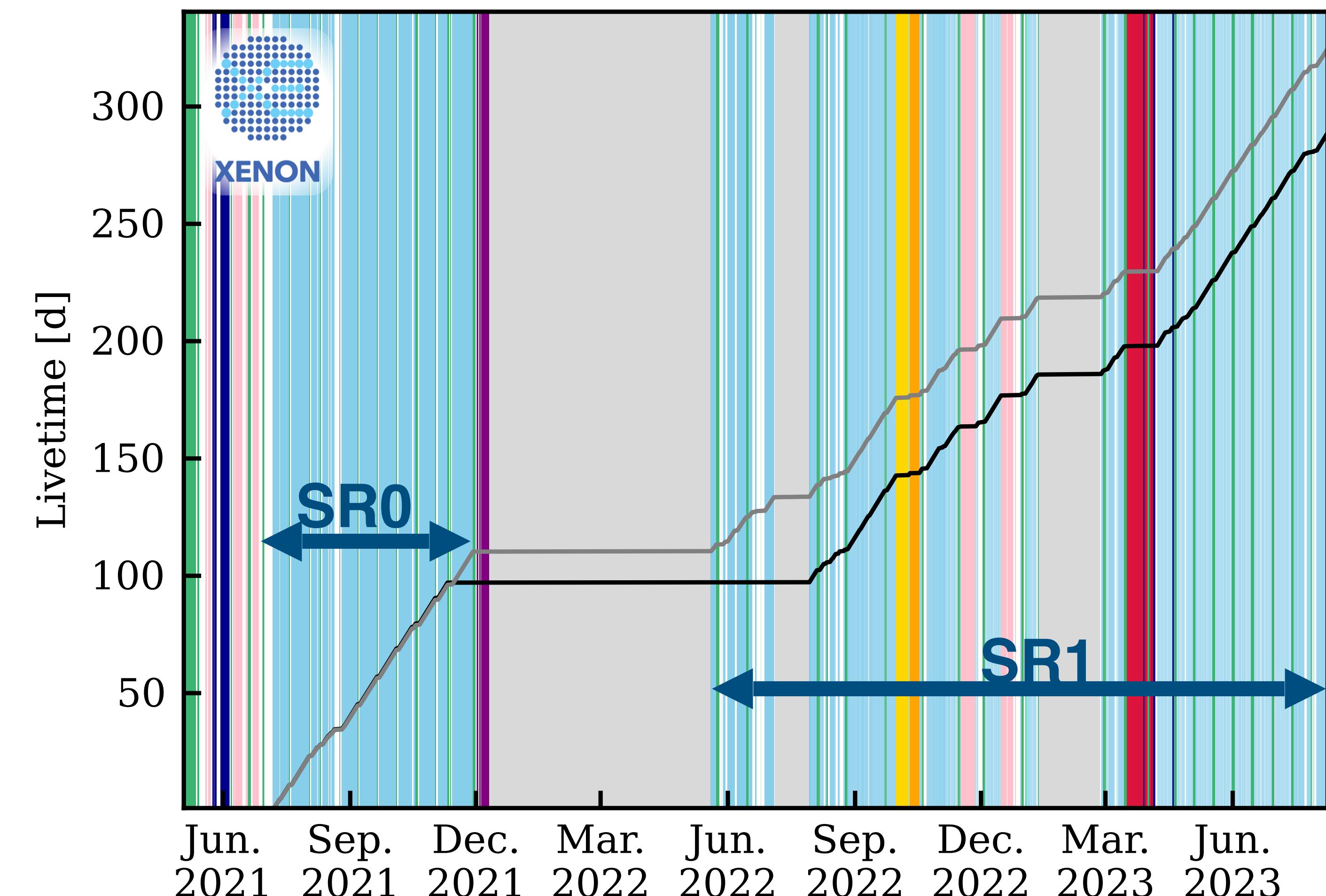
2022.05 – 2023.08

## GdSO loading into veto detector

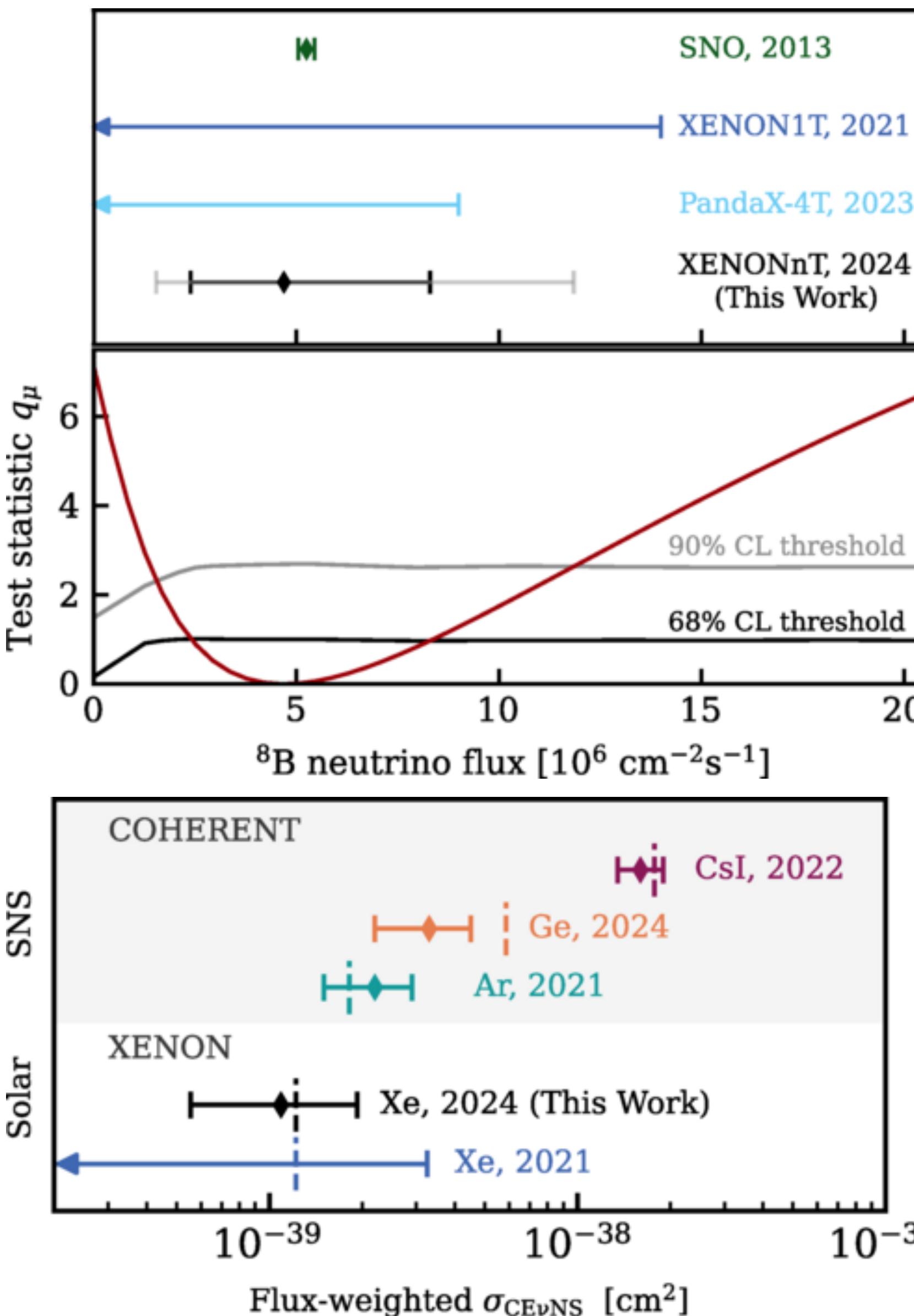
## Science Run 2 (SR2)

2023.10 –

2024  
2025



# Recent results: Solar ${}^8\text{B}$ CEvNS

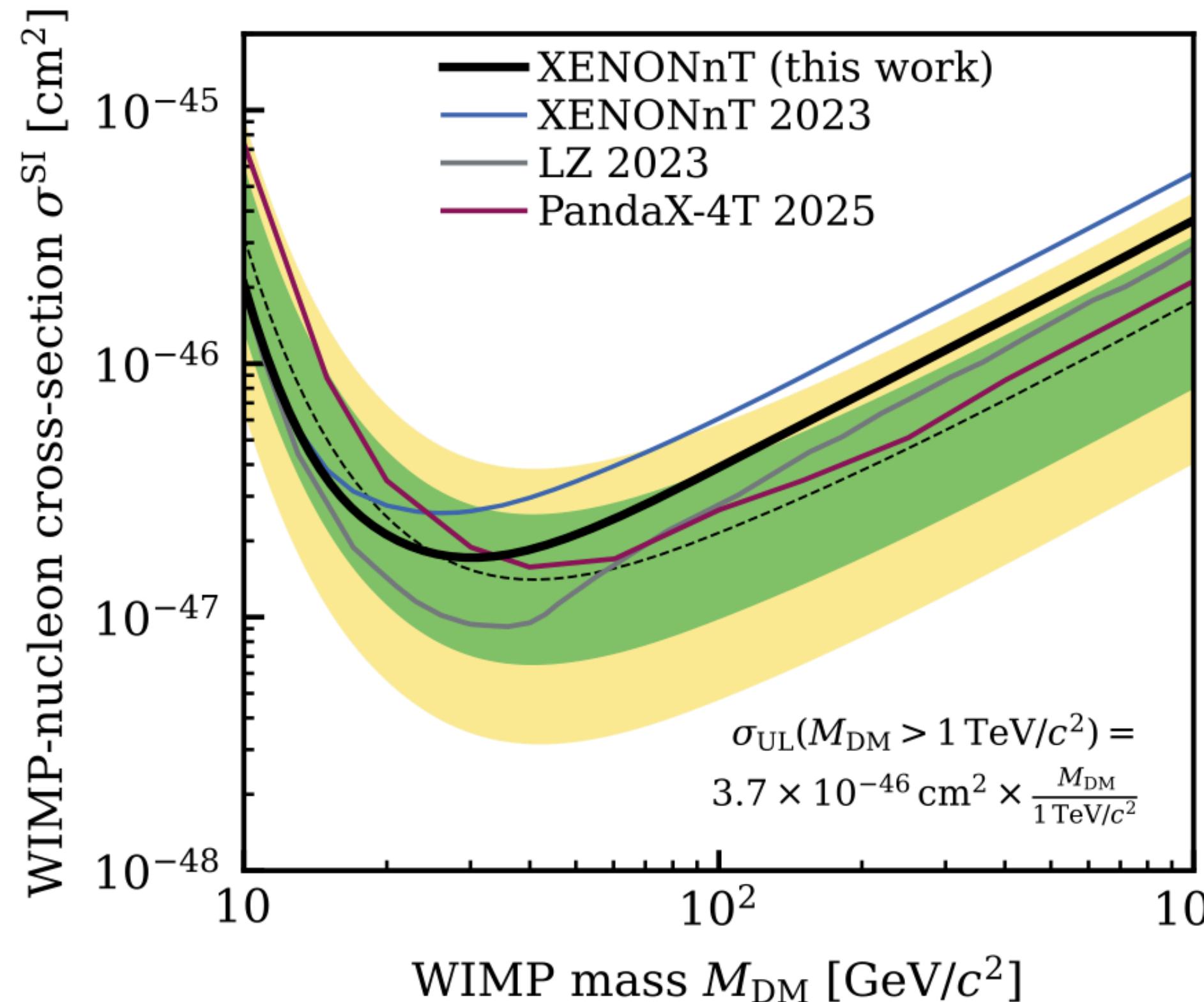


First indication of Solar  ${}^8\text{B}$   $\nu$  CEvNS  
(Coherent Elastic  $\nu$ -Nucleus Scattering)

- 2-fold S1 → lower threshold
- Main BG: Accidental Coincidence (AC)  
: mistakenly paired S1 and S2
- **$2.73\sigma$**  significance

Component	Expectation	Best fit
AC (SR0)	$7.5 \pm 0.7$	$7.4 \pm 0.7$
AC (SR1)	$17.8 \pm 1.0$	$17.9 \pm 1.0$
ER	$0.7 \pm 0.7$	$0.5^{+0.7}_{-0.6}$
Neutron	$0.5^{+0.2}_{-0.3}$	$0.5 \pm 0.3$
Total background	$26.4^{+1.4}_{-1.3}$	$26.3 \pm 1.4$
${}^8\text{B}$	$11.9^{+4.5}_{-4.2}$	$10.7^{+3.7}_{-4.2}$
Observed	37	

# Recent results: WIMPs



----- median sensitivity  
█  $\pm 1\sigma$   
█  $\pm 2\sigma$   
neutron BG is suppressed by  
veto detector (pure water Cherenkov)  
w/ tagging efficiency of 53%

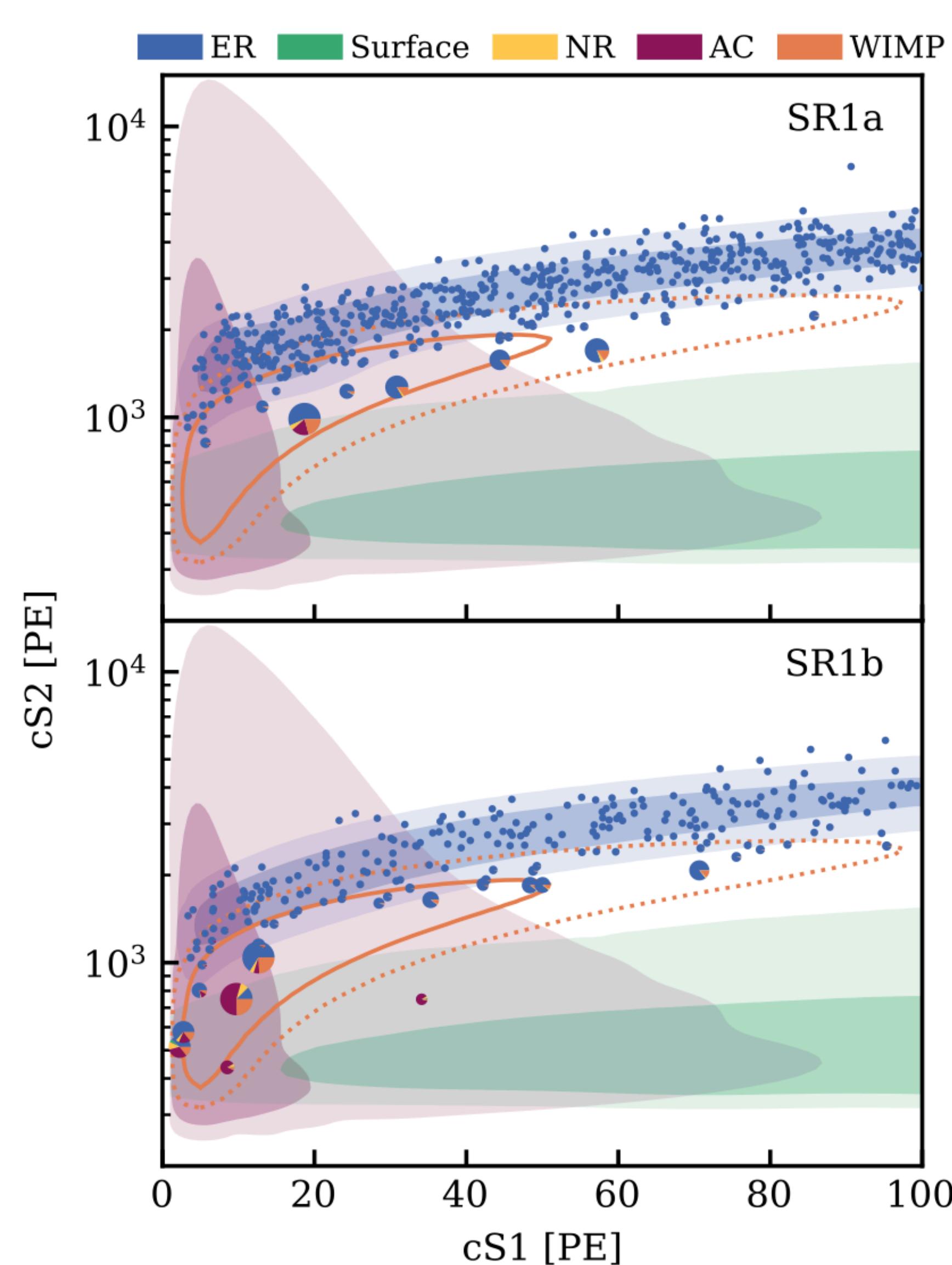
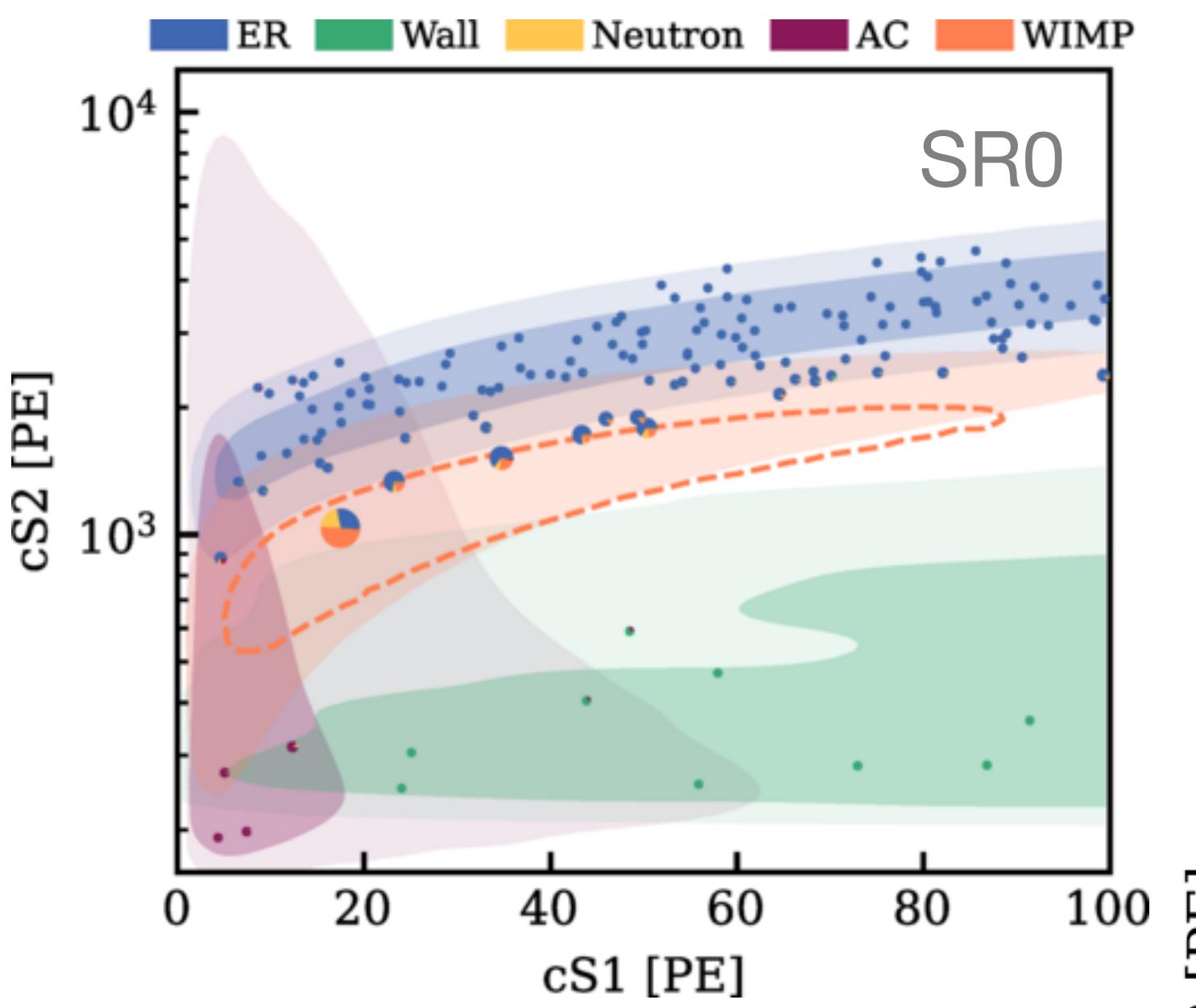
## Nominal WIMPs (Weakly Interacting Massive Particles) search

- No significant signal
- Main BG: ER leakage

Kr abundant period

	SR0		SR1a		SR1b	
	Nominal	Best fit	Nominal	Best fit	Nominal	Best fit
ER (flat)	134	136 $\pm$ 12	430 $\pm$ 30	450 $\pm$ 20	151 $\pm$ 11	154 $\pm$ 10
ER ( <sup>3</sup> H-like)	–	–	62	40 $\pm$ 30	101	80 <sup>+18</sup> <sub>-17</sub>
ER ( <sup>37</sup> Ar)	–	–	58 $\pm$ 6	55 $\pm$ 5	–	–
Neutron	0.7 $\pm$ 0.3	0.6 $\pm$ 0.3	0.47 $\pm$ 0.19	0.45 $\pm$ 0.19	0.7 $\pm$ 0.3	0.7 $\pm$ 0.3
CE $\nu$ NS (solar)	0.16 $\pm$ 0.05	0.16 $\pm$ 0.05	0.010 $\pm$ 0.003	0.010 $\pm$ 0.003	0.019 $\pm$ 0.006	0.019 $\pm$ 0.006
CE $\nu$ NS (atm.+DSNB)	0.04 $\pm$ 0.02	0.04 $\pm$ 0.02	0.024 $\pm$ 0.012	0.024 $\pm$ 0.012	0.05 $\pm$ 0.02	0.05 $\pm$ 0.02
AC	4.3 $\pm$ 0.9	4.4 <sup>+0.9</sup> <sub>-0.8</sub>	2.12 $\pm$ 0.18	2.10 $\pm$ 0.18	3.8 $\pm$ 0.3	3.8 $\pm$ 0.3
Surface	13 $\pm$ 3	11 $\pm$ 2	0.43 $\pm$ 0.05	0.42 $\pm$ 0.05	0.77 $\pm$ 0.09	0.76 $\pm$ 0.09
Total background	152	152 $\pm$ 12	553	550 $\pm$ 20	257	239 $\pm$ 15
WIMP (200 GeV/c <sup>2</sup> )	–	1.8	–	1.1	–	2.1
Observed	152		560		245	

# Backgrounds



**ER backgrounds**

**222Rn**  
**85Kr**  
**(<sup>3</sup>H)**

→ Distillation

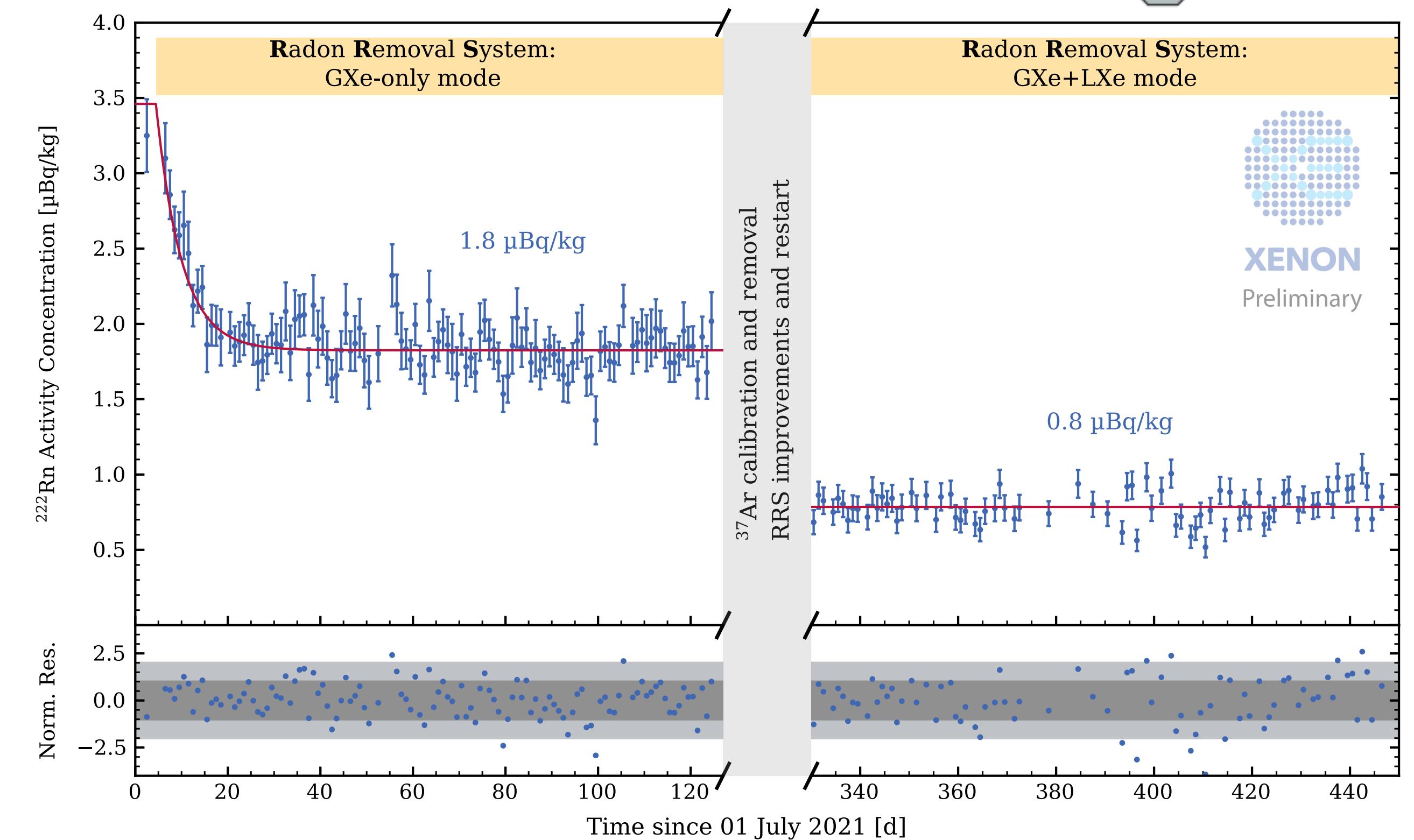
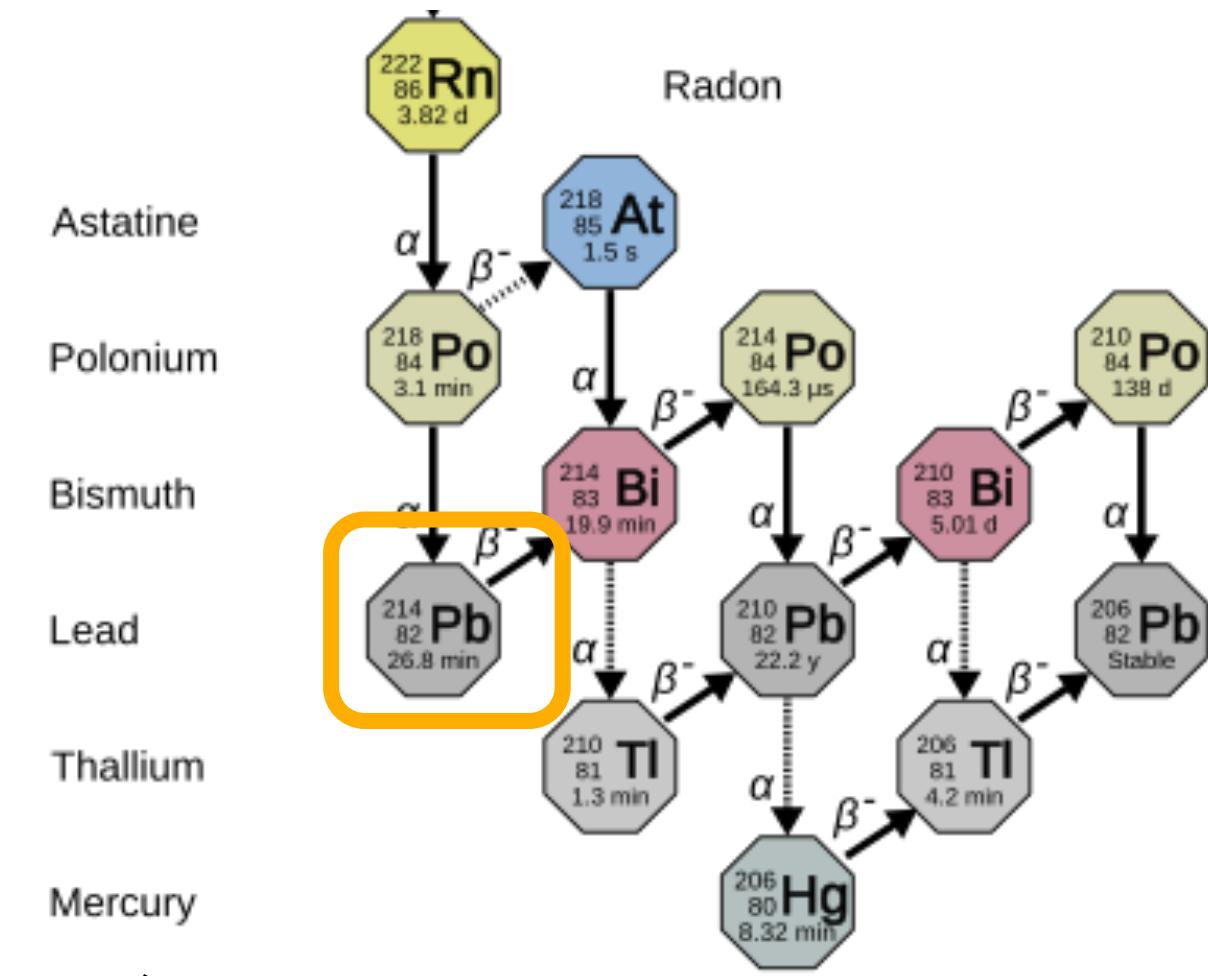
**NR background**

**neutrons**

→ Tag by  
Gd-Water Cherenkov

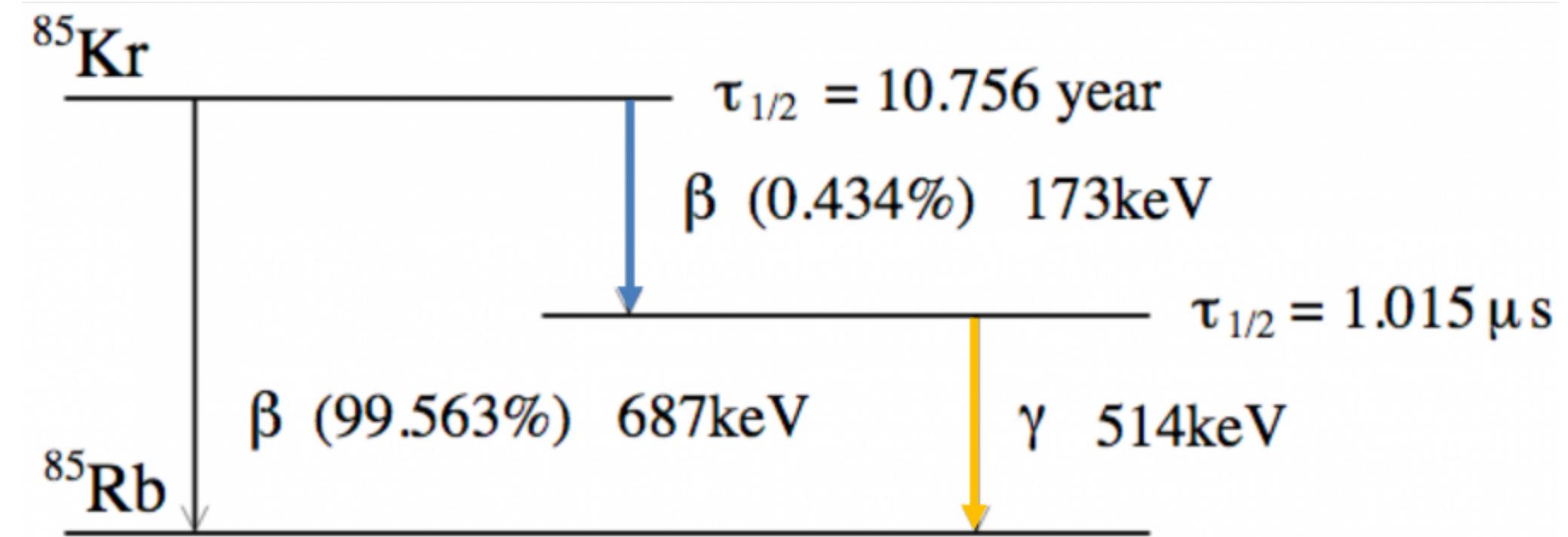
# 222Rn

- $\beta^-$ -decay of daughter  $^{214}\text{Pb}$  : flat ER background
- Emanation from detector materials  
→ continuous removal by online distillation



# 85Kr

- $\beta$ -decay : flat ER background
- Offline distillation → reduced to < ppt
- rate is constrained in two ways



## Rare gas mass spectroscopy

Measuring  $^{\text{nat}}\text{Kr}$  concentration in sampled Xe

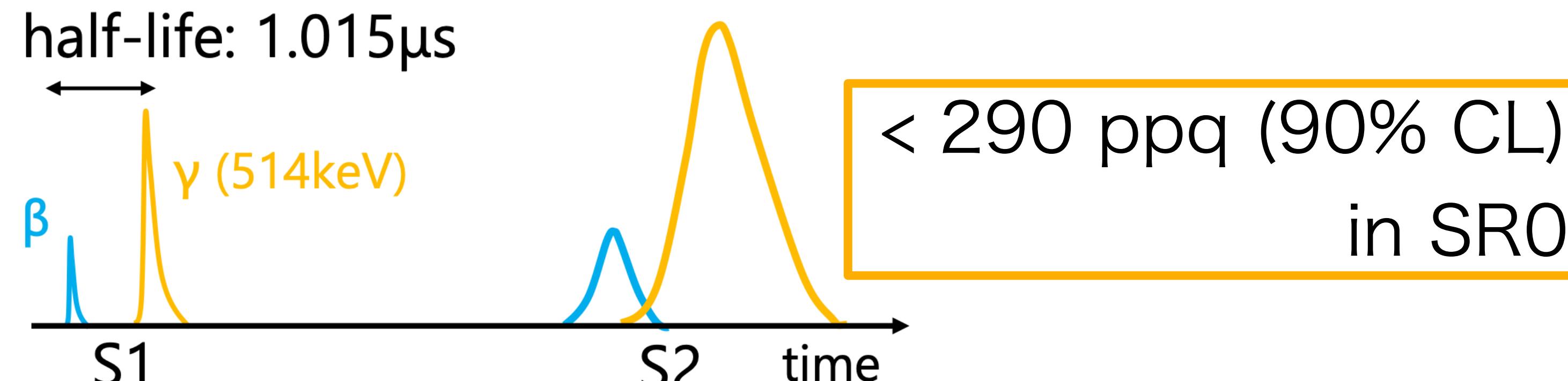
- $^{85}\text{Kr}/^{\text{nat}}\text{Kr} \sim 2 \times 10^{-11}$
- high sensitivity  
~ < 100 ppq level

**56±36 ppq in SR0**

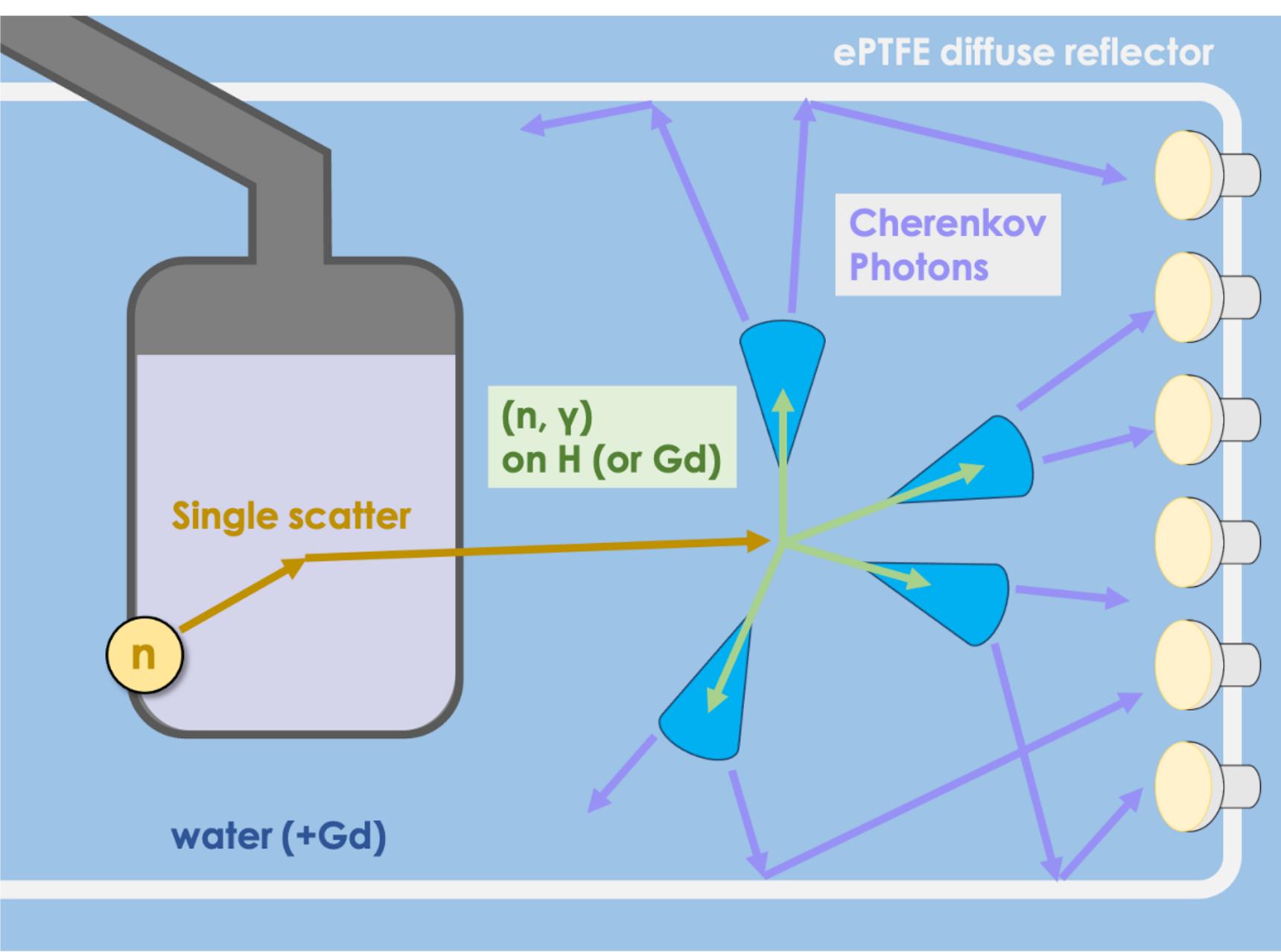
## Direct constraint by rare decay mode

Y. Kaminaga (D1, ICRR)

- Directly constrain  $^{85}\text{Kr}$  abundance by observing  $\beta + \gamma$  mode
- delayed coincidence search
  - independent of  $^{85}\text{Kr}/^{\text{nat}}\text{Kr}$  rate



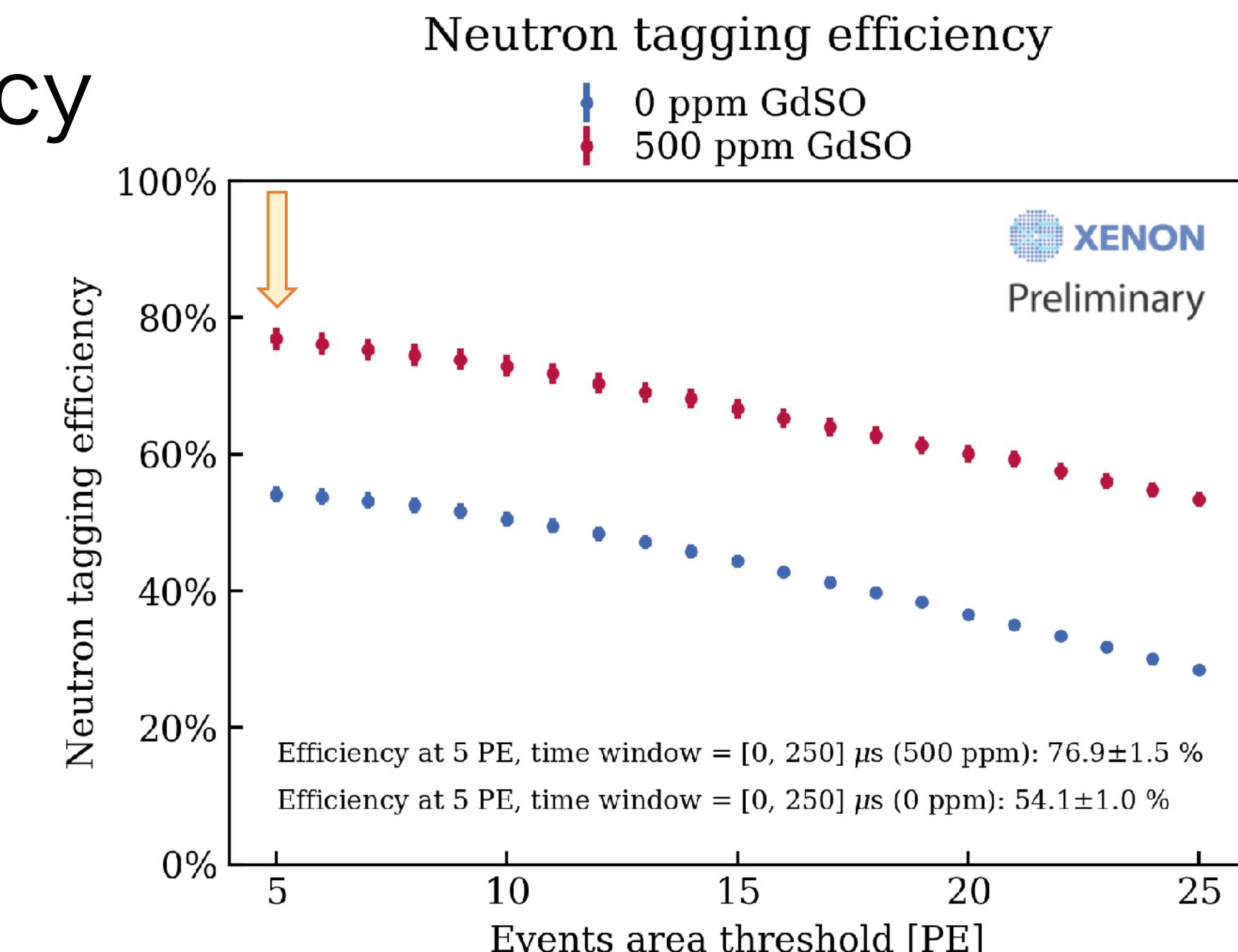
# Neutron veto detector



- Tagging neutrons after scattering in TPC
- Gd water Cherenkov detector
  - : techniques of EGADS, SK-Gd
- 120 Low-RI 8" PMTs (Hamamatsu R5912)
- Diffuse reflection by ePTFE panels
  - : > 99% reflectivity for ~ 300 nm

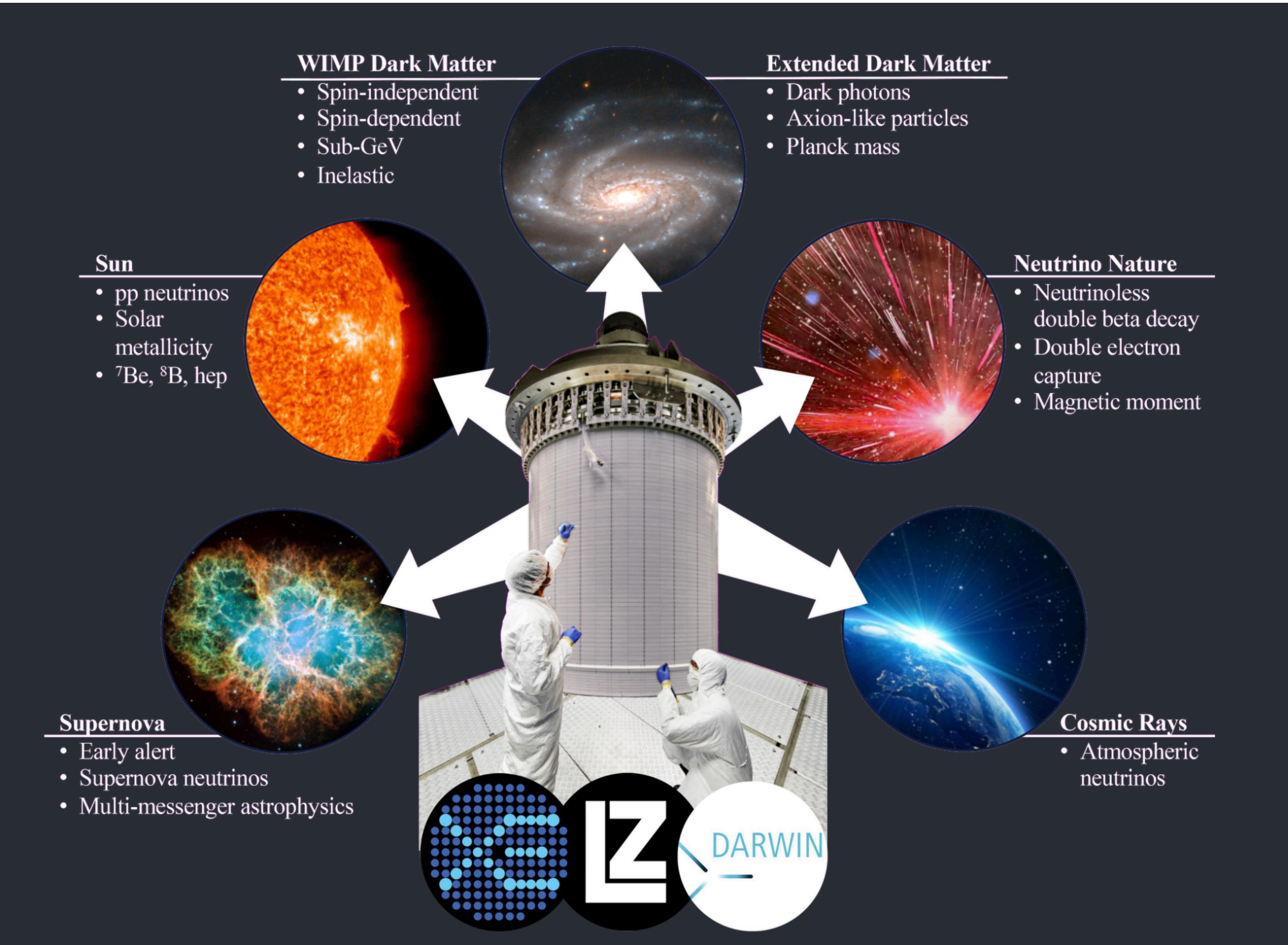
77% tagging efficiency  
@0.02% Gd

expecting 90%  
tagging efficiency  
@0.2% Gd



**XLZD**

## Joining the efforts from XENON/DARWIN and LZ

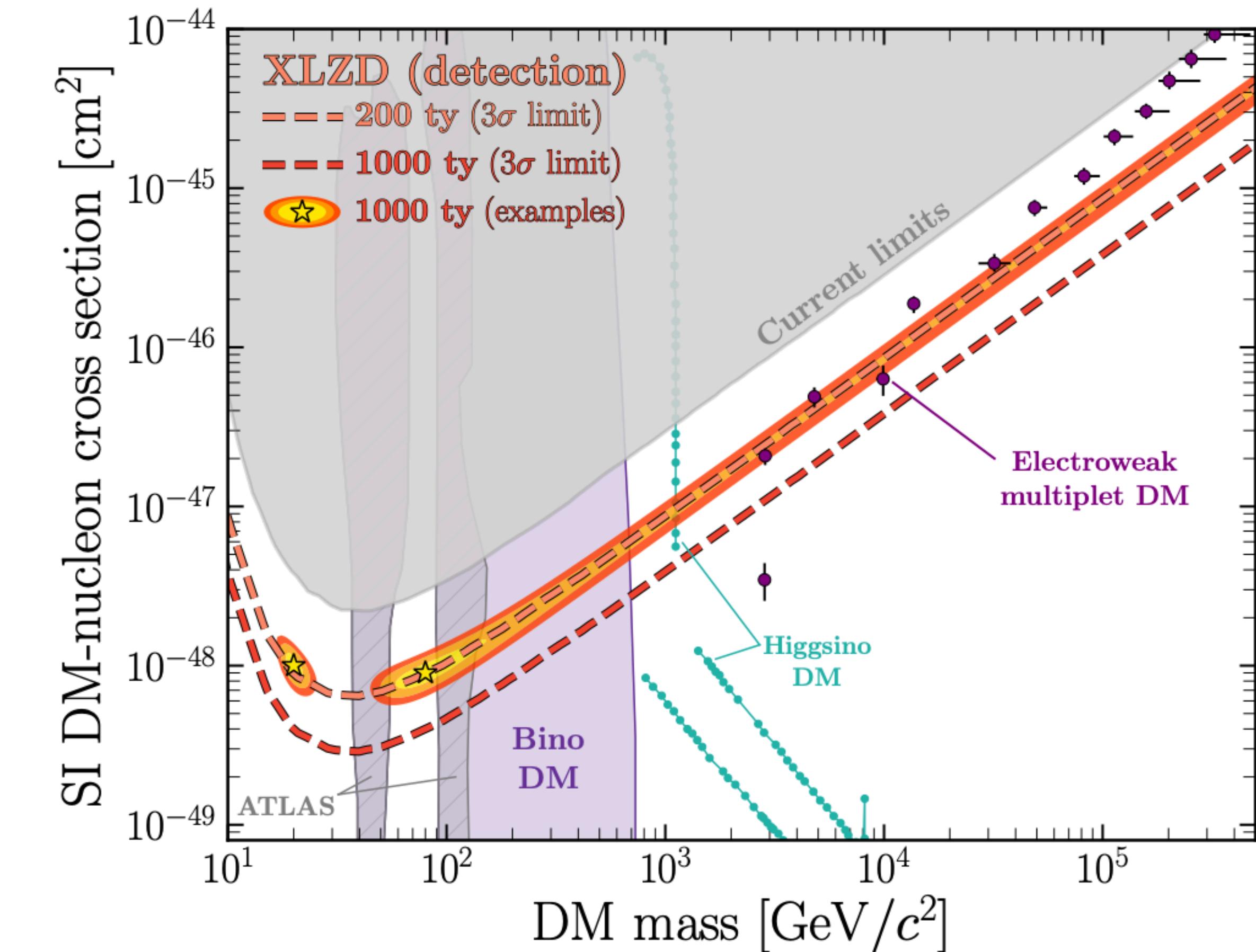
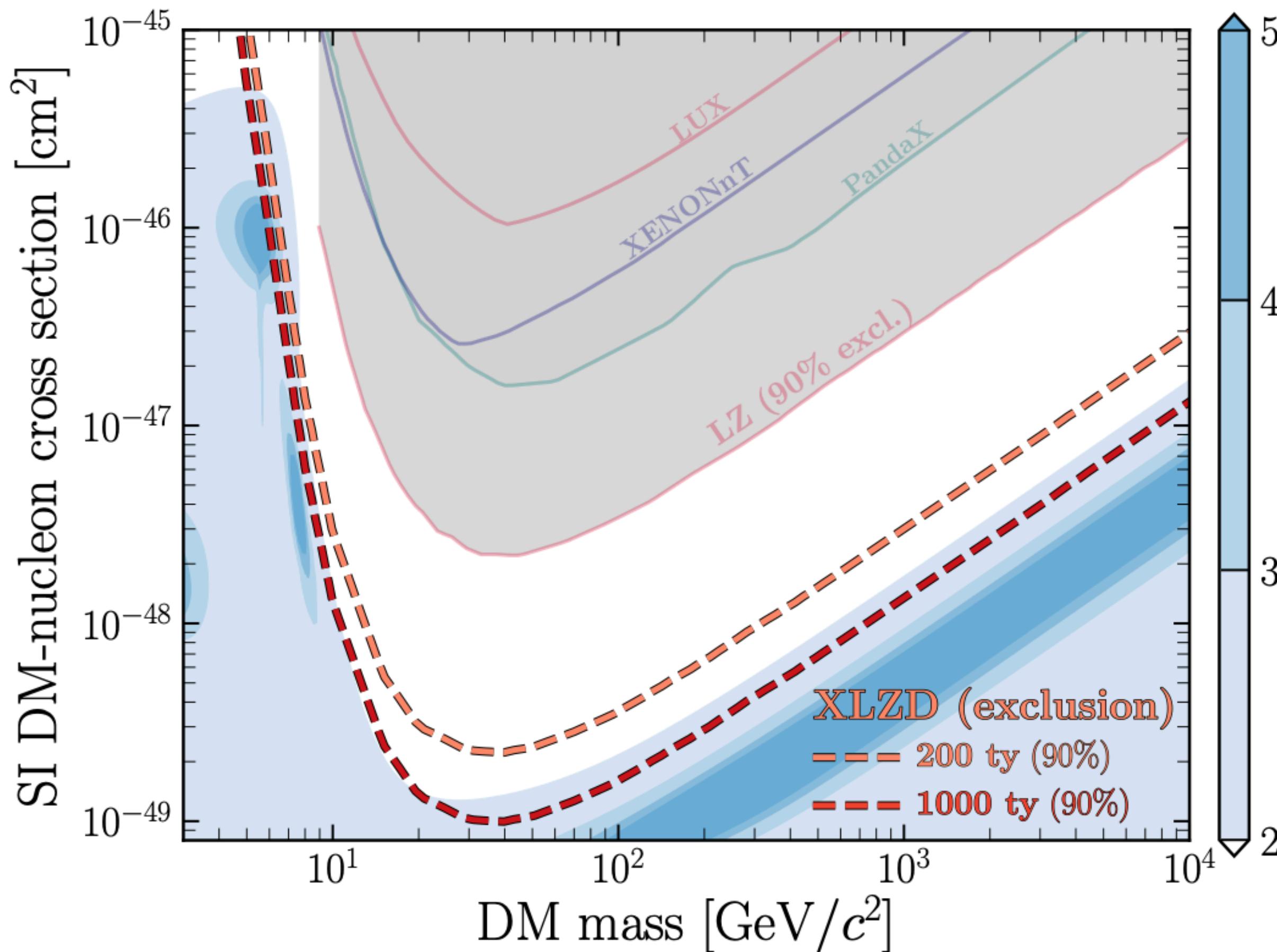


- Collaboration initiated in 2024
- step by step scaling  $40 \text{ t} \rightarrow 60 \text{ t} (\text{or } 80 \text{ t})$
- aiming  $1000 \text{ t}\cdot\text{y}$  exposure
- starting in first half of 2030s

- ✓ Searching WIMPs under “Neutrino fog”
- ✓ other various physics

# XLZD expected sensitivity

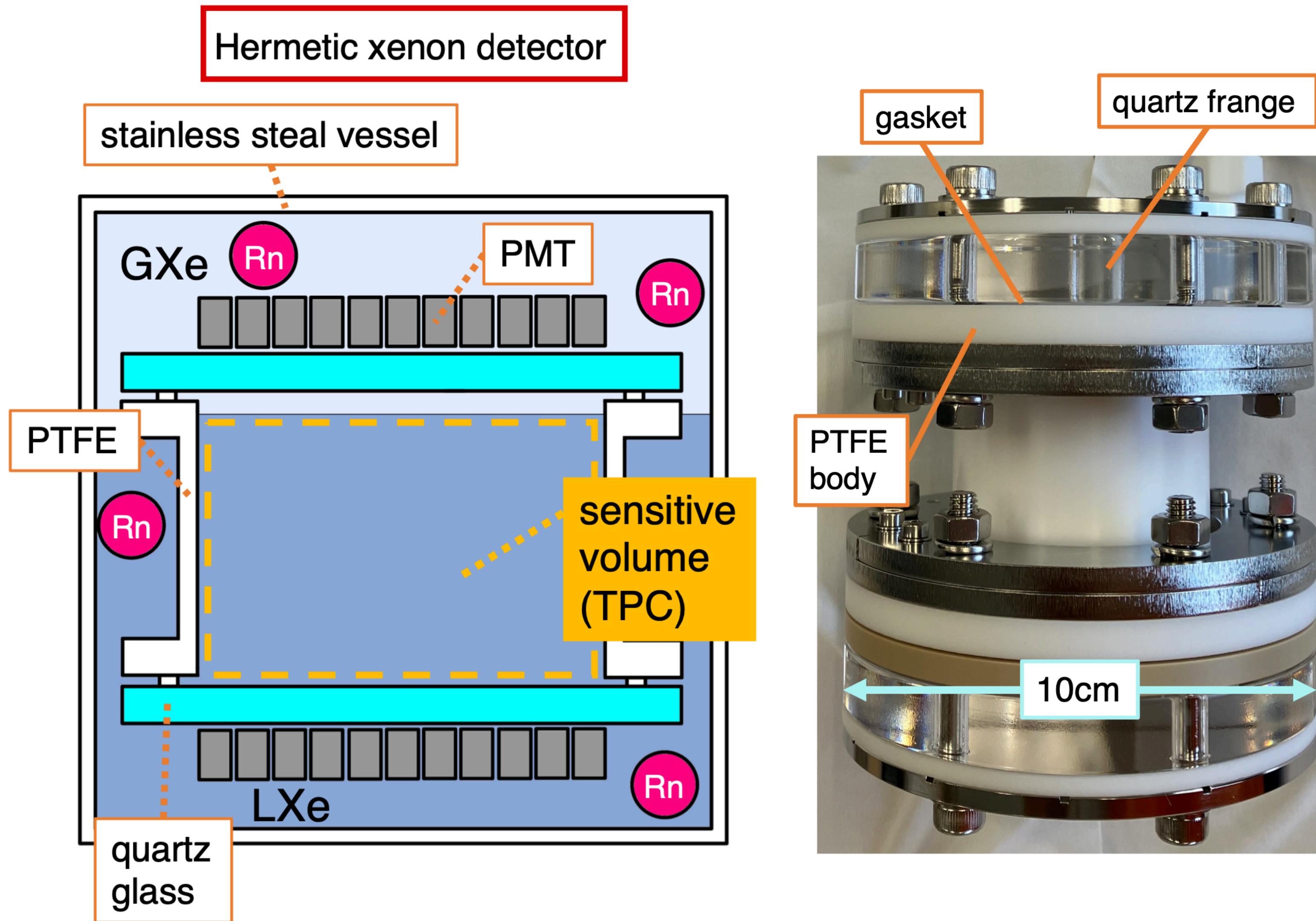
arXiv:2410.17137



To achieve these sensitivities,  
**Radioactive BGs must be negligible** compared with neutrino backgrounds!!

# Hermetic TPC

talk at Wakate workshop by R. Miyata (M2, Nagoya)  
poster by K. Fujikawa (M1, Nagoya)



## Target

reduce  $^{222}\text{Rn} < \sim 0.1 \mu\text{Bq/kg}$   
→ block emanated radon

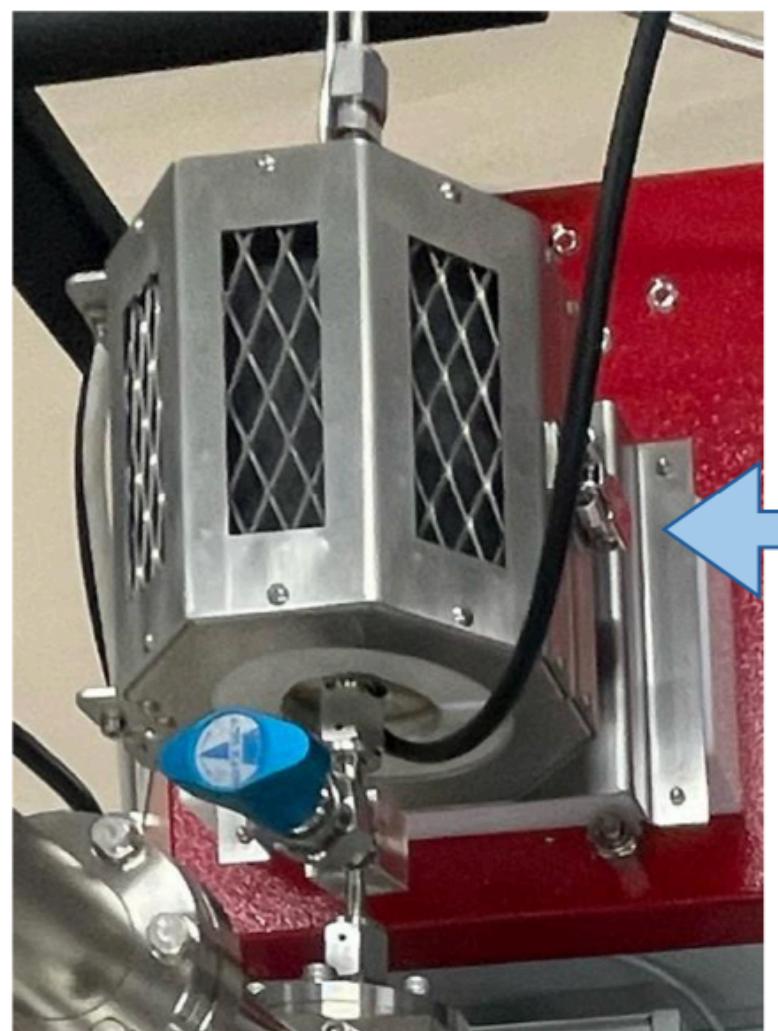
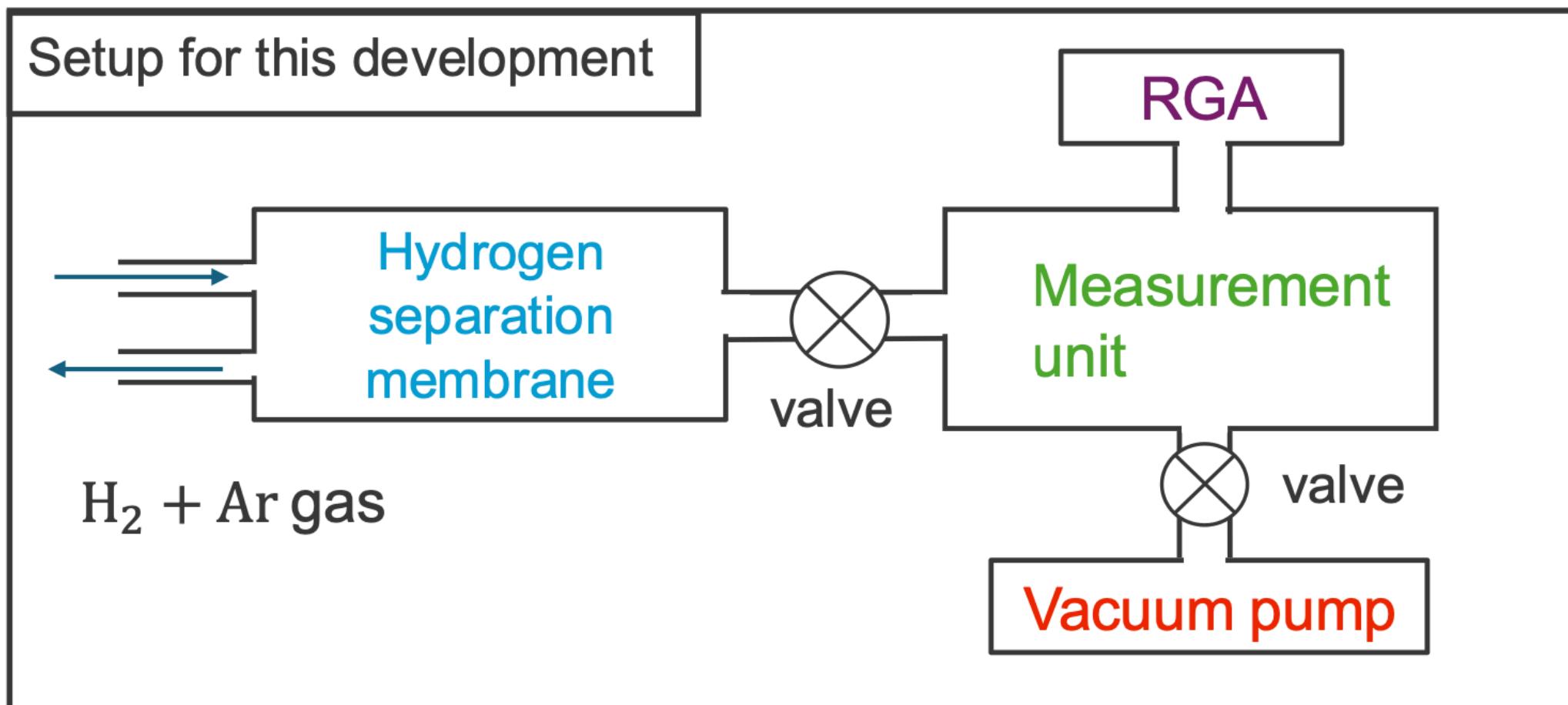
Sealing performance was  
tested with small setup

→ Radon reduction rate of  
 **$(9.1 \pm 1.5) \times 10^{-3}$**

scaled to XLZD size

→  **$^{222}\text{Rn} < 8.17 \times 10^{-3} \mu\text{Bq/kg}$**

# <sup>3</sup>H concentration measurement



Hydrogen separation  
membrane module



Hydrogen  
separation  
membrane

- <sup>3</sup>H: important BG component especially for Low-E ER new physics search
- Currently no method to constrain due to the low concentration

## Goal

Developing the method to measure  
**ppt level** hydrogen in xenon gas

→ hydrogen separation membrane + RGA

working principle is checked with  
O(10 ppb) H<sub>2</sub>/Ar mixture gas

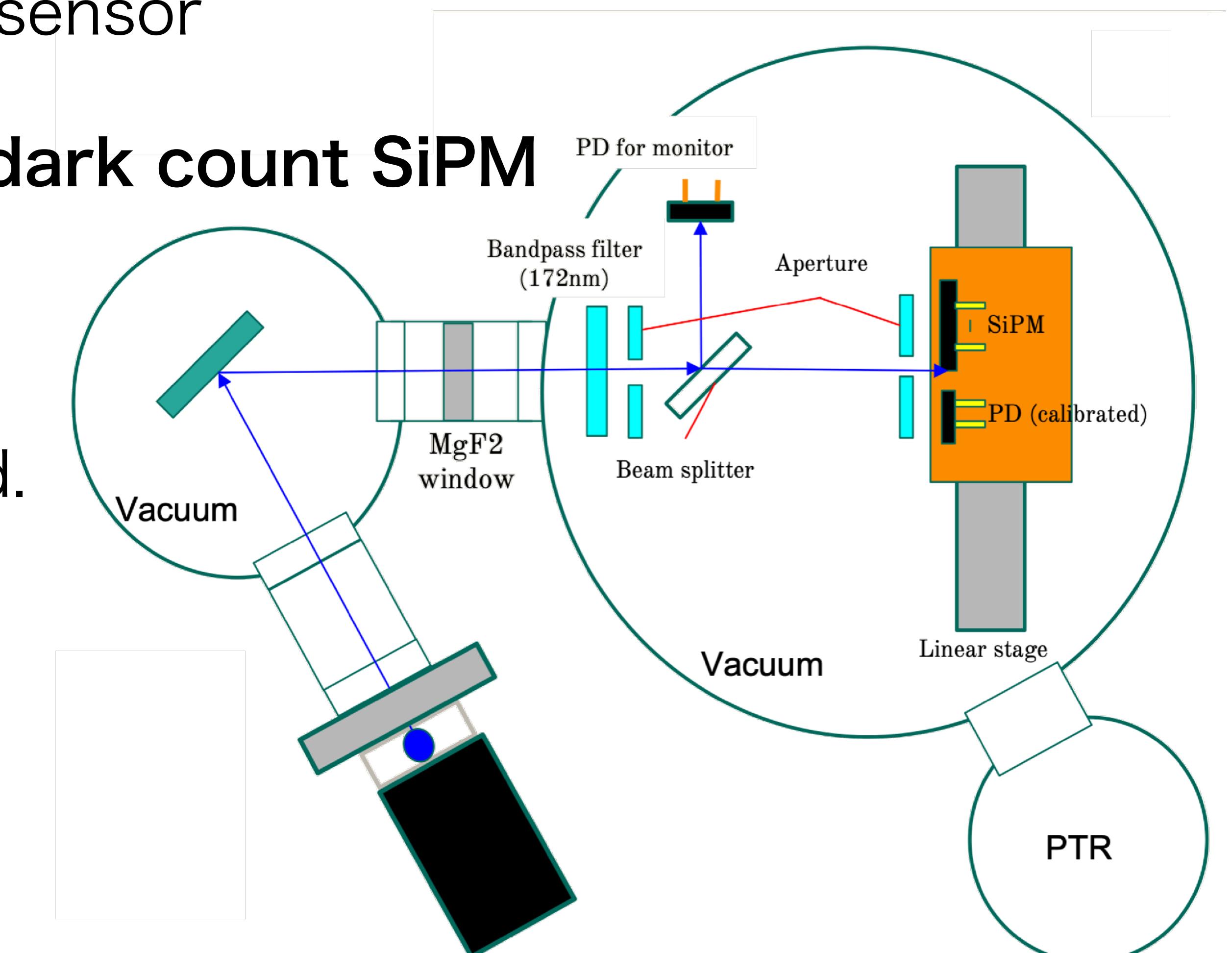
# Photosensor R&D

U/Th in TPC components should be lowered to reduce neutron BGs.

- Stem of PMTs: one of the dominant U/Th source  
→ Development of new Low-RI photosensor

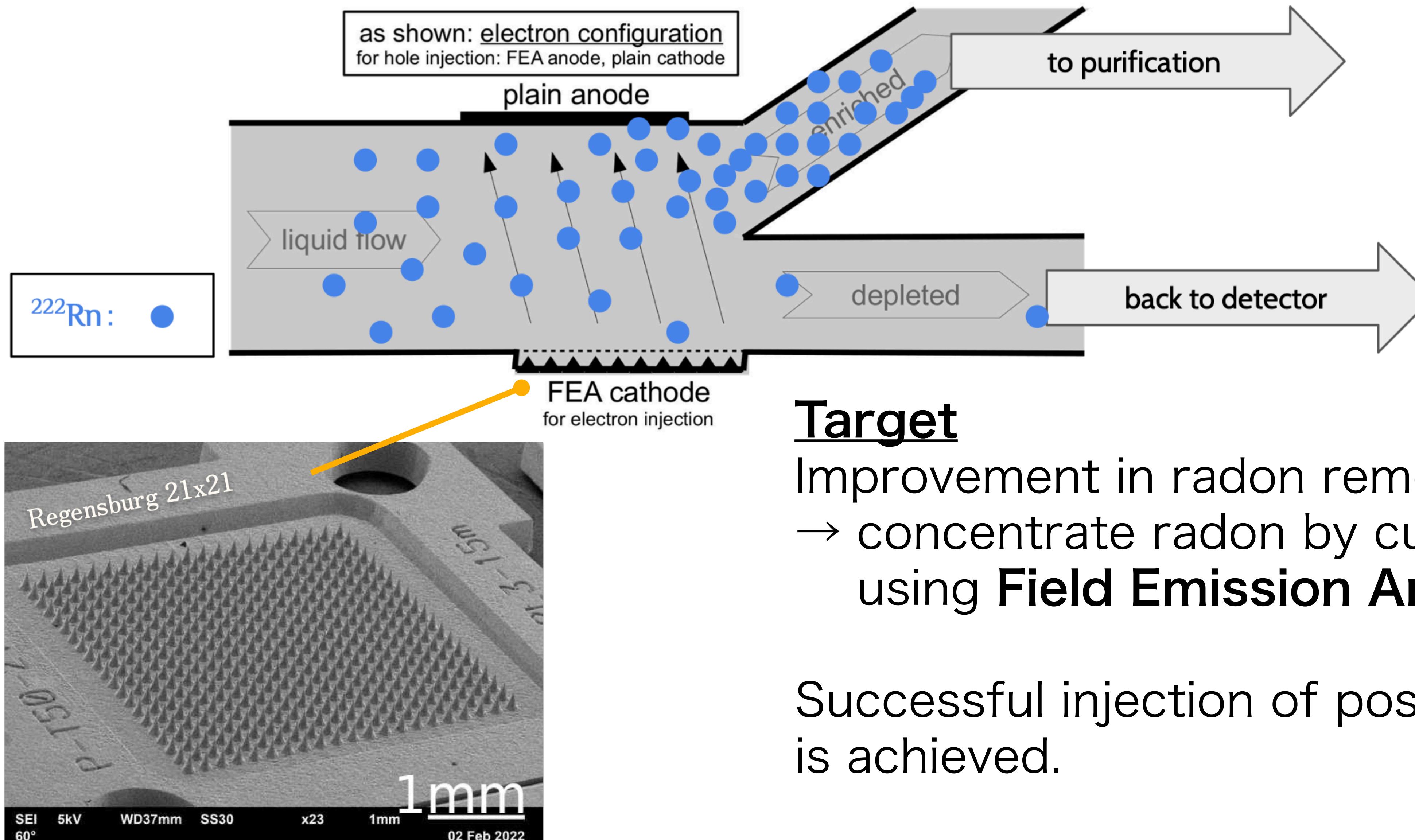
Possible solution: **VUV sensitive low dark count SiPM**

Dedicated PDE measurement system under liq. Xe temperature is developed.



# 222Rn removal with FEA

posters by Caio (M2, IPMU)  
and Xiaoxin (IPMU)



# Summary

## XENONnT

- Under operation with world competing sensitivity
- WIMPs search w/ SR0 + SR1 data  
 $\sigma_{\text{SI}} < 1.7 \times 10^{-47} \text{ cm}^2$  (90%CL) @30 GeV
- Background mitigation
  - ER: distillation of radioactive noble gases
  - NR: neutron tagging with Gd-Water Cherenkov detector

## XLZD

- Next generation dark matter project with 40--80 t LXe
- WIMPs search below the neutrino fog
- R&D for lowering radioactive BGs under neutrino level